# Size distributions of shocks and static avalanches from the functional renormalization group 

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#### Abstract

Interfaces pinned by quenched disorder are often used to model jerky self-organized critical motion. We study static avalanches, or shocks, defined here as jumps between distinct global minima upon changing an external field. We show how the full statistics of these jumps is encoded in the functional-renormalizationgroup fixed-point functions. This allows us to obtain the size distribution $P(S)$ of static avalanches in an expansion in the internal dimension $d$ of the interface. Near and above $d=4$ this yields the mean-field distribution $P(S) \sim S^{-3 / 2} e^{-S / 4 S_{m}}$, where $S_{m}$ is a large-scale cutoff, in some cases calculable. Resumming all one-loop contributions, we find $P(S) \sim S^{-\tau} \exp \left(C\left(S / S_{m}\right)^{1 / 2}-\frac{B}{4}\left(S / S_{m}\right)^{\delta}\right.$ ), where $B, C, \delta$, and $\tau$ are obtained to first order in $\epsilon=4-d$. Our result is consistent to $O(\epsilon)$ with the relation $\tau=\tau_{\zeta}:=2-\frac{2}{d+\zeta}$, where $\zeta$ is the static roughness exponent, often conjectured to hold at depinning. Our calculation applies to all static universality classes, including random-bond, random-field, and random-periodic disorders. Extended to long-range elastic systems, it yields a different size distribution for the case of contact-line elasticity, with an exponent compatible with $\tau=2-\frac{1}{d+\zeta}$ to $O(\epsilon=2-d)$. We discuss consequences for avalanches at depinning and for sandpile models, relations to Burgers turbulence and the possibility that the relation $\tau=\tau_{\zeta}$ be violated to higher loop order. Finally, we show that the avalanche-size distribution on a hyperplane of codimension one is in mean field (valid close to and above $d=4$ ) given by $P(S) \sim K_{1 / 3}(S) / S$, where $K$ is the Bessel- $K$ function, thus $\tau_{\text {hyper plane }}=\frac{4}{3}$.

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## I. INTRODUCTION

A hallmark of complex nonlinear systems, as well as systems with quenched disorder or inhomogeneities, is that the response to an applied field is very often not smooth but involves jumps, bursts, or avalanches. This is true for domain walls in a magnet responding to a change in external magnetic field, leading to the Barkhausen noise [1,2], the flux lattice in type-II superconductors upon varying the field [3], the contact line of a liquid partially wetting a disordered substrate when emptying the container $[4,5]$, or piles of granular material when adding grains [6]. Other examples are motion of cracks in brittle materials, dry friction [7], and earthquakes $[8,9]$. This jerky behavior often arises as a nonequilibrium phenomenon, but it may also occur at equilibrium, in systems with many metastable states, as they switch from one global minimum to another when an external perturbation is applied. The statistics of these jumps ubiquitously exhibits scale invariance and self-organized criticality as in sandpiles [10], with power-law tails for the probability of rare large events: if one defines an event size $S$, the probability distribution behaves as $P(S) \sim S^{-\tau}$, up to some largesize cutoff $S_{m}$, e.g., imposed by the finite system size.

An outstanding question is the degree of universality of the jump statistics. This can, in principle, be answered in particular prototype models, but even then it turns out to be rather difficult to obtain analytical results. One such class of models is sandpile automata which are dynamical systems where small events can trigger large avalanches [10-13]. There, some beautiful results have been obtained analytically, but the full avalanche statistics, including the distribution of avalanche sizes $P(S)$, has not yet been obtained in the cases of physical interest, such as spatial dimensions $d=2,3$. Mean-field (MF) theories [14,15] predict $\tau=3 / 2$
and various scaling arguments have been constructed [10-13, 16,17], not all in mutual agreement especially for $d$ $=2$. Among these, the conjecture $\tau=2-2 / d$ seems to be the best guess $[18,19]$ in $d=3$, but most often one has resorted to extensive numerics.

Another class of prototype models is random-field (RF) Ising magnets (RFIMs) in an external field $H$. There, one may study either the changes in the ground state as $H$ is varied or the nonequilibrium zero-temperature dynamical evolution from an initial state. Tuning the system near the critical values of field and disorder where the macroscopic magnetization jump vanishes, one can study finite-size avalanches, and these are found to exhibit self-organized criticality. Extensive numerical work has been carried out to determine avalanche-cluster statistics [20]. In parallel, the avalanche size distribution was studied [21] using a field theoretical renormalization group (RG) approach in an expansion in $d=6-\epsilon$. However, since the RG used there is based on the dimensional reduction property, which is well known to fail for even simpler disordered problems, as well as for the static RFIM, the status of the result in Ref. [21], $\tau=3 / 2+O\left(\epsilon^{2}\right)$ remains to be clarified. Furthermore it has been recently argued, mostly on the basis of numerics, that the avalanche statistics for the RFIM ground state and for its nonequilibrium dynamics are not distinct but belong to the same universality class [22]. Hence, another outstanding question is to clarify the possible differences between equilibrium and nonequilibrium avalanches in the RFIM and in a broader class of models.

Elastic systems pinned by a quenched random substrate provide yet another frequently used model for avalanche phenomena. In the statics they are known to exhibit glassy phases where the Gibbs measure, while localized near a minimal energy configuration, does exhibit jumps at low
temperature due to the presence of many low-lying metastable states. As an external force $f$ is applied, the system exhibits a depinning transition: at $T=0$ it starts moving at nonzero velocity only for $f>f_{c}$. The motion near the depinning threshold $f_{c}$ has been studied extensively numerically and is known to proceed by avalanches [23-26]. These are actually not obvious to define nonambiguously in the moving phase; hence some of the literature on the subject is qualitative. Understanding these avalanche statistics is however a challenging question since elastic plates driven by springs exhibit stick-slip motion, which provide a first step to model complex systems such as earthquakes In Ref. [8] a very simple model for avalanches, mean field in inspiration, was proposed and solved and yields again the value $\tau=3 / 2$. Surprisingly, there is to this day no first-principles derivation of even this simple mean-field behavior within this class of models. Besides extensive numerics, the main theoretical result is a conjecture for the size exponent near depinning, equivalent to $\tau=\tau_{\zeta}=2-2 /(d+\zeta)$, based on a scaling argument and some unproven assumptions [27,28]. Thus there is the need for an analytic tool to approach this problem. The functional renormalization group (FRG), a powerful fieldtheoretical method to deal with disordered elastic systems in the statics [29-44] and in the driven dynamics [27,45-53], as well as with random field models [54-61], has been quite successful at computing, for instance, the roughness properties of the displacement field $u(x)$ in a dimensional $d=4-\epsilon$ expansion. Following the procedure proposed in [62-64], the numerical determination of the predicted FRG fixed-point functions in dimensions $d=1,2,3$ has confirmed the theory to high precision, up to two-loop accuracy both in the statics [65] and at quasistatic depinnings [66]. This was an important test, given that the field theory is quite nonstandard: the coupling constant is a function of the field, $\Delta(u)=-R^{\prime \prime}(u)$, the renormalized second cumulant of the pinning force, and the zero-temperature effective action is nonanalytic, e.g., $\Delta(u)$ exhibits a cusp at $u=0$. However, until now no calculation of avalanche distributions has been even attempted using FRG. In fact, possible difficulties in handling the fast-jump motion within the field theory have been emphasized [46], and it was unclear whether it was possible at all.

The aim of this paper is to show that one can extract the avalanche statistics from the FRG in a controlled way. It is $a$ priori quite involved as it requires the calculation of the nonanalytic part (the cusp) of all cumulants of the renormalized disorder, which encode the jump distribution. However, it turns out to be feasible in the end, as we find that remarkably simple self-consistent equations are obeyed by suitable generating functions. As an application we compute the distribution of avalanche sizes from first principles in an expansion in $d=4-\epsilon$. We derive the mean-field result and obtain a one loop, i.e., $O(\epsilon)$ prediction for $P(S)$. Here we study the case of static avalanches, also called shocks, which are defined as jumps between distinct global minima of the energy upon changing an external field. The most convenient setting is to add an external harmonic well with a variable central position, i.e., consider an elastic interface in a random potential tied to a harmonic spring. As a function of the spring position, the center of mass of the interface changes in discrete jumps. We study all static universality classes including
random-bond (RB), random-field, and random-periodic (RP) disorders, as well as long-range (LR) elastic systems. We obtain not only the exponent $\tau$ but the full scaling function of the avalanche distribution. The latter is (almost fully) universal with respect to small scales, and the dependence in the large scales, i.e., in $S / S_{m}$, is calculable. A short summary of some of our main results, together with a parallel numerical study, has recently appeared [26] and here we provide the necessary details of the approach and of the calculation. Although we do not directly study depinning here, there are connections between static and dynamical FRGs, the main idea of the method being similar. As found in a companion study [67] of avalanches at depinning, the results of the present paper are most likely to hold also for depinning to one loop. Since for some observables, such as the roughness exponent $\zeta$, differences between statics and depinning within the FRG appear only at two loops [32], the question of the difference between nonequilibrium and static avalanches remains open and requires a two-loop study [67]. Note that to our knowledge, while avalanches near depinning have received a lot of attention, the static avalanches and shocks defined here have not. Their connection to the Burgers equation [62,63], which provides their $d=0$ limit, helped formulate the problem in a way amenable to FRG calculations.

Let us also point out that there are interesting relations between the depinning in the random periodic class, i.e., charge density waves (CDWs), and sandpile automata [24,68]. Sandpile automata are also related to spanning trees and loop erased random walks $[18,19,69-74]$, and this has lead to recent FRG predictions for the latter [75]. Hence the present work is also relevant to sandpile models. An important issue is the status of the conjecture $\tau=\tau_{\zeta}$, which in sandpile models reads $\tau=\tau_{\zeta=0}$ consistent with a value $\zeta=0$ for CDW depinning. Of course, in the present paper the question becomes whether such a relation also exists in the statics. Finally, let us note that our results extend beyond disordered systems. Indeed in some complex dynamical systems, the randomness lies in the initial condition, and avalanches or shocks appear in the course of a nonlinear deterministic evolution. This is the case for the sandpile automata (e.g., in the conserved energy ensemble) as well as for the problem of decaying Burgers turbulence. The latter is indeed intimately connected to the pinning problem, and to the FRG, as is recalled below. Therefore the methods discussed in this paper should find applications in a much broader class of nonlinear complex systems, possibly including turbulence and spin glasses.

The outline of this paper is as follows. In Sec. II we define the model of an elastic manifold in a random potential and a harmonic well. In Sec. III we give the expected scaling form for shock distributions. In Sec. IV we detail the connection to FRG functions. In Sec. V we present the tree-level calculation which leads to the mean-field result. In Secs. VI and VII we present the one-loop resummation and in Sec. VIII we discuss the result for $P(S)$. The local avalanche-size distribution is discussed in Sec. IX. Extensions to nonlocal elasticity are given in Sec. X. More technical material is relegated to various appendixes.

Note that we will not review manifolds in random potentials, for which we refer to [76,77], nor static FRG methods,
for which we refer to [78] for a pedagogical introduction and to $[36,63]$ for details of the FRG calculations.

## II. MODEL

We consider an elastic interface parametrized by a one component $(N=1)$ displacement field noted $u(x) \equiv u_{x}$. It is subjected to a random potential $V(u, x)$ and to a harmonic well centered at $u_{x}=w_{x}$. It is defined by the Hamiltonian

$$
\begin{equation*}
\mathcal{H}[u ; w]=\frac{1}{2} \int_{q} g_{q}^{-1}\left|u_{q}-w_{q}\right|^{2}+\int_{x} V\left(u_{x}, x\right), \tag{1}
\end{equation*}
$$

where we denote $\int_{x_{d}}=\int d^{d} x$ and the Fourier transform $u_{q}$ $=\int_{x} e^{i q x} u_{x}$ and $\int_{q}=\int \frac{d^{d} q}{(2 \pi)^{d}}$. A cutoff at small scale $x \sim a$ or $q$ $\sim \Lambda$ is implicit everywhere. In most of the paper we focus on the choice of a standard local elasticity,

$$
\begin{equation*}
g_{q}^{-1}=K q^{2}+m^{2} \tag{2}
\end{equation*}
$$

and often set $K=1$ for convenience. The bare disorder is assumed to be short range (SR) in internal space, statistically translationally invariant, with a bare second cumulant,

$$
\begin{equation*}
{\overline{V(u, x) V\left(u^{\prime}, 0\right)}}^{c}=\delta\left(x-x^{\prime}\right) R_{0}\left(u-u^{\prime}\right) \tag{3}
\end{equation*}
$$

whose precise form is unimportant, apart from some global features which determine the universality classes, mainly (i) RB with $R_{0}(u)$ as a SR function, (ii) RF with $R_{0}(u) \sim-\sigma|u|$ as a LR function, and (iii) RP with $R_{0}(u)$ as a periodic function of period (arbitrarily) set to unity.

To study the equilibrium statics we define the renormalized potential $\hat{V}$ as the free energy of the system and the renormalized force as its functional derivative with respect to $w_{x}$,

$$
\begin{gather*}
e^{-\hat{V}[w] / T}=\int D u_{x} e^{-\mathcal{H}[u ; w] / T},  \tag{4}\\
\hat{V}_{x}^{\prime}[w]=\int_{x^{\prime}} g_{x x^{\prime}}^{-1}\left(w_{x^{\prime}}-\left\langle u_{x^{\prime}}\right\rangle\right),  \tag{5}\\
g_{x x^{\prime}}=\int_{k} g_{k} e^{i k\left(x-x^{\prime}\right)}, \tag{6}
\end{gather*}
$$

where $\langle\cdots\rangle$ denotes thermal averages over $\mathcal{H}[u ; w]$ in a given disorder realization and $[\cdots]$ is reserved for arguments of functionals, while $(\cdots)$ is used for arguments of functions. For further details and notations we refer to [62-66] where this model and the observables where discussed and studied.

In this paper we are mostly interested in energy minimization as $w$ is varied in a given realization of the random potential $V(x, u)$, i.e., the $T=0$ problem. At $T=0$ the minimum-energy configuration is denoted $u_{x}[w]$, and $u_{x}(w)$ for $w$ is uniform, thus dropping the expectation value. $\hat{V}[w]$ becomes the minimum energy,

$$
\begin{equation*}
\hat{V}[w]=\min _{u_{x}} \mathcal{H}[u ; w] \tag{7}
\end{equation*}
$$

and for a uniform $w_{x}=w$ one defines the ground-state energy of the system per unit volume,

$$
\begin{equation*}
\hat{V}(w):=L^{-d} \hat{V}\left[\left\{w_{x}=w\right\}\right] \tag{8}
\end{equation*}
$$

$L^{d}$ being the volume of the system. Its derivative with respect to $w$,

$$
\begin{align*}
\hat{V}^{\prime}(w) & =m^{2}(w-u(w))  \tag{9}\\
u(w) & =L^{-d} \int_{x} u_{x}(w) \tag{10}
\end{align*}
$$

coincides with the force per unit volume exerted by the spring. We are ultimately interested in the small-mass limit $m \rightarrow 0$ where scale invariance becomes manifest. As shown previously $[30,65,76,77,79,80]$, the optimal interface is statistically self-affine with $\overline{(u(x)-u(0))^{2}} \sim|x|^{2 \zeta}$ and a roughness exponent $\zeta$ which depends on the class of disorder and with a $\epsilon=4-d$ expansion [32]: $\zeta=\epsilon / 3$ for RF, $\zeta=0$ for RP, and $\zeta=0.2083 \epsilon+0.00686 \epsilon^{2}$ for $\mathrm{RB}(\zeta=2 / 3$ in $d=1)$. This holds for scales $L_{c}<L<L_{m}$, where $L_{c}$ is the Larkin length (here of the order of the microscopic cutoff) and $L_{m} \sim 1 / m$, the large scale cutoff induced by the harmonic well. It is useful to picture the interface as a collection of $\left(L / L_{m}\right)^{d}$ regions pinned almost independently.

Note that in this paper we study the static problem where the interface finds the global energy minimum for each $w$. The function $u(w)$ is then a single valued monotonically increasing function of $w$. One can also study the quasistatic dynamic problem where $w(t)$ grows very slowly as a function of time, and the interface visits a deterministic sequence of metastable states [64]. In that case $u(w)$ is history dependent, although due to the no-passing theorem [81] different initial conditions converge to the same asymptotic trajectory as $w$ is increased by an amount larger than $L_{m}^{\zeta}$ [66]. In the $m \rightarrow 0$ limit this is a method to study avalanches near the depinning transition. It is different from the more standard method where the force is increased infinitesimally while remaining below threshold $f<f_{c}$, and in a sense it is a cleaner method since it produces a steady state for avalanches without getting closer to threshold, i.e., without changing the cutoff length which remains $L_{m} \sim 1 / \mathrm{m}$. The study of this case requires dynamical FRG and will be performed in [67].

## III. OBSERVABLES

We now define a few useful observables and notations, focusing on the simplest case of a parabola centered around a uniform $w_{x}=w$. We also discuss their expected scaling form in the limit $m \rightarrow 0$, confirmed later via the FRG analysis.

## A. Energy fluctuations

The renormalized potential defined in Eq. (4) is a random function of $w$ and of the bare random potential $V(u, x)$. Hence one can define its cumulants as averages over the bare random potential,

$$
{\overline{\hat{V}}\left(w_{1}\right) \hat{V}\left(w_{2}\right)}^{c}=L^{-d} R\left(w_{1}-w_{2}\right)
$$

$$
\begin{equation*}
{\overline{\hat{V}\left(w_{1}\right) \cdots \hat{V}\left(w_{n}\right)}}^{c}=L^{-(n-1) d}(-1)^{n} \hat{S}^{(n)}\left(w_{1}, \ldots, w_{n}\right) . \tag{11}
\end{equation*}
$$

We denote $\bar{S}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$ the same expectation values for nonconnected averages and use that $\hat{R}=R$. Note that connected averages scale with system size as indicated, while nonconnected averages do not. In the limit $m \rightarrow 0$ one expects that they take the scaling form $\hat{S}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$ $\sim m^{d-n \theta \hat{s}^{(n)}}\left(m^{\zeta} w_{1}, \ldots, m^{\zeta} w_{n}\right)$ with $\theta=d-2+2 \zeta$.

## B. Force fluctuations

One defines the cumulants of the renormalized pinning force,

$$
\begin{gather*}
m^{4}{\overline{\left(u\left(w_{1}\right)-w_{1}\right)\left(u\left(w_{2}\right)-w_{2}\right)}}^{c}=L^{-d} \Delta\left(w_{1}-w_{2}\right)  \tag{12}\\
m^{2 n}{\overline{\left(u\left(w_{1}\right)-w_{1}\right) \cdots\left(u\left(w_{n}\right)-w_{n}\right)}}^{c} \\
=(-1)^{n}{\overline{\hat{V}^{\prime}\left(w_{1}\right), \ldots, \hat{V}^{\prime}\left(w_{n}\right)}}^{c} \\
\quad=L^{-(n-1) d}(-1)^{n} \hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right) \tag{13}
\end{gather*}
$$

and similarly with $\bar{C}^{(n)}$ for nonconnected averages. (The same remark as above applies for the system size dependence.) In the first line of Eq. (12) the connected average can be replaced by a nonconnected average since $\overline{u\left(w_{1}\right)-w_{1}}=0$ (as shown by parity and statistical translational invariance of the disorder). One has the relation $\hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$ $=(-1)^{n} \partial_{w_{1}} \ldots \partial_{w_{n}} \hat{S}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$ and $\Delta(w)=-R^{\prime \prime}(w)$. In the limit $m \rightarrow 0$ one expects that they take the scaling form $\hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right) \sim m^{2 n-(n-1) d-n \zeta} \hat{c}^{(n)}\left(m^{\zeta} w_{1}, \ldots, m^{\zeta} w_{n}\right)$, and

$$
\begin{equation*}
\Delta(w)=A_{d} m^{\epsilon-2 \zeta} \widetilde{\Delta}\left(m^{\zeta} w\right) \tag{14}
\end{equation*}
$$

where a convenient choice of the constant $A_{d}$ is given below [see Eq. (B3)].

## C. Shock observables

Consider a uniform $w$. It is reasonable to assume (and confirmed by numerical studies $[26,65])$ that $u_{x}(w)$ consists of smooth parts, which become constant in the scaling limit $u_{x}(w) \sim m^{-\xi}$ as $m \rightarrow 0$, and jumps, also called shocks, or static avalanches, that it can be decomposed as

$$
\begin{gather*}
u_{x}(w)=\sum_{i} S_{i}^{x} \theta\left(w-w_{i}\right)  \tag{15}\\
u(w)=\frac{1}{L^{d}} \sum_{i} S_{i} \theta\left(w-w_{i}\right),  \tag{16}\\
S_{i}=\int_{x} S_{i}^{x} \tag{17}
\end{gather*}
$$

Here $S_{i}$ is the "size" of the shock labeled $i$ and $\theta(x)$ is the unit-step function. Each environment corresponds a unique set of $u_{x}(w)$ and a unique set $\left\{\left(w_{i}, S_{i}\right)\right\}$. From this one defines the normalized probability of an (infinite) sequence
$p\left(\cdots ; w_{1}, S_{1} ; \cdots w_{n}, S_{n} ; \cdots\right)$. One can also define the onepoint probability density,

$$
\begin{equation*}
\rho(S, w)=\overline{\sum_{i} \delta\left(S-S_{i}\right) \delta\left(w-w_{i}\right)}, \tag{18}
\end{equation*}
$$

so that the total number of shocks in an interval of size $w$ is $N_{w}=\int d S \int_{0}^{w} \rho(S, w)$. We will assume that $N_{w}$ is proportional to $w$. This is equivalent to

$$
\begin{equation*}
\rho(S, w)=\rho_{0} P(S)=\rho(S), \quad \rho_{0}=\int_{S_{\min }}^{\infty} d S \rho(S), \tag{19}
\end{equation*}
$$

where $P(S) d S$ is the normalized probability that a given shock has a size in the interval $[S, S+d S]$ and $\rho_{0} d w$ is the average total number of shocks in an interval $d w$, i.e., the shock density, assumed here to be a finite number. Note that the mass $m$ provides a scale ensuring convergence at large $S$, while the notation $S_{\text {min }}$ refers to the avalanche size at the small scale cutoff of the model, which in some cases will be explicitly needed to ensure convergence at small $S$ of our scaling forms (see below). The statistical translational invariance of the disorder together with parity implies

$$
\begin{equation*}
\overline{u^{\prime}(w)} \equiv \frac{1}{L^{d}} \overline{\sum_{i} S_{i} \delta\left(w-w_{i}\right)} \equiv \frac{1}{L^{d}} \int_{S_{\min }}^{\infty} d S S \rho(S)=1 \tag{20}
\end{equation*}
$$

This also gives

$$
\begin{equation*}
L^{-d} \rho_{0}\langle S\rangle=1 \tag{21}
\end{equation*}
$$

where here and below $\left\langle S^{n}\right\rangle:=\int d S S^{n} P(S)$ are normalized moments.

We now consider the $m \rightarrow 0$ limit. One expects and later verifies that the shock distribution takes the following scaling form (see [24,26]):

$$
\begin{equation*}
\rho(S)=L^{d} m^{\rho} S^{-\tau} \widetilde{\rho}\left(S m^{d+\zeta}\right) \tag{22}
\end{equation*}
$$

where $\widetilde{\rho}(s=0)$ is a constant and $\widetilde{\rho}(s)$ has a fast decay to zero at large $s$, suppressing large avalanches. This form involves $a$ priori two exponents $\rho$ and $\tau$. However if one assumes $\tau$ $<2$, which is found to always hold here, constraint (20) implies a relation between the two exponents obtained by writing, as $m \rightarrow 0$,

$$
\begin{gather*}
\frac{1}{L^{d}} \int_{S_{\min }}^{\infty} d S S \rho(S)=m^{\rho-(2-\tau)(d+\zeta)} \int_{0}^{\infty} d s s s^{-\tau} \widetilde{\rho}(s)=1 \\
\rho=(2-\tau)(d+\zeta), \quad \int_{0}^{\infty} d s s^{1-\tau} \widetilde{\rho}(s)=1 \tag{23}
\end{gather*}
$$

It also implies a constraint on the scaling function. From now on we denote $S_{m}$ the scale at which the avalanche sizes are cut by the mass $m$. It behaves at small mass as

$$
\begin{equation*}
S_{m}=c m^{-(d+\zeta)} \tag{24}
\end{equation*}
$$

For now it is defined up to a multiplicative constant $c$, for which we make a convenient choice below.

We must now distinguish two cases where the distribution is qualitatively different:
(i) $\tau<1$. In that case the total shock density

$$
\begin{equation*}
\rho_{0}=L^{d} m^{(d+\zeta)} \int_{0}^{\infty} d s s^{-\tau} \widetilde{\rho}(s) \tag{25}
\end{equation*}
$$

is given by a convergent integral at small $s=c S / S_{m}$ (it always converges at large $s$ ). Hence the normalized avalanche size distribution $P(S)$ takes the simple scaling form [82],

$$
\begin{gather*}
P(S)=S_{m}^{-1} p\left(S / S_{m}\right)  \tag{26}\\
p(x)=c^{1-\tau} x^{-\tau} \widetilde{\rho}(c x) \frac{1}{\int_{0}^{\infty} d y y^{-\tau} \widetilde{\rho}(y)} \tag{27}
\end{gather*}
$$

controlled by a unique scale $S_{m}$. The scaling function $p(x)$ for the probability satisfies

$$
\begin{equation*}
\int_{0}^{\infty} p(x) d x=1, \quad \int_{0}^{\infty} x p(x) d x=\frac{1}{c \int_{0}^{\infty} d y y^{-\tau} \widetilde{\rho}(y)} \tag{28}
\end{equation*}
$$

It is itself a probability [therefore the notation $p(x)$ ].
Two known examples with $\tau<1$ are (for review see [63]): (1) the Sinai model, which corresponds to the $d=0$ limit of the RF case, i.e., $\zeta=4 / 3$, and also to the decaying Burgers equation with initial uncorrelated velocities. There the scaling function $p(s)$ is known, with $\tau=1 / 2$ and $p(s)$ $\sim s^{5 / 2} \exp \left(-\alpha s^{3}\right)$ at large $s$, (2) for $d=0$ short-range random potential, also known as Kida turbulence in Burgers literature, which corresponds to $\zeta=1$. There $p(s)=\frac{1}{2} s e^{-s^{2} / 4}$.
(ii) $2>\tau>1$. Here the total shock density in the smallmass limit $S_{\text {min }} \ll S_{m}$ is controlled by small shocks,

$$
\begin{equation*}
\rho_{0}=\int_{S_{\min }}^{\infty} \rho(S) d S \approx L^{d} m^{\rho} \frac{S_{\min }^{1-\tau}}{1-\tau} . \tag{29}
\end{equation*}
$$

The normalized size distribution can then be written for $S \gg S_{\text {min }}$ as

$$
\begin{gather*}
P(S)=C\langle S\rangle S_{m}^{-2}\left(S / S_{m}\right)^{-\tau} f\left(S / S_{m}\right),  \tag{30}\\
C^{-1}=\int_{0}^{\infty} d x x^{1-\tau} f(x) \tag{31}
\end{gather*}
$$

with $C f(x)=c^{2-\tau} \widetilde{\rho}(c x)$ using Eq. (21). Since $f$ is not normalizable, it is not a probability, hence the notation. The normalization integral (31) of $\int \operatorname{dSSP}(S)$ converges at small $S$ and does not depend on the small-scale cutoff [83]. However the integral $\int d S P(S)$ is divergent if we extend form (30) to small $S$. It means that

$$
\begin{equation*}
\langle S\rangle \sim k S_{m}^{2-\tau} S_{\min }^{\tau-1}, \tag{32}
\end{equation*}
$$

with $k=(1-\tau) /(C f(0))$. This is an approximate value, obtained setting $\int_{S_{\text {min }}}^{\infty} d S P(S)=1$ and extending form (30) down to $S_{\min }$. Note that although $\langle S\rangle=\int d S S P(S)$ is a convergent integral at small $S$, its actual value depends on the precise cutoff at small scale. This is because, for $\tau>1$, almost all
avalanches are in size of the order of the UV cutoff $S_{\text {min }}$. Despite that fact however, all moments $\left\langle S^{p}\right\rangle$ with $p>\tau-1$ are controlled by rare avalanches of size $\sim S_{m}$, the large-scale cutoff. For $\tau>1$ it is only this part of the distribution, i.e., for $S \gtrdot S_{\min }$, which is universal, up to a multiplicative constant.

Finally, in the case $\tau=1$ both phenomena are present. We will not study this case here.

## IV. CONNECTION WITH FRG FUNCTIONS

The hypothesis of a finite density of shocks implies that the functions $\hat{C}^{(n)}\left(v_{1}, \ldots, v_{n}\right)$ and $\bar{C}^{(n)}\left(v_{1}, \ldots, v_{n}\right)$ are continuous and have no ambiguities at coinciding points; i.e., choosing a given order for its arguments one can take the limit of coinciding arguments, and the result does not depend on the chosen order.

The derivatives of $\hat{C}^{(n)}\left(v_{1}, \ldots, v_{n}\right)$ however are distributions and they do contain information about shock statistics as we now show.

## A. Cusp

Let us start with second moment (12). Taking derivatives with respect to $w_{1}$ and $w_{2}$ one finds

$$
\begin{align*}
-m^{-4} & L^{-d} \Delta^{\prime \prime}\left(w_{1}-w_{2}\right) \\
= & \overline{u^{\prime}\left(w_{1}\right) u^{\prime}\left(w_{2}\right)}-1 \\
= & L^{-2 d} \overline{\sum_{i} S_{i}^{2} \delta\left(w_{1}-w_{i}\right)} \delta\left(w_{1}-w_{2}\right) \\
& \quad+L^{-2 d} \overline{\sum_{i \neq j} S_{i} S_{j} \delta\left(w_{1}-w_{i}\right) \delta\left(w_{2}-w_{j}\right)}-1, \tag{33}
\end{align*}
$$

where we have used the decomposition in shocks [Eq. (16)] and separated in the double sum the contributions from the same shock and from different shocks.

Hence the second derivative is the sum of a $\delta$ singularity and a smooth part,

$$
\begin{align*}
-\Delta^{\prime \prime}(w)= & m^{4} L^{-d} \int d S S^{2} \rho(S) \delta(w) \\
& +m^{4} L^{-d} \int d S_{1} d S_{2} S_{1} S_{2} \rho_{c}\left(S_{1}, S_{2}, w\right) \tag{34}
\end{align*}
$$

where

$$
\begin{align*}
& \rho_{c}\left(S_{1}, S_{2}, w_{1}-w_{2}\right) \\
& \quad={\bar{\sum} \sum_{i \neq j} \delta\left(S_{1}-S_{i}\right) \delta\left(S_{2}-S_{j}\right) \delta\left(w_{1}-w_{i}\right) \delta\left(w_{2}-w_{j}\right)^{c} .}^{c} . \tag{35}
\end{align*}
$$

Note that $\rho(S)$ and the connected joined two-shock size density $\rho_{c}\left(S_{1}, S_{2}\right)$ are proportional to the size of the system. Integrating around zero yields [84]

$$
\begin{equation*}
-2 \Delta^{\prime}\left(0^{+}\right)=m^{4} L^{-d} \int d S S^{2} \rho(S)=m^{4} \frac{\left\langle S^{2}\right\rangle}{\langle S\rangle} \tag{36}
\end{equation*}
$$

using Eq. (21). This provides a rather nice interpretation of the FRG function $\Delta(w)$ in terms of shocks: the cusp gives directly the second moment of the shock size and the remaining smooth part of the function gives the correlation of the sizes of two (distinct) shocks at different points. A nonlocal generalization of this formula is given in Eq. (F12).

Let us now verify the scaling for small $m$. From the above one finds

$$
\begin{equation*}
-2 \Delta^{\prime}\left(0^{+}\right)=\widetilde{c} m^{4} S_{m}=c \widetilde{c} m^{\epsilon-\zeta} \tag{37}
\end{equation*}
$$

with $\widetilde{c}:=c^{-1} \int_{0}^{\infty} d s s^{2-\tau} \widetilde{\rho}(s)=\int_{0}^{\infty} d s s^{2-\tau} f(s) / \int_{0}^{\infty} d s s^{1-\tau} f(s)$ (the latter for $\tau>1$ only). This has the expected fixed-point scaling (14) with $-\widetilde{\Delta}^{\prime}\left(0^{+}\right)=c \widetilde{c} /\left(2 A_{d}\right)$. Below, we will make the convenient choice $\tilde{c}=2$ to fix the arbitrariness in definition of $S_{m}$.

This correspondence can be extended to higher moments. For illustration, let us consider the third moment,

$$
\begin{align*}
\overline{u^{\prime}\left(w_{1}\right) u^{\prime}\left(w_{2}\right) u^{\prime}\left(w_{3}\right)}= & L^{-3 d} \int d S S^{3} \rho(S) \delta\left(w_{1}-w_{2}\right) \delta\left(w_{1}-w_{3}\right)+\left[\delta\left(w_{1}-w_{2}\right) L^{-3 d} \overline{\sum_{i \neq j} S_{i}^{2} S_{j} \delta\left(w_{1}-w_{i}\right) \delta\left(w_{3}-w_{j}\right)}+2 \text { perm }\right] \\
& +L^{-3 d} \overline{\sum_{i \neq j \neq k \neq i} S_{i} S_{j} S_{k} \delta\left(w_{1}-w_{i}\right) \delta\left(w_{2}-w_{j}\right) \delta\left(w_{3}-w_{k}\right)} \tag{38}
\end{align*}
$$

From this expression one can integrate $\int_{0}^{w} d w_{1} \int_{0}^{w} d w_{2} \int_{0}^{w} d w_{3}$ and obtain

$$
\begin{equation*}
\overline{[u(w)-u(0)]^{3}}=L^{-3 d} \int d S S^{3} \rho(S) w+O\left(w^{2}\right) \tag{39}
\end{equation*}
$$

This can be generalized to any order, as discussed in Appendix A.

Note that in dimension $d=0$ the field $m^{2}(w-u(w))$ identifies with the velocity field of a one-dimensional fluid which evolves in time $t=m^{-2}$ according to the Burgers equation with random initial conditions. The linear cusp of the third moment [Eq. (39)] is the analog for Burgers turbulence to the famous exact result of Kolmogorov for Navier Stokes in the inertial range, both models exhibiting similarities. We will not discuss these connections further here (see Refs. [62,63]), but since the manifold problem is a $d>0$ generalization of the Burgers equation, we will call Kolmogorov moments such as Eq. (39) and their generalizations.

## B. Kolmogorov cumulants and generating function

We now generalize the results of the previous section and construct a very useful generating function which resums all shock-size moments.

Since we are interested in the jump sizes, it is convenient to define the Kolmogorov cumulants (by analogy with the famous K41 paper [85], as discussed above),

$$
\begin{gather*}
\overline{[u(w)-w-u(0)]^{n}}{ }^{c}=m^{-2 n} L^{-(n-1) d} K^{(n)}(w),  \tag{40}\\
K^{(2)}(w)=2(\Delta(0)-\Delta(w)), \tag{41}
\end{gather*}
$$

where all $K^{(n)}$ have a large- $L$ limit and $K^{(1)}=0$. We find that they are proportional to moments of the shock-size distribution,

$$
\begin{align*}
K^{(n)}(w) & =m^{2 n} L^{-d} \int d S S^{n} \rho(S)|w|(\operatorname{sgn} w)^{n}+O\left(w^{2}\right) \\
& =m^{2 n} \frac{\left\langle S^{n}\right\rangle}{\langle S\rangle}|w|(\operatorname{sgn} w)^{n}+O\left(w^{2}\right) \tag{42}
\end{align*}
$$

Note that the leading small-w coefficient, the cusp, is the same for $\overline{[u(w)-w-u(0)]^{n}}{ }^{c}, \overline{[u(w)-w-u(0)]^{n}}$, and $\overline{[u(w)-u(0)]^{n}}$. We thus study the generating function

$$
\begin{align*}
\overline{e^{\lambda L^{d}[u(w)-w-u(0)]}}-1 & =\sum_{n=2}^{\infty} \frac{\lambda^{n}}{n!} L^{n d} \overline{[u(w)-w-u(0)]^{n}} \\
& =\sum_{n=2}^{\infty} \frac{\lambda^{n}}{n!} \int d S S^{n} \rho(S)|w|(\operatorname{sgn} w)^{n}+O\left(w^{2}\right) . \tag{43}
\end{align*}
$$

The following generating function hence has a finite large- $L$ limit:

$$
\begin{align*}
G(\lambda): & \left.=L^{-d} \overline{\left(e^{\lambda L^{d}[u(w)-w-u(0)]}\right.}-1\right) \\
= & L^{-d} \int \rho(S)(\cosh (\lambda S)-1)|w| \\
& +L^{-d} \int \rho(S)(\sinh (\lambda S)-\lambda S) w+O\left(w^{2}\right) . \tag{44}
\end{align*}
$$

For positive $w>0$ this yields

$$
\begin{gather*}
G(\lambda)=\hat{Z}(\lambda) w+O\left(w^{2}\right),  \tag{45}\\
\hat{Z}(\lambda)=\frac{1}{\langle S\rangle}\left(\left\langle e^{\lambda S}\right\rangle-1-\lambda\langle S\rangle\right) . \tag{46}
\end{gather*}
$$

Hence by computing the cusp of $G(\lambda)$ one has direct access to the characteristic function of the shock-size distribution. Formula (44) is also derived in Appendix A by another method. In the following we reserve the notation
$Z(\lambda):=\lambda+\hat{Z}(\lambda)$ and the notation $\tilde{Z}$ for the rescaled version of $Z$ (see below).

## V. TREE AND IMPROVED TREE CALCULATION

We now compute the cumulants $\hat{C}\left(w_{1}, \ldots, w_{n}\right)$ and from them, the Kolmogorov cumulants and the generating function $G(\lambda)$. We use several methods which give equivalent results: a calculation using replicas, one without replicas, and a graphical representation using static or dynamic graphs.

In this section we obtain the form of $G(\lambda)$ to lowest order in the $\epsilon$ expansion (i.e., zeroth order). It is essentially a treelevel calculation, although some loops can be incorporated, as we discuss. In a diagrammatic language, one defines suitable trees where the building blocks contain loops, which are resummed here. Since we restrict to a uniform $w$, i.e., to zero external momentum, all trees carry zero momentum $q=0$. This is why the calculations to this order look very similar to calculations on a $d=0$ toy model.

## A. Method using replicas

The cumulants $\hat{C}$ are contained in the generating functional $W[w]$ for connected correlations (for more definitions see Ref. [62]). We focus on a uniform $w_{x}=w$, in which case $W[w]=L^{d} W(w)$ and

$$
\begin{align*}
W(w)= & \frac{m^{2}}{2 T} \sum_{a} w_{a}^{2}+\frac{1}{2!T^{2}} \sum_{a b} R\left(w_{a b}\right) \\
& +\sum_{p=3}^{\infty} \frac{1}{p!T^{p}} \sum_{a_{1} \cdots a_{p}} \hat{S}^{(p)}\left(w_{a_{1}}, \ldots, w_{a_{p}}\right), \tag{47}
\end{align*}
$$

where $w_{a b}=w_{a}-w_{b}$. It can be obtained via a Legendre transform from the effective action per unit volume, i.e., $\Gamma(u)$ $=L^{-d} \Gamma[u]$ for a uniform field $u_{x}=u$,

$$
\begin{gather*}
W(w)+\Gamma(u)=\frac{m^{2}}{T} \sum_{a} u_{a} w_{a},  \tag{48}\\
\Gamma(u)= \\
\frac{m^{2}}{2 T} \sum_{a} u_{a}^{2}-\frac{1}{2!T^{2}} \sum_{a b} R\left(u_{a b}\right)  \tag{49}\\
\\
\quad-\sum_{p=3}^{\infty} \frac{1}{p!T^{p}} \sum_{a_{1} \cdots a_{p}} S^{(p)}\left(u_{a_{1}}, \ldots, u_{a_{p}}\right) .
\end{gather*}
$$

The $\Gamma$ cumulants are naturally computed in the field theory as the sum of all one-particle irreducible (1PI) graphs, as in [36]. Each $S^{(p)}$ can be computed in an expansion in powers of $R$ (at $T=0$ it is the usual loop expansion). They have the property that to lowest order $S^{(3)} \sim R^{3}$ and $S^{(p)} \sim R^{p}$. Hence near $d=4$ where $R \sim \epsilon, S^{(p)} \sim \epsilon^{p}$. Note that the $\hat{S}^{(p)}$ have a different counting, e.g., $S^{(3)} \sim \epsilon^{2}$. An important property, arising from the Legendre transform, is that $W(w)$ can be written as a sum of tree diagrams with vertices made of $\Gamma$.

The improved tree approximation consists in setting the higher cumulants of the effective action $\Gamma$ to zero, $S^{(p)}=0$ for
$p>2$, and then perform the Legendre transform to obtain the effective action $W(w)$ and $\hat{S}^{(p)}$, i.e., to use

$$
\begin{equation*}
\Gamma_{\text {tree }}(u)=\frac{m^{2}}{2 T} \sum_{a} u_{a}^{2}-\frac{1}{2!T^{2}} \sum_{a b} R\left(u_{a b}\right) . \tag{50}
\end{equation*}
$$

It is "improved" as one keeps the exact two-replica part $R$, which itself has a loop expansion. For instance, from the above discussion one has $\Gamma=\Gamma_{\text {tree }}+O\left(\epsilon^{3}\right)$.

To perform the Legendre transform one must invert the relation $m^{2} w_{a} / T=\Gamma_{\text {tree }}^{\prime}\left(u_{a}\right)$, i.e., find the function $u_{a}(w)$, solution of

$$
\begin{equation*}
u_{a}(w)=w_{a}+\frac{1}{T m^{2}} \sum_{b} R^{\prime}\left(u_{a}(w)-u_{b}(w)\right) . \tag{51}
\end{equation*}
$$

This solution can be expanded in the number of free replica sums,

$$
\begin{equation*}
u_{a}(w)=w_{a}+\sum_{a_{1}} u_{a, a_{1}}^{(1)}(w)+\sum_{a_{1}, a_{2}} u_{a, a_{1}, a_{2}}^{(2)}(w)+\cdots \tag{52}
\end{equation*}
$$

which is also an expansion in powers of $R$. Inserting in Eq. (51) generates recursion relations,

$$
\begin{gather*}
u_{a, a_{1}}^{(1)}=\frac{1}{\operatorname{Tm}^{2}} R^{\prime}\left(w_{a a_{1}}\right),  \tag{53}\\
u_{a, a_{1}, a_{2}}^{(2)}=\frac{1}{\operatorname{Tm}^{2}} \operatorname{sym}_{a_{1}, a_{2}} R^{\prime \prime}\left(w_{a a_{1}}\right)\left(u_{a, a_{2}}^{(1)}-u_{a_{1}, a_{2}}^{(1)}\right),  \tag{54}\\
u_{a, a_{1}, a_{2}, a_{3}}^{(3)}=\frac{1}{\operatorname{Tm}^{2}} \operatorname{sym}_{a_{1}, a_{2}, a_{3}}\left[\frac{1}{2} R^{\prime \prime \prime}\left(w_{a a_{1}}\right)\left(u_{a, a_{2}}^{(1)}-u_{a_{1}, a_{2}}^{(1)}\right)\right. \\
\left.\times\left(u_{a, a_{3}}^{(1)}-u_{a_{1}, a_{3}}^{(1)}\right)+R^{\prime \prime}\left(w_{a a_{1}}\right)\left(u_{a, a_{2}, a_{3}}^{(2)}-u_{a_{1}, a_{2}, a_{3}}^{(2)}\right)\right] . \tag{55}
\end{gather*}
$$

This is easy to automatize using MATHEMATICA, the combinatorics being similar to expanding $R^{\prime}\left(\Sigma_{p} b_{p} x^{p}\right)$, with the additional difficulty of attributing new labels to repeated indices. Once the terms in expansion (52) are known, since $m^{2} u_{a} / T=W^{\prime}\left(w_{a}\right)$, one obtains the derivatives of the energy cumulants,

$$
\begin{gather*}
u_{a, a_{1}}^{(1)}(w)=\partial_{w_{a}} R\left(w_{a a_{1}}\right)  \tag{56}\\
u_{a, a_{1}, \ldots, a_{n}}^{(n)}(w)=\frac{1}{(n-1)!} \partial_{w_{a}} \hat{S}^{(n+1)}\left(w_{a}, w_{a_{1}}, \ldots, w_{a_{n}}\right) \tag{57}
\end{gather*}
$$

Upon further derivation, we obtain the cumulants of the force, $\hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right)=(-1)^{n} \partial_{w_{1}} \cdots \partial_{w_{n}} \hat{S}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$, as

$$
\begin{align*}
& \hat{C}^{(n+1)}\left(w_{a}, w_{a_{1}}, \ldots, w_{a_{n}}\right) \\
& \quad=(-1)^{n}(n-1)!\partial_{w_{a_{1}}} \cdots \partial_{w_{a_{n}}} u_{a, a_{1}, \ldots, a_{n}}^{(n)}(w) . \tag{58}
\end{align*}
$$

We display the obtained explicit forms for the third and fourth cumulants,

$$
\begin{equation*}
\hat{S}^{(3)}\left(w_{1}, w_{2}, w_{3}\right)=\frac{3}{m^{2}} \operatorname{sym}_{123} R^{\prime}\left(w_{12}\right) R^{\prime}\left(w_{13}\right), \tag{59}
\end{equation*}
$$

$$
\begin{equation*}
\hat{S}^{(4)}\left(w_{1}, w_{2}, w_{3}, w_{4}\right)=\frac{12}{m^{4}} \operatorname{sym}_{1234} R_{12}^{\prime \prime}\left(R_{13}^{\prime}-R_{23}^{\prime}\right) R_{14}^{\prime} \tag{60}
\end{equation*}
$$

for the energy, already given in [62]. We use the shorthand notation $R_{12}=R\left(w_{12}\right)$. They have a simple graphical representation in terms of tree diagrams with $R$ vertices. For the corresponding force cumulants this gives [with everywhere $\left.\Delta(u)=-R^{\prime \prime}(u)\right]$

$$
\begin{align*}
-\hat{C}^{(3)}\left(w_{1}, w_{2}, w_{3}\right)= & \frac{6}{m^{2}} \operatorname{sym}_{123} \Delta^{\prime}\left(w_{12}\right) \Delta\left(w_{13}\right), \\
\hat{C}^{(4)}\left(w_{1}, w_{2}, w_{3}, w_{4}\right)= & \frac{12}{m^{4}} \operatorname{sym}_{1234}\left[2 \Delta_{12} \Delta_{13}^{\prime} \Delta_{14}^{\prime}\right. \\
& +\Delta_{12} \Delta_{23}^{\prime} \Delta_{14}^{\prime}+2 \Delta_{12} \Delta_{13}^{\prime} \Delta_{34}^{\prime} \\
& \left.+\Delta_{13} \Delta_{14} \Delta_{12}^{\prime \prime}-\Delta_{13} \Delta_{24} \Delta_{12}^{\prime \prime}\right] \tag{61}
\end{align*}
$$

with $\Delta_{12}=\Delta\left(w_{12}\right)$ and so on. Note that this expression for the third force cumulant was tested against numerics near depinning in Ref. [66] (within the improved tree approximation, relation (61) is the same for statics and depinning with, however, different functions $\Delta(w)$ in each case).

The present method allows us to compute significantly higher cumulants using MATHEMATICA. We do not display the full form of $\hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$, which are very tedious, but we give the resulting Kolmogorov cumulants,

$$
\begin{gather*}
\hat{K}^{(2)}(w)=2[\Delta(0)-\Delta(w)],  \tag{62}\\
m^{2} \hat{K}^{(3)}(w)=-12 \Delta^{\prime}(w)[\Delta(0)-\Delta(w)],  \tag{63}\\
m^{4} \hat{K}^{(4)}(w)=120[\Delta(0)-\Delta(w)] \Delta^{\prime}(w)^{2} \\
-48[\Delta(0)-\Delta(w)]^{2} \Delta^{\prime \prime}(w),  \tag{64}\\
m^{6} \hat{K}^{(5)}(w)=-80[\Delta(0)-\Delta(w)]\left\{21 \Delta^{\prime}(w)^{3}\right. \\
-24[\Delta(0)-\Delta(w)] \Delta^{\prime}(w) \Delta^{\prime \prime}(w) \\
 \tag{65}\\
\left.+2[\Delta(0)-\Delta(w)]^{2} \Delta^{\prime \prime \prime}(w)\right\}, \\
m^{8} \hat{K}^{(6)}(w)=-480[\Delta(0)-\Delta(w)]\left(-63 \Delta^{\prime}(w)^{4}\right. \\
+138[\Delta(0)-\Delta(w)] \Delta^{\prime}(w)^{2} \Delta^{\prime \prime}(w)-22[\Delta(0) \\
-\Delta(w)]^{2} \Delta^{\prime}(w) \Delta^{\prime \prime \prime}(w)+[\Delta(0)-\Delta(w)]^{2}  \tag{66}\\
\left.\times\left\{-18 \Delta^{\prime \prime}(w)^{2}+[\Delta(0)-\Delta(w)] \Delta^{\prime \prime \prime \prime}(w)\right\}\right) .
\end{gather*}
$$

We expect, from the assumption of a shock density (see the discussion above), that the $\hat{C}^{(n)}$ are continuous functions of their arguments. We checked this explicitly. Hence there are no ambiguities at coinciding points, i.e., to perform the limits one can choose any order for the arguments, with the result being independent of the chosen order.

One first checks that the values at zero vanish,

$$
\begin{equation*}
{\overline{[u(w)-w]^{n}}}^{c}={\overline{u(0)^{n}}}^{c} \propto \hat{C}^{(n)}(0,0, \ldots, 0)=0 . \tag{67}
\end{equation*}
$$

Hence the distribution of the center-of-mass position in the quadratic well is Gaussian to lowest order in $\epsilon$, i.e., to the improved tree approximation. To this order the calculation is the same in statics and dynamics. As we see below corrections to the Gaussian arise to one loop order (see [51,86] for some results on the deviations of the distribution of the interface width to the Gaussian at depinning). At quasistatic depinning Eq. (67) also gives the cumulants of the distribution of the critical force [64]. Hence it is also Gaussian to this order. This is consistent with Ref. [53] where deviations from Gaussian were found and computed to one loop. In fact, Eq. (67) validates these calculations within the present well-controlled setting of a quadratic well.

We now expand the above result for $\hat{K}^{(n)}(w)$ to small argument and find

$$
\begin{equation*}
\hat{K}^{(n)}(w)=a_{n}\left(-\Delta^{\prime}\left(0^{+}\right)\right)^{n-1} w(\operatorname{sgn} w)^{n} m^{4-2 n} \tag{68}
\end{equation*}
$$

with

$$
\begin{gather*}
a_{2}=2, \quad a_{3}=12, \quad a_{4}=120, \quad a_{5}=1680 \\
a_{6}=30240, \ldots \tag{69}
\end{gather*}
$$

Note that $a_{n+1} / a_{n}=4 n-2$. Defining $b_{n}:=a_{n} / n$ !, the first coefficients are

$$
\begin{equation*}
b_{2}=1, \quad b_{3}=2, \quad b_{4}=5, \quad b_{5}=14, \quad b_{6}=42 \tag{70}
\end{equation*}
$$

## B. Method without replicas

An equivalent method to this order is as follows. One notices that each cumulant $\hat{C}^{(n)}$ is computed to lowest nonvanishing order in $R$ (or $\Delta$ ). Hence it is formally equivalent to start from a replicated action $\mathcal{S}$ containing $R$ only (i.e., a bare disorder with the substitution $R_{0} \rightarrow R$ ) and compute the moments of $u(w)-w$ each to lowest order in perturbation theory. Hence the calculation of the previous paragraph is equivalent to the following one in dimension $d=0$ : denote $u(w)$ the minimum of the toy model,

$$
\begin{equation*}
\frac{1}{m^{2}} H_{\text {toy }}=\frac{1}{2}(u-w)^{2}+\frac{\lambda}{m^{2}} V(u), \tag{71}
\end{equation*}
$$

where $V(u)$ is a Gaussian random potential of correlator $R(u)$. The minimum satisfies

$$
\begin{equation*}
u(w)=w+\frac{\lambda}{m^{2}} F(u(w)), \tag{72}
\end{equation*}
$$

with $F(u)=-V^{\prime}(u)$ and $\lambda$ has been introduced to count the powers in $V$. Compute each moment defined as

$$
\begin{align*}
&(-1)^{n} m^{-2 n} L^{-(n-1) d} \hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right) \\
&:={\overline{\left(u\left(w_{1}\right)-w_{1}\right) \cdots\left(u\left(w_{n}\right)-w_{n}\right)}}^{c} \\
& \quad=\left(\frac{\lambda}{m^{2}}\right)^{n} \overline{F\left(u\left(w_{1}\right)\right) \cdots F\left(u\left(w_{n}\right)\right)} \tag{73}
\end{align*}
$$

perturbatively in $\lambda$ (in fact, in $\lambda / m^{2}$ ) to lowest nontrivial order, which is $O\left(\lambda^{2 n-2}\right)$. The factor $L^{-(n-1) d}$ can be omitted since $d=0$. One uses iteratively (72) and the Wick theorem with $\overline{F(w) F\left(w^{\prime}\right)}=\Delta\left(w-w^{\prime}\right)$ with all other cumulants of $F(w)$ set to zero (i.e., a Gaussian $F$ ). Using mathematica we have reproduced most of the results of the previous section.

Note that this is different from the standard perturbative expansion of the toy model which yields dimensional reduction, i.e., a trivial perturbation expansion involving only $\Delta(0)=-R^{\prime \prime}(0)$. The difference is that one computes here cumulants $\hat{C}^{(n)}$ at different points and only at the end the limit of coinciding arguments is taken using a nonanalytic $\Delta(u)$ with a cusp. Nevertheless ambiguous terms at intermediate stages may be generated. To the lowest order studied here, those represent possible contributions to $\bar{C}^{(n)}$ which depend on a smaller number of points and they cancel in the calculation of the connected correlations $\hat{C}^{(n)}$.

## C. Graphical representation

It is useful for the following to give a graphical representation of the results of the two previous sections. Define

$$
\begin{align*}
\mathcal{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right) & =(-1)^{n} m^{-2 n} \hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right) \\
& =L^{(n-1) d}{\overline{\left(u\left(w_{1}\right)-w_{1}\right) \cdots\left(u\left(w_{n}\right)-w_{n}\right)}}^{c} . \tag{74}
\end{align*}
$$

For instance, we write the two-point correlation as

$$
\begin{equation*}
\left.\left.\mathcal{C}^{(2)}\left(w_{1}, w_{2}\right)=\emptyset \quad\right\rfloor \quad\right\rfloor=\frac{1}{m^{4}} \Delta\left(w_{1}-w_{2}\right) \tag{75}
\end{equation*}
$$

The graphical notations here and in all diagrams of this type are as follows: there are external legs with points on the top labeled by integers 1 to $n$, corresponding to positions $w_{1}$ to $w_{n}$ and external fields $u\left(w_{i}\right)-w_{i}$, here $n=2$. The $\Delta$ (or equivalently $R$ ) vertices are double vertices (nonlocal in $w$ ) with two points joined by a dotted line and can be interpreted equivalently as in the statics or in the dynamics (we have checked equivalence to the order we are working). In the statics, they are $R$ vertices, giving $-R^{\prime \prime}$ after contraction with the $u$ fields. In the dynamic formulation, they are $\Delta$ vertices and the two lines exiting a vertex are directed to the top and end up being equal to a static propagator: in the real dynamics they are response function, which usually are denoted with an arrow; we do not show the arrow here but they are always implicitly toward the top of the picture. Here they are always evaluated at zero frequency. Thus the lines are static propagators, evaluated here all at $q=0$, hence giving a factor of $1 / \mathrm{m}^{2}$. For a generalization to nonzero external momenta see Appendix F.

With these diagrammatic rules there is a single diagram to represent the third cumulant,

$$
\begin{equation*}
\mathcal{C}^{(3)}=\left\lvert\, \ldots \ldots-\ldots-\cdots=\frac{6}{m^{8}} \operatorname{sym} \Delta_{12} \Delta_{13}^{\prime}\right., \tag{76}
\end{equation*}
$$

where the combinatoric factor comes from the six inequivalent ways to assign three labels to external legs. The result
agrees with Eq. (61). Similarly there are five diagrams for the fourth cumulant,

which reproduce each term in Eq. (61), all with a factor 4 ! $=24$ and a factor $1 / 2$ from the three diagrams symmetric in exchange of a pair among the four labels.

To compute the $n$th cumulant of $(u(w)-w-u(0))$, we must evaluate $\mathcal{C}^{(n)}$ for $w_{i} \rightarrow w$ minus $w_{i} \rightarrow 0$, for each $i$. Writing this operation as $\mathcal{K} f\left(w_{1}, \ldots, w_{n}\right):=\prod_{i=1}^{n}\left[\left(w_{i} \rightarrow w\right)\right.$ $\left.-\left(w_{i} \rightarrow 0\right)\right] f\left(w_{1}, \ldots, w_{n}\right)$, this gives

$$
\begin{align*}
\hat{K}^{(2)}(w)=\mathcal{K} \mid \quad & =2[\Delta(0)-\Delta(w)] \\
& \approx-2 \Delta^{\prime}\left(0^{+}\right)|w| \tag{78}
\end{align*}
$$

where from now on we extract the $m$ dependence from the lines of the graphs, which are hence set to 1 . Note that there are four choices to assign 0 or $w$ to each leg, hence four terms. For the third cumulant one finds

$$
\begin{align*}
m^{2} \hat{K}^{(3)}(w)=\mathcal{K} \mid \quad \bullet \cdots & =-12[\Delta(0)-\Delta(w)] \Delta^{\prime}(w) \\
& \approx 12 \Delta^{\prime}\left(0^{+}\right)^{2} w \tag{79}
\end{align*}
$$

which produces $6 \times 4=24$ terms.
Similarly one finds

$$
\begin{align*}
m^{4} \hat{K}^{(4)}(w)= & \mathcal{K}\left[\mathcal{C}^{(4)}\right] \\
= & 120[\Delta(0)-\Delta(w)] \Delta^{\prime}(w)^{2} \\
& -48[\Delta(0)-\Delta(w)]^{2} \Delta^{\prime \prime}(w) \\
\approx & 120 \Delta^{\prime}\left(0^{+}\right)|w| \tag{80}
\end{align*}
$$

in terms of the five diagrams in Eq. (77). These results agree with those of the two previous sections. Here too the contractions are unambiguous, as long as all $w_{i}$ are different. It would appear naively that it is equivalent to apply $\mathcal{K}$ to an unsymmetrized expression or to its symmetrized form. This is not true, in fact, because of ambiguities at coinciding points. One must be very careful to apply $\mathcal{K}$ to the symmetrized expression and then take the limit, i.e.,

$$
\begin{align*}
(-1)^{n} \hat{K}^{(n)}(w)= & \mathcal{K} \hat{C}^{(n)} \\
= & \lim _{\delta_{i}, \delta_{i}^{\prime} \rightarrow 0} \prod_{i=1}^{n}\left[\left(w_{i} \rightarrow w+\delta_{i}\right)-\left(w_{i} \rightarrow \delta_{i}^{\prime}\right)\right] \\
& \times \hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right), \tag{81}
\end{align*}
$$

where the $\delta_{i}, \delta_{i}^{\prime}$ are taken to zero in such a manner that all arguments remain distinct and the order is fixed, the $\hat{C}^{(n)}$ being fully symmetric functions of their arguments. We have checked here and in the one-loop extension given below that
the result does not depend on the order as required by the continuity of the force correlations.

## D. Recursion relation and resummation

It is clear from the above results [Eqs. (78)-(80) and (66)] that the leading term in the expansion of $\hat{K}^{(n)}(w)$ in small $w$ is obtained from those terms which contain a single factor of $\Delta(w)-\Delta(0)$ since the latter is of order $w$. There are also terms such as $[\Delta(w)-\Delta(0)]^{p}$ with $p \geq 2$ in the expressions for $\hat{K}^{(n)}(w)$, but they are of higher order [since, e.g., $\Delta^{\prime \prime}(w)$ $\rightarrow \Delta^{\prime \prime}\left(0^{+}\right)$has a finite limits at $\left.w=0^{+}\right]$. Therefore, final result (68) for the coefficient of the cusp depends only on the part of $\hat{C}^{(n)}$ of the form $[\Delta(w)-\Delta(0)] \Delta^{\prime}(w)^{n-2}$ and is the diagram with one and only one terminating (lower) $\Delta$ vertex. This is not the case, e.g., for the third and fifth diagrams in Eq. (77) which have $p=2$.

Let us call $\mathrm{C}^{(n)}:=(-1)^{n} \hat{C}^{(n)}$ the part which contains only one factor of $\Delta\left(w_{i}\right)$ for some $i$ [we do not write explicitly the $-\Delta(0)$ but remind that every $\Delta(w)$ comes with a $-\Delta(0)]$. If we restrict to this part we can write a recursion relation using the tree structure. Indeed, to construct $\mathbf{C}^{(n)}$ we can either:
(i) take $\mathrm{C}^{(n-1)}$ and a single (as yet unlabeled) leg $u(w)$ $-w$ and attach a $\Delta$ vertex at the bottom (hence without derivatives) to each of these two elements. It results in a diagram with again only one lower $\Delta$, e.g.,


The combinatorial factor is $n$ (here $n=4$ ) for labeling the newly added leg.
(ii) starting from the fourth cumulant, take two already constructed cumulants (trees) of size $l \geqslant 2$ and $(n-l) \geqslant 2$ and glue them together using again a $\Delta$ vertex at the bottom, e.g.,


Now the combinatorial factor is $\binom{n}{l}$ for choosing the group of indices for the first and second elements, with the restriction that $l<n-l$. If $n=l / 2$, an additional factor of $1 / 2$ appears. It is more systematic to sum over all pairs, i.e., $l$ $=2, \ldots, n-2$, and divide this sum by 2 . This can be summarized in the following schematic recursion relation for $\mathrm{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$, which by construction is a function of the set of $\Delta_{i j}=\Delta\left(w_{i}-w_{j}\right)$ and $\Delta_{i j}^{\prime}=\Delta^{\prime}\left(w_{i}-w_{j}\right)$,

$$
\begin{align*}
\mathrm{C}^{(n)}= & \sum_{i<j} \frac{\partial \mathrm{C}^{(n-1)}}{\delta \Delta_{i j}} \Delta_{i j}^{\prime}\left(\Delta_{i, n+1}-\Delta_{j, n+1}\right) n \\
& +\sum_{l=2}^{n-2} \sum_{i<j} \frac{\partial \mathrm{C}^{(l)}}{\delta \Delta_{i j}} \sum_{k<m} \frac{\partial \mathrm{C}^{(n-l)}}{\delta \Delta_{k m}} \Delta_{i j}^{\prime} \Delta_{k m}^{\prime} \\
& \times\left(\Delta_{i, k}+\Delta_{j, m}-\Delta_{i, m}-\Delta_{j, k} \frac{1}{2}\binom{n}{l} .\right. \tag{84}
\end{align*}
$$

We have temporarily set the mass $m=1$ to not burden notations. We remark that the first term has formally the same form as the others, except that one of the derivative
terms does not exist. Since $n=\binom{n}{1}=\binom{n}{n-1}=\frac{1}{2}\left[\binom{n}{1}+\binom{n}{n-1}\right]$, and when adding the two new terms $l=1$ and $l=n-1$, the combinatorics is identical. We can indeed take advantage of this feature. To do so, we go to the Kolmogorov cumulants and set as in Eq. (68) and for $w>0, \quad m^{2 n-4} \hat{K}^{(n)}(w) \approx a_{n}\left[-\Delta^{\prime}\left(0^{+}\right)\right]^{n-2}[\Delta(0)-\Delta(w)], \quad$ i.e., $m^{2 n-4} \hat{K}^{(l) \prime}\left(0^{+}\right)=a_{l}\left[-\Delta^{\prime}\left(0^{+}\right)\right]^{l-1}$. We note that gluing $\hat{K}^{(l)}$ and $\hat{K}^{(n-l)}$ yields $\hat{K}^{(l) \prime}\left(0^{+}\right) \hat{K}^{(n-l) \prime}\left(0^{+}\right)[\Delta(0)-\Delta(w)]$ times the combinatoric factor $\binom{n}{l}$. Indeed one glued part should be at $w_{i}$ $=w$ and the other one at $w=0$. Hence we can convert Eq. (84) into a recursion relation for the $a_{n}$. Inserting ansatz (68), the above recursion relation becomes

$$
\begin{equation*}
a_{n}=\sum_{l=1}^{n-1}\binom{n}{l} a_{l} a_{n-l}, \quad a_{1}=1 . \tag{85}
\end{equation*}
$$

We note that it correctly reproduces series (69). The series $b_{n}:=a_{n} / n!$ then satisfies the recursion relation,

$$
\begin{equation*}
b_{n}=\sum_{l=1}^{n-1} b_{l} b_{n-l} \tag{86}
\end{equation*}
$$

Hence the generating function

$$
\begin{equation*}
\widetilde{Z}_{0}(\lambda):=\sum_{n=1}^{\infty} b_{n} \lambda^{n} \tag{87}
\end{equation*}
$$

satisfies the equation

$$
\begin{equation*}
\tilde{Z}(\lambda)=\lambda+\tilde{Z}(\lambda)^{2} \tag{88}
\end{equation*}
$$

It has solution

$$
\begin{equation*}
\tilde{Z}(\lambda)=\tilde{Z}_{0}(\lambda):=\frac{1}{2}(1-\sqrt{1-4 \lambda}) \tag{89}
\end{equation*}
$$

Let us now recall that from Eqs. (40) and (68) one has at small $w$,

$$
\begin{align*}
\frac{1}{n!} \overline{[u(w)-w-u(0)]^{n}} & =L^{(1-n) d} b_{n}\left[-\frac{\Delta^{\prime}\left(0^{+}\right)}{m^{4}}\right]^{n-1} \\
& \times w(\operatorname{sgn}(w))^{n-1}+O\left(w^{2}\right) \tag{90}
\end{align*}
$$

Following the discussion of Sec. IV B we find:
(i) The generating function defined in Eqs. (44) and (45), for $w>0$ from $G(\lambda)=\hat{Z}(\lambda) w+O\left(w^{2}\right)$ satisfies the scaling form

$$
\begin{equation*}
\hat{Z}(\lambda)=\frac{1}{S_{m}} \widetilde{Z}\left(\lambda S_{m}\right)-\lambda \tag{91}
\end{equation*}
$$

where from now on we define $\left[\right.$ recall $\left.\Delta^{\prime}\left(0^{+}\right)<0\right]$,

$$
\begin{equation*}
S_{m}=\frac{\left|\Delta^{\prime}\left(0^{+}\right)\right|}{m^{4}} \tag{92}
\end{equation*}
$$

Since Eq. (36) is exact, Eqs. (91) and (93) are valid to all orders, as they amount to the choice $\tilde{Z}(\lambda)=\lambda+\lambda^{2}+O\left(\lambda^{3}\right)$, i.e., they fix the coefficient $\lambda^{2}$ in $\widetilde{Z}$ to 1 . Scaling means that in the limit $m \rightarrow 0$ the function $\widetilde{Z}$ becomes $m$ independent,
whose validity is discussed below. In summary, using exact relation (36) the definition of $S_{m}$ chosen in this paper is

$$
\begin{equation*}
S_{m}:=\frac{\left\langle S^{2}\right\rangle}{2\langle S\rangle} \tag{93}
\end{equation*}
$$

(ii) In the (improved) tree level approximation the function $\tilde{Z}(\lambda)$ satisfies the self-consistent equation [Eq. (88)], with solution $\widetilde{Z}_{0}(\lambda)$ given in Eq. (89), but with $\Delta^{\prime}(0)$ renormalized. Hence we find the generating function of the avalanche size distribution,

$$
\begin{align*}
\hat{Z}(\lambda) & :=\frac{1}{\langle S\rangle}\left[\left\langle e^{\lambda S}\right\rangle-1-\lambda\langle S\rangle\right] \\
& =\frac{1}{S_{m}}\left[\frac{1}{2}\left(1-\sqrt{1-4 \lambda S_{m}}\right)-\lambda S_{m}\right] \tag{94}
\end{align*}
$$

with $S_{m}$ given by Eq. (93). Note that the presence of the factor $1 /\langle S\rangle$ indicates that one can only obtain information about the distribution $P(S)$ up to an unknown multiplicative factor $\langle S\rangle$, consistent with the discussion in Sec. III C. We now analyze this result.

## E. Moments of avalanche sizes and universal ratios

The easiest quantities to extract from Eq. (90) are the moments of the size distribution $P(S)$. From Eq. (94) one finds for $n \geqslant 2$,

$$
\begin{equation*}
\frac{\left\langle S^{n}\right\rangle}{\langle S\rangle}=a_{n} S_{m}^{n-1} \tag{95}
\end{equation*}
$$

The coefficients $a_{n}$ and $b_{n}$ can be calculated explicitly for $n>1$,

$$
\begin{gather*}
a_{n}=\left.\left(\frac{\partial}{\partial \lambda}\right)^{n}\left[\frac{1}{2}(1-\sqrt{1-4 \lambda})-\lambda\right]\right|_{\lambda=0}=2 \frac{(2 n-3)!}{(n-2)!}, \\
b_{n}=\frac{a_{n}}{n!}=2 \frac{(2 n-3)!}{(n-2)!n!} . \tag{96}
\end{gather*}
$$

Although there are some cases where the scale $S_{m}$ given by Eq. (93) can be calculated (see below), usually it contains a nonuniversal amplitude. Hence it is interesting to form universal ratios independent of any scale, such as

$$
\begin{equation*}
r_{n}:=\frac{\left\langle S^{n-1}\right\rangle\left\langle S^{n+1}\right\rangle}{\left\langle S^{n}\right\rangle^{2}} . \tag{97}
\end{equation*}
$$

One finds their value at tree level,

$$
\begin{gather*}
r_{n}=1+\frac{2}{2 n-3},  \tag{98}\\
r_{2}=3, \quad r_{3}=\frac{5}{3}, \quad r_{4}=\frac{7}{5}, \quad r_{5}=\frac{9}{7}, \ldots \tag{99}
\end{gather*}
$$

## F. Distribution $\boldsymbol{P}(\boldsymbol{S})$

Let us now perform the inverse Laplace transform of $Z(\lambda)$ to obtain $P(S)$. One must be careful since one expects a form
for $P(S)$ which is not normalizable in the absence of a small- $S$ cutoff; hence we work with $S P(S) /\langle S\rangle$ which is normalized. From Eq. (94) one can write (setting momentarily $S_{m} \rightarrow 1$ )

$$
\begin{equation*}
\langle S\rangle^{-1} \int_{0}^{\infty} d S P(S)\left(e^{\lambda S}-1\right)=\frac{1}{2}(1-\sqrt{1-4 \lambda}) \tag{100}
\end{equation*}
$$

where the $S$ integral converges at small $S$. This is equivalent to

$$
\begin{equation*}
\langle S\rangle^{-1} \int_{0}^{\infty} d S S P(S) e^{\lambda S}=\frac{1}{\sqrt{1-4 \lambda}} \tag{101}
\end{equation*}
$$

as one can check by integrating over $\lambda$ on both sides. Using $\int_{-i \infty}^{i \infty} \frac{d \lambda}{2 \pi i} e^{\left(s-s_{0}\right) \lambda}=\delta\left(s-s_{0}\right)$ one has by inverse Laplace

$$
\begin{align*}
\langle S\rangle^{-1} S P(S) & =\int_{-i \infty}^{i \infty} \frac{d \lambda}{2 \pi i} \frac{e^{-\lambda S}}{\sqrt{1-4 \lambda}} \\
& =\frac{1}{\pi} \int_{1 / 4}^{\infty} d \lambda \frac{1}{\sqrt{4 \lambda-1}} e^{-\lambda S}=\frac{e^{-S / 4}}{2 \sqrt{\pi}} S^{-1 / 2} \tag{102}
\end{align*}
$$

upon deforming the contour around the branch cut. Restoring all factors yields the final result

$$
\begin{equation*}
P(S)=\frac{\langle S\rangle}{2 \sqrt{\pi}} S_{m}^{-1 / 2} S^{-3 / 2} e^{-S /\left(4 S_{m}\right)} \tag{103}
\end{equation*}
$$

As discussed in Sec. III C this is expected to be valid for $S \gtrdot S_{\text {min }}$ in the limit of small $m$, large $S_{m}$. Note that, in general, the exponent $\tau$ can be extracted from the behavior of $Z(\lambda)$ for $\lambda \rightarrow-\infty$, which is dominated by small avalanches and reads, in the scaling regime,

$$
\begin{equation*}
Z(\lambda)=\lambda+\hat{Z}(\lambda) \sim-S_{m}^{\tau-2}|\lambda|^{\tau-1} \tag{104}
\end{equation*}
$$

On the other hand, from its definition (45) and (91) we see that $\lambda+Z(\lambda)$ must converge to $-1 /\langle S\rangle \sim-S_{\min }^{1-\tau} S_{m}^{\tau-2}$ at $\lambda \rightarrow$ $-\infty$; thus the crossover occurs for $\lambda \sim 1 / S_{\min }$. This estimate is consistent with relation (32). For larger values of $-\lambda, Z(\lambda)$ becomes nonuniversal with respect to UV details and is out of reach of the present method.

We can now compute the moments for arbitrary real $n$ $>1 / 2$ by direct integration from Eq. (103). One finds

$$
\begin{equation*}
\frac{\left\langle S^{n}\right\rangle}{\langle S\rangle}=\frac{2^{2 n-2} \Gamma\left(n-\frac{1}{2}\right)}{\sqrt{\pi}} S_{m}^{n-1}, \tag{105}
\end{equation*}
$$

which agrees with the analytic continuations of the moments obtained above.

## G. Discussion of the result: Mean-field theory

Results (93) and (103) for $P(S)$ imply the value of the avalanche distribution exponent,

$$
\begin{equation*}
\tau=3 / 2 \tag{106}
\end{equation*}
$$

It agrees with the results obtained in: (i) a mean-field toy model for dynamic avalanches at depinning [8], (ii) a mean-


FIG. 1. Example of a diagram at MF level, as generated by Eq. (142) at $\alpha=0$. It contains a correction to disorder, i.e., $\Delta^{\prime}\left(0^{+}\right)$at one loop (shaded in gray).
field argument given in the context of a nonequilibrium random-field Ising model [21], and (iii) mean-field calculations developed for sandpile models [14,15]. This form hence seems rather robust as a mean-field result.

Here however, it is derived from first principles using FRG from elastic model (1). It was obtained by resumming all "improved tree" diagrams, i.e., trees made of fully dressed vertices $R$. In the standard bare perturbation theory it includes diagrams as represented in Fig. 1. Hence it can also predict the scale $S_{m}$ given by Eq. (93).

For the elastic model, results (93) and (103) for $P(S)$ are valid for dimensions $d \geqslant 4$. Convergence to this result as $m$ $\rightarrow 0$ requires a one-loop analysis, as discussed below. The only point to discuss is thus the scale $S_{m}$, i.e., the $m$ dependence of $\Delta^{\prime}\left(0^{+}\right)$. For that we use some of the discussion about elastic manifolds for $d>4$ and the associated (Wilson) FRG flow, given in Appendix H of [42], and some specific results for the present model, summarized in Appendix B. The roughness exponent is $\zeta=0$, and the manifold has a finite width. If bare disorder is smooth, it should be sufficiently strong for metastability and (typical) shocks to exist [a cusp in $\Delta(u)$ then develops]. Alternatively one can consider weak rough disorder (i.e., with a cusp) or smooth but with a very short correlation length. In these cases one has large avalanches with

$$
\begin{equation*}
S_{m}=\frac{\left|\Delta^{* \prime}\left(0^{+}\right)\right|}{m^{4}} \tag{107}
\end{equation*}
$$

where for $d>4,\left|\Delta^{* \prime}\left(0^{+}\right)\right|$is a nonuniversal number (see Appendix $B$ ). One can compare this scale with the fluctuations of the center of mass, such that $L^{d} \overline{u(0)^{2}}=\Delta^{*}(0) / m^{4}$, also nonuniversal. In $d=4$ one recovers universality and finds

$$
\begin{equation*}
S_{m} \approx 8 \pi^{2}\left|\hat{\Delta}^{* \prime}\left(0^{+}\right)\right| m^{-4}\left[\ln \left(\frac{m}{m_{0}}\right)\right]^{1-\zeta_{1}} \tag{108}
\end{equation*}
$$

at small $m$, while, for comparison, $L^{d} \overline{u(0)^{2}}=8 \pi^{2} \hat{\Delta}^{*}\left(0^{+}\right)$ $\times m^{-4}\left[\ln \left(\frac{m}{m_{0}}\right)\right]^{1-2 \zeta_{1}}$, where $\zeta_{1}$ and all fixed-point values are given in Appendix B.

In the end it may not be surprising that mean-field theory is obtainable as a summation of trees. However, this had not been done previously within the FRG. It is quite remarkable that the result for the mean-field generating function $\widetilde{Z}_{0}(\lambda)$ in Eq. (89) is identical to the generating function of the number of rooted binary planar trees with $n+1$ leaves [87], also known as the Catalan numbers $C_{n} \equiv b_{n+1}$. The latter also appear for the number of rainbow diagrams in RNA folding, the number of Dyck words of length $2 n$, the number of ways how $n+1$ factors can be completely parenthesized, the num-
ber of monotonic paths along the edges of a grid with $n \times n$ square cells, and many more. It would be interesting to obtain a more microscopic understanding of the relation between the diagrammatic trees and the structure of the (static or dynamic) avalanche processes in $d \geqslant 4$.

This task may be easier to carry out for the dynamics. Intuitively an avalanche starts with a seed and each event may or may not trigger new events. If the latter do not influence much each other, as expected in mean field, the process indeed looks like a tree. These features are captured by the simple model studied in Ref. [8]. There, an avalanche starting at time $t=0$ and lasting until $t=T$, of total size $S$ $=\sum_{t=0}^{\infty} S_{t}$ is a succession of jumps of size $S_{t}$ each being the sum of $n_{t}$ independent events, i.e., $S_{t}=\sum_{i=1}^{n_{t}} s_{i}$. The number of events $n_{t}$ has a Poisson distribution of average $\left\langle n_{t}\right\rangle=\rho S_{t-1}$, i.e., depends only on the jump at time $t-1$ (with $n_{0}=1$ and $S_{t}=0$ for $\left.t \geqslant T\right)$. If the sizes of the independent avalanches are simply $s_{i}=1$ this model is the famous Galton process [88], which describes the evolution of a population of size $S_{t}$ with Poissonian distribution of numbers of offsprings. In this model criticality arises as $\rho$ is increased up to the threshold for an infinite avalanche. Below the threshold the distribution of the avalanche size $S$ is exactly the one obtained here by summing the trees. The $S^{-3 / 2}$ power law distribution may then be understood as the distribution of return time to the origin of a random walk. Interestingly, there is also a $d=0$ model which captures this mean-field physics and its connection to return time of a random walk. It is the celebrated ABBM model [89] for domain wall motion, represented as a particle in a Brownian random force landscape. We recently computed [90] the avalanche statistics and the renormalized FRG force correlators for this model, and the similarities with the mean-field results obtained here are striking. These analogies involve mostly the zero momentum, $q=0$ structure, but in Appendix F we present a spatial generalization of our tree summation, which may help us to understand the relation between trees and avalanches in mean field.

## VI. LOOP CORRECTIONS

## A. General method

Until now the force correlations $\hat{C}^{(n)}$, and from them the $K^{(n)}$ and the avalanche size moments $\left\langle S^{n}\right\rangle /\langle S\rangle$, were computed to improved tree level, i.e., setting $S^{(n)}=0$ for $n \geqslant 3$ in effective action (49). After Legendre transform this gives each $\hat{C}^{(n)}$ to the lowest order in an expansion in powers of $\Delta$ (or $R$ ), and later, using the fixed-point values for $\Delta$ to lowest order in $\epsilon$.

It is useful to describe a systematic procedure to compute these quantities to higher order in the expansion in powers of $R$ in the statics (or $\Delta$ in the dynamics). For each specific calculation, an intuitive diagrammatic representation is given below.
(i) One first computes all the functions $S^{(n)}, n>3$, in the effective action in an expansion in powers of $R$ [91]. They are given by the sum of all one-particle irreducible (n-replica) diagrams with $R$ as vertices, i.e., the sum of all loops. For instance, to leading order to which we restrict below one finds $S^{(n)} \sim R^{n}$ with a unique one-loop integral $I_{n}$,

$$
\begin{gather*}
I_{n}=\int_{k} \frac{1}{\left(k^{2}+m^{2}\right)^{n}}=m^{d-2 n} \widetilde{I}_{n},  \tag{109}\\
\widetilde{I}_{n}=\int_{k} \frac{1}{\left(k^{2}+1\right)^{n}} \tag{110}
\end{gather*}
$$

for $d<2 n$, and only the replica combinatorics is nontrivial, e.g., for $n=3,4$

$$
\begin{align*}
S_{123}^{(3)}= & 3 I_{3} \operatorname{sym}\left[\left(\mathrm{R}_{12}^{\prime \prime}\right)^{2} \mathrm{R}_{13}^{\prime \prime}-\frac{1}{3} \mathrm{R}_{12}^{\prime \prime} \mathrm{R}_{23}^{\prime \prime} \mathrm{R}_{31}^{\prime \prime}\right],  \tag{111}\\
S_{1234}^{(4)}= & 3 I_{4} \operatorname{sym}\left[2\left(\mathrm{R}_{12}^{\prime \prime}\right)^{2} \mathrm{R}_{13}^{\prime \prime}\left(2 \mathrm{R}_{14}^{\prime \prime}+\mathrm{R}_{24}^{\prime \prime}\right)\right. \\
& \left.-4 \mathrm{R}_{12}^{\prime \prime} \mathrm{R}_{13}^{\prime \prime} \mathrm{R}_{23}^{\prime \prime} \mathrm{R}_{14}^{\prime \prime}+\mathrm{R}_{12}^{\prime \prime} \mathrm{R}_{23}^{\prime \prime} \mathrm{R}_{34}^{\prime \prime} \mathrm{R}_{14}^{\prime \prime}\right], \tag{112}
\end{align*}
$$

where here $R^{\prime \prime}(w):=R^{\prime \prime}(w)-R^{\prime \prime}(0), R_{a b}^{\prime \prime}=R^{\prime \prime}\left(w_{a}-w_{b}\right), S_{123}^{(3)}$ $=S^{(3)}\left(w_{1}, w_{2}, w_{3}\right)$, and so on. The formula for general $n$ was given in [35] and takes the form for the $n$-replica term,

$$
\begin{gather*}
S_{a_{1}, \ldots, a_{n}}^{(n)}=\frac{(n-1)!}{2} I_{n}\left[\operatorname{tr}\left(W^{n}\right)-\sum_{a} R_{a a_{1}}^{\prime \prime} \cdots R_{a a_{n}}^{\prime \prime}\right] \\
W_{a b}=\delta_{a b} \sum_{a_{1}} R_{a a_{1}}^{\prime \prime}-R_{a b}^{\prime \prime} \tag{113}
\end{gather*}
$$

i.e., it is a trace over replica indices, with the $n+1$ replica term is subtracted. This formula yields Eqs. (111) and (112); the explicit form for $n=5$ was given in [42].
(ii) One then performs the Legendre transform (49) from $S^{(n)}$ to $\hat{S}^{(n)}$. From them one can then obtain $\hat{C}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$ $=(-1)^{n} \partial_{w_{1}} \cdots \partial_{w_{n}} \hat{S}^{(n)}\left(w_{1}, \ldots, w_{n}\right)$. Legendre transformation simply means that each $\hat{S}^{(n)}$ is a sum of all ( $n$-replica) tree diagrams which can be drawn using $R$ or any of the $S^{(n)}, n$ $>3$, at the vertices of the tree. Since when forming a tree the number of replica can only increase, there are simple exact formulas for the lowest moments,

$$
\begin{align*}
\hat{S}^{(3)}\left(w_{123}\right) & =S_{0}^{(3)}\left(w_{123}\right)+S^{(3)}\left(w_{123}\right)  \tag{114}\\
\hat{S}^{(4)}\left(w_{1234}\right)= & S_{0}^{(4)}\left(w_{1234}\right)+S^{(4)}\left(w_{1234}\right) \\
& +\frac{12}{m^{2}} \operatorname{sym} R^{\prime}\left(w_{14}\right) \partial_{w_{1}} S^{(3)}\left(w_{123}\right), \tag{115}
\end{align*}
$$

as can be checked by explicit Legendre transform [62,63]. Here the $S_{0}^{(n)} \sim R^{n-1}$ are the result of the improved tree approximation, i.e., the trees made of only $R$ at the vertices, as described above, giving Eq. (59) for the two lowest moments. Note that $\hat{S}^{(3)}$ contains a single $I_{3}$ loop integral while $\hat{S}^{(4)}$ contains both $I_{3}$ and $I_{4}$ and so on. Since everywhere we consider uniform $w$, the trees are at zero external momenta (hence the factor $1 / \mathrm{m}^{2}$ ), and loop momenta only flow inside each $S^{(n)}$ vertex. In principle using Eq. (115) and inserting each $S^{(n)}$ to the required order in $R$ allow us to compute the $\hat{C}^{(n)}$ to any desired order in $\Delta$ and in $\epsilon$.

Here we only want $\hat{S}^{(n)}$ to order $R^{n}$, hence we do not need all trees made of $S^{(n)}$ and $R$ but only those trees containing
one $S^{(p)}$ and one or more $R$, i.e., it with schematically the form

$$
\begin{equation*}
\hat{S}^{(n)}=\hat{S}_{0}^{(n)}+S^{(n)}+\sum_{p=1}^{n-3} \operatorname{sym} S^{(n-p)} R^{p} \tag{116}
\end{equation*}
$$

We first perform an explicit calculation of third and fourth moments to one loop before attempting the resummation. Note that to one-loop order, an equivalent procedure is to perform perturbation theory in the bare disorder $\Delta_{0}$, in which case an additional term $I_{2}$ will appear in each $\hat{C}^{(n)}$, with trees based on a single $\Delta_{0}$ loop. Inserting the one-loop expression of $\Delta_{0}$ as a function of $\Delta$ this term disappears and one recovers the same result as with the above method.

We need below the momentum integrals $I_{n}$. One has $\tilde{I}_{n}$ $\sim \int_{k} \int_{0}^{\infty} d s \exp \left(-s-s k^{2}\right) s^{n-1} /(n-1)!$; hence,

$$
\begin{equation*}
\frac{\tilde{I}_{n}}{\epsilon \widetilde{I}_{2}}=\frac{\Gamma(n-d / 2)}{2 \Gamma(n) \Gamma(3-d / 2)} \xrightarrow{d \rightarrow 4} \frac{\Gamma(n-2)}{2 \Gamma(n)}=\frac{1}{2(n-1)(n-2)}, \tag{117}
\end{equation*}
$$

where $\epsilon \widetilde{I}_{2}$ is discussed in Appendix B. Note that all these momentum integrals are IR finite for $d<6$. This will remain true to any number of loop as all diagrams entering in $\hat{C}^{(n)}$ are superficially IR convergent by power counting for $d \leq 4$. If one uses perturbation theory in the bare disorder, they will however contain diverging subdiagrams, starting at two-loop order. These divergences should be removed by the counterterms for the disorder, i.e., the replacement of $\Delta_{0}$ as a function of $\Delta$. Therefore, corrections to $\tau$ should come from finite diagrams and appear when summing all diagrams. On the other hand, for small $n$, the moments are interesting since they can be compared to the numerics.

## B. Third cumulant

From the above expressions [Eqs. (111) and (115)], we obtain

$$
\begin{align*}
-\hat{C}_{1 \text {-loop }}^{(3)}\left(w_{1}, w_{2}, w_{3}\right)= & 6 I_{3} \operatorname{sym}\left\{\Delta^{\prime}\left(w_{12}\right)^{2} \Delta^{\prime}\left(w_{13}\right)\right. \\
& +\left[\Delta\left(w_{12}\right)-\Delta(0)\right]\left[\Delta^{\prime}\left(w_{13}\right) \Delta^{\prime \prime}\left(w_{12}\right)\right. \\
& \left.\left.+\Delta^{\prime}\left(w_{12}\right) \Delta^{\prime \prime}\left(w_{13}\right)+\Delta^{\prime}\left(w_{23}\right) \Delta^{\prime \prime}\left(w_{13}\right)\right]\right\} \tag{118}
\end{align*}
$$

It is useful to indicate the graphical representation of each term using the conventions described in Sec. V C (suppressing for clarity below all propagator and momentum loop factors). There are three nonvanishing one-loop contributions,



They reproduce above result (118) after symmetrization. One additional diagram,

vanishes after symmetrization in the statics.
Note that the total result for $\hat{C}^{(3)}$ should be continuous and unambiguous. Note that $\hat{C}^{(3)}(0,0,0)=0$ while this is not the case for the dynamics (it gives the third cumulant of the critical force).

From there we compute the Kolmogorov cumulant $K^{(3)}(w)$. One finds for each piece


$$
=12 I_{3}[\Delta(0)-\Delta(w)] \Delta^{\prime}(w)\left[\Delta^{\prime \prime}(w)+\Delta^{\prime \prime}(0)\right]
$$

$$
\begin{equation*}
\approx-24 I_{3} \Delta^{\prime}\left(0^{+}\right)^{2} \Delta^{\prime \prime}(0) w \tag{123}
\end{equation*}
$$



$$
\begin{equation*}
=12 I_{3}[\Delta(w)-\Delta(0)] \Delta^{\prime}(w)\left[\Delta^{\prime \prime}(0)-\Delta^{\prime \prime}(w)\right]=O\left(w^{2}\right), \tag{124}
\end{equation*}
$$

and

$$
\begin{align*}
\mathcal{K} & =12 I_{3} \Delta^{\prime}(w)\left[\Delta^{\prime}\left(0^{+}\right)^{2}-\Delta^{\prime}(w)^{2}\right] \\
& =-24 I_{3} \Delta^{\prime}\left(0^{+}\right)^{2} \Delta^{\prime \prime}(0) w+O\left(w^{2}\right) \tag{125}
\end{align*}
$$

Together, they give


The final result for the small-w behavior of $K^{(3)}(w)$ is

$$
\begin{align*}
K^{(3)}(w) & =K_{\text {tree }}^{(3)}+K_{1-\text { loop }}^{(3)}  \tag{127}\\
& =\frac{12}{m^{2}} \Delta^{\prime}\left(0^{+}\right)^{2}\left[1-4 m^{2} I_{3} \Delta^{\prime \prime}\left(0^{+}\right)\right] w+\cdots . \tag{128}
\end{align*}
$$

Hence, using Eq. (42) we obtain the third moment of avalanche sizes,

$$
\begin{align*}
\frac{\left\langle S^{3}\right\rangle}{\langle S\rangle} & =\frac{12}{m^{8}} \Delta^{\prime}\left(0^{+}\right)^{2}\left[1-4 m^{2} I_{3} \Delta^{\prime \prime}\left(0^{+}\right)\right]+O\left(\Delta^{4}\right) \\
& =12 S_{m}^{2}\left(1-4 \widetilde{\Delta}^{\prime \prime}\left(0^{+}\right) \frac{\tilde{I}_{3}}{\epsilon \widetilde{I}_{2}}\right) \tag{129}
\end{align*}
$$

We now use that

$$
\begin{equation*}
\frac{\widetilde{I}_{3}}{\epsilon \widetilde{I}_{2}}=\frac{1}{4}, \quad \widetilde{\Delta}^{\prime \prime}\left(0^{+}\right)=\frac{\epsilon-\zeta}{3}+O\left(\epsilon^{2}\right) \tag{130}
\end{equation*}
$$

where the second equality is valid at the fixed point (see Appendix B). Using the exact relation $\left\langle S^{2}\right\rangle /\langle S\rangle$ $=-2 \Delta^{\prime}\left(0^{+}\right) / m^{4}=2 S_{m}$, we thus obtain the universal ratio to one loop,

$$
\begin{equation*}
r_{1-\text { loop }}^{(2)}=\frac{\langle S\rangle\left\langle S^{3}\right\rangle}{\left\langle S^{2}\right\rangle^{2}}=3\left[1-\widetilde{\Delta}^{\prime \prime}(0)\right]=3\left[1-\frac{1}{3}(\epsilon-\zeta)\right] . \tag{131}
\end{equation*}
$$

## C. Fourth and higher cumulants

Similar calculations, either directly from Eq. (113) or by calculating the sum of all diagrams at one-loop order, we find the corrections for the fourth Kolmogorov cumulant,

$$
\begin{aligned}
K_{1-\mathrm{loop}}^{(4)}(w)= & 12\left(-6 \Delta^{\prime}(w)^{2}\left\{\Delta^{\prime}\left(0^{+}\right)^{2}+4[\Delta(0)-\Delta(w)] \Delta^{\prime \prime}(w)\right\}-\Delta^{\prime}\left(0^{+}\right)^{4}+2[\Delta(0)-\Delta(w)]^{2} \Delta^{\prime \prime}(w)^{2}+7 \Delta^{\prime}(w)^{4}\right) I_{4} \\
& +48\left(\left\{16[\Delta(w)-\Delta(0)] \Delta^{\prime \prime}(w)-3 \Delta^{\prime}\left(0^{+}\right)^{2}\right\} \Delta^{\prime}(w)^{2}+4[\Delta(0)-\Delta(w)]^{2} \Delta^{(3)}(w) \Delta^{\prime}(w)\right. \\
& \left.+3 \Delta^{\prime}(w)^{4}+2[\Delta(0)-\Delta(w)] \Delta^{\prime \prime}(w)\left\{\Delta^{\prime}\left(0^{+}\right)^{2}+2[\Delta(0)-\Delta(w)] \Delta^{\prime \prime}(w)\right\}\right) I_{3} .
\end{aligned}
$$

In the limit of small $w$,

$$
\begin{equation*}
K_{1-\text { loop }}^{(4)}(w)=\left(480 I_{4}+960 \frac{I_{3}}{m^{2}}\right) \Delta^{\prime}\left(0^{+}\right)^{3} \Delta^{\prime \prime}\left(0^{+}\right)|w|+O\left(w^{2}\right) \tag{132}
\end{equation*}
$$

Adding the tree-level part and using Eq. (42) we obtain the fourth moment of avalanche sizes,

$$
\begin{align*}
\frac{\left\langle S^{4}\right\rangle}{\langle S\rangle} & =\frac{120}{m^{12}}\left|\Delta^{\prime}\left(0^{+}\right)\right|^{3}\left[1-4\left(m^{4} I_{4}+2 m^{2} I_{3}\right) \Delta^{\prime \prime}\left(0^{+}\right)\right] \\
& =120 S_{m}^{3}\left[1-4 \frac{\widetilde{I}_{4}+2 \widetilde{I}_{3}}{\epsilon I_{2}} \widetilde{\Delta}^{\prime \prime}\left(0^{+}\right)\right] \tag{133}
\end{align*}
$$

It yields for the universal ratio,

$$
\begin{align*}
r_{1-\mathrm{loop}}^{(3)}=\frac{\left\langle S^{4}\right\rangle\left\langle S^{2}\right\rangle}{\left\langle S^{3}\right\rangle^{2}} & =\frac{5}{3}\left[1-4 \frac{\widetilde{I}_{4}}{\epsilon \widetilde{I}_{2}} \Delta^{\prime \prime}(0)\right] \\
& =\frac{5}{3}\left[1-\widetilde{\Delta}^{\prime \prime}(0)\left(1-\frac{d}{6}\right)\right] \\
& =\frac{5}{3}\left[1-\frac{(\epsilon-\zeta)(6-d)}{18}\right] \tag{134}
\end{align*}
$$

with MATHEMATICA we have computed $\hat{C}^{(n)}$ and $K^{(n)}$ up to $n=6$. The expressions are tedious. Let us just note that one finds the limits at zero argument,

$$
\begin{equation*}
\hat{C}^{(2 p)}(0, \ldots, 0)=-(2 p-1)!I_{2 p} \Delta^{\prime}\left(0^{+}\right)^{2 p} \tag{135}
\end{equation*}
$$

for even $n \geqslant 4$ and 0 for odd cumulants. This allows us to compute the probability distribution of the center of mass fluctuations. The calculation and results are given in Appendix G. It bears some similarities with the one for the distribution of the critical force at depinning in [53].

Let us give here only the part of the Kolmogorov cumulants proportional to $I_{n}$. They read [denoting here $\Delta(v)-\Delta(0)$ by $\Delta(v)]$

$$
\begin{align*}
\hat{K}^{(5)}(v)= & -240 I_{5} \Delta^{\prime}(v)\left\{-\Delta^{\prime}\left(0^{+}\right)^{4}+3 \Delta^{\prime}(v)^{4}+4 \Delta(v)^{2} \Delta^{\prime \prime}(v)^{2}\right. \\
& \left.-2 \Delta^{\prime}(v)^{2}\left[\Delta^{\prime}\left(0^{+}\right)^{2}-8 \Delta(v) \Delta^{\prime \prime}(v)\right]\right\}, \tag{136}
\end{align*}
$$

$$
\begin{align*}
\hat{K}^{(6)}(v)= & 240 I_{6}\left\{-\Delta^{\prime}\left(0^{+}\right)^{6}+31 \Delta^{\prime}(v)^{6}+4 \Delta(v)^{3} \Delta^{\prime \prime}(v)^{3}\right. \\
& -15 \Delta^{\prime}(v)^{4}\left[\Delta^{\prime}\left(0^{+}\right)^{2}-16 \Delta(v) \Delta^{\prime \prime}(v)\right] \\
& \left.-15 \Delta^{\prime}(v)^{2}\left[\Delta^{\prime}\left(0^{+}\right)^{4}-8 \Delta(v)^{2} \Delta^{\prime \prime}(v)^{2}\right]\right\} \tag{137}
\end{align*}
$$

Expanding at small $w>0$ one finds, for the part proportional to $I_{n}$,

$$
\begin{equation*}
s \hat{K}^{(n)}(w)=-A_{n} I_{n}\left[-\Delta^{\prime}\left(0^{+}\right)\right]^{n-1} \Delta^{\prime \prime}(0) w+O\left(w^{2}\right) \tag{138}
\end{equation*}
$$

with coefficients

$$
\begin{align*}
& A_{2}=6, \quad A_{3}=48, \quad A_{4}=480 \\
& A_{5}=5760, \quad A_{6}=80640, \ldots \tag{139}
\end{align*}
$$

Note that $A_{n+1} / A_{n}=2 n+4$. This suggests that

$$
\begin{equation*}
A_{n}=2^{n-2}(n+1)! \tag{140}
\end{equation*}
$$

which is supported by a heuristic argument in Appendix C.

## VII. RESUMMATION OF ALL DIAGRAMS WITH A SINGLE LOOP: SELF-CONSISTENT EQUATION

## A. Self-consistent equation

We can now perform a resummation of all cumulants to one-loop order. In view of the results of the previous sections we can absorb all factors of $m$ and $\widetilde{\Delta}^{\prime}\left(0^{+}\right)$by using the rescaled generating function $\widetilde{Z}(\lambda)$ defined in Eqs. (91) and (93). One has

$$
\begin{equation*}
\widetilde{Z}(\lambda)=\sum_{n=1}^{\infty} \frac{c_{n}}{n!} \lambda^{n} \tag{141}
\end{equation*}
$$

where the $c_{n}$ is extracted from the Kolmogorov cumulants at small $w$ and at tree level $c_{n}=a_{n}$. In the previous section we have obtained the contribution to $c_{n}$ proportional to $\tilde{I}_{n}$, involving the coefficients $A_{n}$. Remarkably, the remaining corrections (i.e., the terms in $c_{n}$ proportional to $\tilde{I}_{p}$ with $3 \leqslant p$ $<n)$ can be generated automatically by taking advantage of the tree structure. This is illustrated graphically as follows:


Here, the shaded blob represents $\widetilde{Z}(\lambda)$. It must contain all sets of trees made with $R$ vertices joined together at their basis either by a $R$ vertex or by a $S^{(n)}$ vertex with $n>3$, i.e., a loop, in the spirit of Eq. (116). Expressed as an equation it reads

$$
\begin{equation*}
\widetilde{Z}(\lambda)=\lambda+\widetilde{Z}(\lambda)^{2}+\alpha \sum_{n \geqslant 3}(n+1) 2^{n-2} i_{n} \widetilde{Z}(\lambda)^{n} \tag{142}
\end{equation*}
$$

where we have defined $i_{n}:=\widetilde{I}_{n} /\left(\epsilon \widetilde{I}_{2}\right)$ and substituted the factor $(n+1) 2^{n-2}=A_{n} / n!$. Here and below we denote

$$
\begin{equation*}
\alpha:=-\epsilon \widetilde{I}_{2} m^{-\epsilon} \Delta^{\prime \prime}\left(0^{+}\right)=-\widetilde{\Delta}^{\prime \prime}\left(0^{+}\right) \tag{143}
\end{equation*}
$$

which is of order $\epsilon$. It was not strictly necessary to make Eq. (142) self-consistent; one could with the same accuracy of $O(\epsilon)$ replace all terms $\widetilde{Z}^{n}$ with $n \geqslant 3$ by $\widetilde{Z}_{0}^{n}$. By making it self-consistent we include more loops. It is however more convenient in view of the form at tree level. It also has a nice interpretation in terms of the generating function of trees with higher- $n$ branching: these occur at a small rate, which is $O(\alpha)$, and $n$ dependent.

## B. Summation of the series and final equation

From the above self-consistent equation one can compute iteratively $\tilde{Z}(\lambda)$ with the result

$$
\begin{align*}
\tilde{Z}(\lambda)= & \lambda+\lambda^{2}+\left[8 \alpha i_{3}+2\right] \lambda^{3}+\frac{1}{24}\left[960 \alpha i_{3}+480 \alpha i_{4}+120\right] \lambda^{4} \\
& +\frac{1}{120}\left[23040 \alpha^{2} i_{3}^{2}+20160 \alpha i_{3}+14400 \alpha i_{4}+5760 \alpha i_{5}\right. \\
& +1680] \lambda^{5}+\frac{1}{720}\left[1290240 i_{3}^{2} \alpha^{2}+806400 i_{3} i_{4} \alpha^{2}\right. \\
& +483840 i_{3} \alpha+403200 i_{4} \alpha+241920 i_{5} \alpha \\
& \left.+80640 i_{6} \alpha+30240\right] \lambda^{6}+O\left(\lambda^{7}\right) . \tag{144}
\end{align*}
$$

One recognizes that the results from the third and fourth cumulants of the previous sections are correctly reproduced. Note that the terms with higher powers of $i_{n}$ contain more than one (nonoverlapping) loops and higher powers of $\alpha$ $=O(\epsilon)$. The terms of order $\alpha$ are all correctly generated.

We now want to compute the infinite sum in Eq. (142). For this we use the following representation of the $i_{n}$ as in the derivation of Eq. (117),

$$
\begin{equation*}
i_{n}=\mathcal{N} \frac{1}{(n-1)!} \int_{0}^{\infty} d s e^{-s} s^{-1-d / 2} s^{n} \tag{145}
\end{equation*}
$$

with $\mathcal{N}=1 /\left[2 \Gamma\left(3-\frac{d}{2}\right)\right]$. One finds

$$
\begin{equation*}
\sum_{n \geqslant 3}(n+1) 2^{n-2} i_{n} z^{n}=\mathcal{N} \int_{0}^{\infty} d s e^{-s} s^{-d / 2-1} \sum_{n>2} \frac{1}{4} \frac{(n+1)}{(n-1)!}(2 z s)^{n} \tag{146}
\end{equation*}
$$

Let us define

$$
\begin{equation*}
\psi(z):=\sum_{n \geqslant 3} \frac{1}{4} \frac{(n+1)}{(n-1)!}(2 z)^{n}=z\left(e^{2 z} z-3 z+e^{2 z}-1\right) . \tag{147}
\end{equation*}
$$

One can write

$$
\begin{align*}
\sum_{n \geqslant 3} & (n+1) 2^{n-2} i_{n} z^{n} \\
& =\mathcal{N} \int_{0}^{\infty} d s e^{-s} s^{-d / 2-1} \psi(s z) \\
& =\frac{z\left\{[2-(d+2) z](1-2 z)^{(d-4) / 2}+3 d z-6 z-2\right\}}{(d-2)(d-4)} \\
& \stackrel{d \rightarrow 4}{\longrightarrow} \frac{1}{2} z[2 z+(1-3 z) \log (1-2 z)] \\
& =2 z^{3}+\frac{5 z^{4}}{3}+2 z^{5}+\frac{14 z^{6}}{5}+\frac{64 z^{7}}{15}+\frac{48 z^{8}}{7}+O\left(z^{9}\right) \tag{148}
\end{align*}
$$

The last line is also obtained by using $i_{n}=\frac{1}{2(n-1)(n-2)}$ from Eq. (117). One can thus rewrite the self-consistent equation $[\mathrm{Eq}$. (142)] as

$$
\begin{align*}
\tilde{Z}(\lambda)= & \tilde{Z} \\
= & \lambda+\widetilde{Z}^{2} \\
& +\alpha \frac{\tilde{Z}\left\{[2-(d+2) \tilde{Z}](1-2 \tilde{Z})^{(d-4) / 2}+3 d \tilde{Z}-6 \tilde{Z}-2\right\}}{(d-2)(d-4)} . \tag{149}
\end{align*}
$$

Close to $d=4$ it becomes

$$
\begin{equation*}
\widetilde{Z}=\lambda+\widetilde{Z}^{2}+\alpha \frac{1}{2} \tilde{Z}[2 \tilde{Z}+(1-3 \tilde{Z}) \log (1-2 \tilde{Z})] \tag{150}
\end{equation*}
$$

This equation has to be inverted in order to get both the exponent $\tau$ and the tail of the distribution $P(S)$ for large $S$. Note that from Eq. (142) it can also be written as

$$
\begin{align*}
\tilde{Z}(\lambda)=: & \tilde{Z}= \\
& \lambda+\widetilde{Z}^{2}+\frac{\alpha}{\epsilon \tilde{I}_{2}} \int_{k}\left[\frac{\widetilde{Z}^{2}}{\left(k^{2}+1-2 \tilde{Z}\right)^{2}}+\frac{\tilde{Z}}{\left(k^{2}+1-2 \tilde{Z}\right)}\right.  \tag{151}\\
& \left.-\frac{\tilde{Z}}{k^{2}+1}-3 \frac{\widetilde{Z}^{2}}{\left(k^{2}+1\right)^{2}}\right] .
\end{align*}
$$

As shown in Appendix C, this formula has a simple graphical interpretation. The last two terms can be thought of as counterterms which fix the coefficients of $\widetilde{Z}$ and $\widetilde{Z}^{2}$ according to the choice we implemented here. (They would be absent if we wrote down an expansion in the bare disorder.)

## VIII. ONE-LOOP RESULTS FOR THE AVALANCHE-SIZE DISTRIBUTION

## A. Avalanche exponent $\tau$

Formula (150) contains all we need in order to retrieve the avalanche size distribution. We recall that

$$
\begin{equation*}
S_{m}^{-1} \widetilde{Z}\left(\lambda S_{m}\right)=\langle S\rangle^{-1}\left(\left\langle e^{\lambda S}\right\rangle-1\right) \tag{152}
\end{equation*}
$$

from Eq. (91), where $S_{m}$ is given by Eq. (93). Note however that Eq. (150) contains only information about sizes of order $S_{m}$, from $S \ll S_{m}$ to $S \gtrdot S_{m}$, but in all cases much larger than the microscopic cutoff, i.e., $S \gg S_{\min }$, which has not been included (and was not needed) in the above analysis.

Let us start with the exponent $\tau$. It can be extracted as follows. There is a critical value $p_{c}<1$ such that all moments $\int d S P(S) S^{p}$ diverge for $p \leqslant p_{c}$. Of course these do not strictly diverge, as they should be in the end cutoff by $S_{\min }$, but as just discussed we can forget here to extract the $\tau$ exponent. This exponent is defined from the behavior $P(S) \sim S^{-\tau}$ for $S_{\text {min }} \ll S \ll S_{m}$. Hence we can identify

$$
\begin{equation*}
\tau=1+p_{c} \tag{153}
\end{equation*}
$$

We start from the identity

$$
\begin{align*}
\left\langle S^{p}\right\rangle & =\frac{1}{\Gamma(-p)} \int_{0}^{\infty} d \lambda \lambda^{-1-p}\left\langle e^{-\lambda S}-1\right\rangle \\
& =\frac{\langle S\rangle}{S_{m}} \frac{1}{\Gamma(-p)} \int_{0}^{\infty} d \lambda \lambda^{-1-p} \widetilde{Z}\left(-S_{m} \lambda\right), \tag{154}
\end{align*}
$$

where in the second line we have used Eq. (152). Since $\tilde{Z} \sim \lambda$ at small $\lambda$ the integral converges at the lower bound. However, it may diverge at the upper bound. Hence $p_{c}$ is obtained from the power-law tail of $\tilde{Z}(\lambda)$ for $\lambda \rightarrow-\infty$. The boundary is for $\widetilde{Z}(-\lambda) \sim-\lambda^{p_{c}}$, where we have indicated the sign of $\widetilde{Z}$. ( $\tilde{Z}$ is negative, due to the dominance of the term -1 in the definition of $\tilde{Z}$ for large negative $\lambda$ ). This is consistent with asymptotic behavior (104) and the discussion there.

From Eq. (150) we obtain, for large negative $\widetilde{Z} \equiv \widetilde{Z}(\lambda)$,

$$
\begin{equation*}
-\lambda \approx \widetilde{Z}^{2}\left[1+\alpha-\frac{3}{2} \alpha \log (-2 \widetilde{Z})\right] \approx(-\widetilde{Z})^{2-(3 / 2) \alpha}+O\left(\alpha^{2}\right) \tag{155}
\end{equation*}
$$

This gives $-\tilde{Z} \approx(-\lambda)^{1 /[2-(3 / 2) \alpha]}=(-\lambda)^{p_{c}}$, from which we identify

$$
\begin{align*}
\tau=1+\frac{1}{2-\frac{3}{2} \alpha} & \approx \frac{3}{2}+\frac{3}{8} \alpha  \tag{156}\\
& =\frac{3}{2}-\frac{3}{8} \widetilde{\Delta}^{\prime \prime}(0) . \tag{157}
\end{align*}
$$

Using relation (B12) of Appendix B, valid for all classes of disorder, we obtain our main result for the avalanche exponent,

$$
\begin{equation*}
\tau=\frac{3}{2}-\frac{\epsilon-\zeta}{8}+O\left(\epsilon^{2}\right) \tag{158}
\end{equation*}
$$

Let us now discuss the significance of this result. First we note that this formula agrees to first order in $\epsilon$ with the conjecture

$$
\begin{equation*}
\tau=\tau_{\zeta}:=2-\frac{2}{d+\zeta} \tag{159}
\end{equation*}
$$

As mentioned in the Introduction, this conjecture was put forward in the study of interface depinning in Ref. [27] in the absence of a mass term, i.e., in the ensemble of fixed applied force $f$. It is based on the assumption that, as the force $f$ is increased toward the threshold $f_{c}$, the mean number of avalanches per interval of force increment is not singular. Translated into the present setting it means that

$$
\begin{equation*}
\frac{d N_{w}}{m^{2} d w}=\frac{\rho_{0}}{m^{2}} \tag{160}
\end{equation*}
$$

remains finite as $m \rightarrow 0$. (We use the notations of Sec. III C where $N_{w}$ is the number of shocks and $\rho_{0}$ is the shock number density.) Note however that in the present setting, although the increase in the force is $m^{2} d w$, it does not bring the system closer to criticality, which is achieved by decreasing the mass. This is why we had at the start two independent exponents $\rho$ and $\tau$. The statistical tilt symmetry (STS) symmetry provides first relation (23) while the conjecture of the finiteness of Eq. (160) is equivalent, using Eq. (29), to the conjecture $\rho=2$ (at least for $\tau>1$ ). Using Eq. (29) again, this
implies the value $\tau=\tau_{\zeta}$. Finally note that a similar conjecture was put forward in the context of sandpiles $[18,19]$, and if one admits the connection to the depinning of a periodic interface $[24,75]$ it is equivalent to the conjecture $\tau=\tau_{\zeta=0}$ for the RP class.

As discussed in Ref. [67], we expect the result for $\tau$ for statics (shocks) and for depinning (avalanches) to coincide to one loop. Hence our result indicates that both indeed coincide with the conjecture $\tau=\tau_{\zeta}$ to one loop. What happens beyond one loop accuracy remains to be elucidated. We know that statics and depinning differ at two loop, e.g., for the $\zeta$ exponent. There are various possibilities for either the statics or the depinning to agree or disagree with the conjecture at two loops and for $d=2,3$ : since $\zeta_{\text {dep }}>\zeta_{\text {stat }}$, one possibility is already excluded, which both conjectures are true and that the exponent $\tau$ is the same in both problems. Some claims that static and dynamic avalanches belong to the same universality classes were indeed put forward in [22] for the RF Ising model. Finally note that our result can also be written as

$$
\begin{equation*}
\rho=2+O\left(\epsilon^{2}\right) \tag{161}
\end{equation*}
$$

and a remaining challenge is to compute the higher orders. Until now we have failed to find, within the field theory, a symmetry or a mechanism which would imply $\rho=2$ to all orders.

Note that exact results for $d=0$ depinning [90], i.e., of a particle, yield $\tau=0$ for the so-called Gumbel class which has $\zeta=2$, hence in disagreement with the conjecture. The result for the other classes (Weibul and Frechet) has also $\tau=0$ but $\zeta \neq 2$, hence also violating the conjecture in its present form. For the statics, exact result for the $d=0$ limit of the RF class, the so-called toy or Sinai Brownian energy landscape, has $\tau=1 / 2$ and $\zeta=4 / 3$ which again satisfy the conjecture [62].

To summarize, for the cases of interest, our result is, with $\epsilon=4-d$ and up to $O\left(\epsilon^{2}\right)$ corrections,

$$
\tau=\frac{3}{2}-\left\{\begin{array}{cc}
\frac{1}{12} \epsilon, & \mathrm{RF}  \tag{162}\\
0.0989627 \epsilon, & \mathrm{RB} \\
\frac{1}{8} \epsilon, & \mathrm{CDW}
\end{array}\right.
$$

where we have used the one-loop values for $\zeta$ in the statics, recalled in Appendix B. To $O(\epsilon)$ these are the same as for depinning, except that in that case the RB class does not exist since two-loop corrections make it flow to RF.

It is useful to quote the expected result if the conjecture holds. In $d=3$,

$$
\tau=\left\{\begin{array}{cc}
\frac{7}{5}, & \mathrm{RF}  \tag{163}\\
1.37783 \pm 0.00005, & \mathrm{RB} \\
\frac{4}{3}, & \mathrm{CDW}
\end{array}\right.
$$

In $d=2$,

$$
\tau=\left\{\begin{array}{cc}
\frac{5}{4}, & \mathrm{RF}  \tag{164}\\
1.17987 \pm 0.00037, & \mathrm{RB} \\
1, & \mathrm{CDW}
\end{array}\right.
$$

In $d=1$,

$$
\tau= \begin{cases}1, & \mathrm{RF}  \tag{165}\\ \frac{4}{5}, & \mathrm{RB} \\ 0, & \mathrm{CDW}\end{cases}
$$

For the RB class we have used the two-loop result $\zeta=0.208298042 \epsilon+0.0068582 \epsilon^{2}+O\left(\epsilon^{3}\right)$ and constructed three different Padé approximants, using the constraint that $\zeta_{d=1}=2 / 3$. This gives $\zeta_{d=3}^{\mathrm{RB}}=0.21454 \pm 0.00028$ and $\zeta_{d=2}^{\mathrm{RB}}$ $=0.43864 \pm 0.0011$. The error bars denote the $1 \sigma$ spread of the three Padé approximants. For RF we use $\zeta=(4-d) / 3$ and for $\operatorname{RP} \zeta=0$, hence there is no uncertainty. These are the
values for the statics. For depinning, if the conjecture holds, the values for the RP (CDW) class are the same as given here. For nonperiodic interfaces $\zeta=\frac{1}{3} \epsilon(1+0.14331 \epsilon)+O\left(\epsilon^{3}\right)$ and one gets $\tau_{d=3}=1.409 \pm 0.001, \tau_{d=2}=1.31 \pm 0.01$, and $\tau_{d=1}=1.2 \pm 0.1$ (in the latter inserting the numerically measured value for $\zeta=1.25$ yields instead $\tau_{d=1}=1.11$ ).

## B. Distribution of avalanche sizes

We now compute the avalanche-size distribution. To do so, we have to invert Eq. (150). To first order in $\alpha$ this can be done order by order by first inserting in the term proportional to $\alpha$ the MF solution, i.e., the solution of Eq. (150) at $\alpha=0$. Solving for $\widetilde{Z}$ then yields

$$
\begin{equation*}
\tilde{Z}(\lambda)=\frac{1}{2}(1-\sqrt{1-4 \lambda})+\frac{[(3 \lambda+\sqrt{1-4 \lambda}-1) \log (1-4 \lambda)-2(2 \lambda+\sqrt{1-4 \lambda}-1)] \alpha}{4 \sqrt{1-4 \lambda}}+O\left(\alpha^{2}\right) \tag{166}
\end{equation*}
$$

Expanding in $\lambda$ gives

$$
\begin{align*}
\tilde{Z}(\lambda)= & {\left[\lambda+\lambda^{2}+2 \lambda^{3}+5 \lambda^{4}+14 \lambda^{5}+42 \lambda^{6}+O\left(\lambda^{7}\right)\right] } \\
& +\left(2 \lambda^{3}+\frac{35 \lambda^{4}}{3}+54 \lambda^{5}+\frac{3472 \lambda^{6}}{15}+O\left(\lambda^{7}\right)\right) \alpha+O\left(\alpha^{2}\right) \tag{167}
\end{align*}
$$

consistent with our previous result (144) if one uses the values of the $i_{n}$ given above. The challenge is to find a distribution $P(S)$ which generates all terms in Eq. (166). By trial and error, one arrives at the following integral representation:

$$
\begin{align*}
\tilde{Z}(\lambda)= & \frac{1}{2 \sqrt{\pi}} \int_{0}^{\infty} d S\left(e^{\lambda S}-1\right) e^{-S / 4} S^{-3 / 2} \\
& \times\left\{1+\frac{1}{16}[\log (S) S+\gamma S+4 S-8 \sqrt{\pi} \sqrt{S}\right. \\
& \left.-6 \log (S)-6 \gamma+4] \alpha+O\left(\alpha^{2}\right)\right\} \tag{168}
\end{align*}
$$

To leading order in $\alpha$, using definition (152), this yields our final result for the avalanche-size distribution at one loop

$$
\begin{equation*}
P(S)=\frac{\langle S\rangle}{2 \sqrt{\pi}} S_{m}^{\tau-2} A S^{-\tau} \exp \left(C \sqrt{\frac{S}{S_{m}}}-\frac{B}{4}\left[\frac{S}{S_{m}}\right]^{\delta}\right) \tag{169}
\end{equation*}
$$

for $S \gtrdot S_{\min }$, with coefficients

$$
\begin{gather*}
A=1+\frac{1}{8}\left(2-3 \gamma_{E}\right) \alpha, \quad B=1-\alpha\left(1+\frac{\gamma_{E}}{4}\right), \\
C=-\frac{1}{2} \sqrt{\pi} \alpha, \quad \alpha=-\frac{1}{3}\left(1-\zeta_{1}\right) \epsilon . \tag{170}
\end{gather*}
$$

$\gamma_{E}=0.577216$ and the exponents read

$$
\begin{equation*}
\tau=\frac{3}{2}+\frac{3}{8} \alpha=\frac{3}{2}-\frac{1}{8}\left(1-\zeta_{1}\right) \epsilon+O\left(\epsilon^{2}\right) \tag{171}
\end{equation*}
$$

$$
\begin{equation*}
\delta=1-\frac{\alpha}{4}=1+\frac{1}{12}\left(1-\zeta_{1}\right) \epsilon \tag{172}
\end{equation*}
$$

Note that the decay of large avalanches becomes stretched exponential, faster than exponential, with an exponent $\delta>1$. For the RF class $\zeta_{1}=1 / 3$, hence $\delta=1+\epsilon / 18+O\left(\epsilon^{2}\right)$. This goes in the right direction to match the exact result $\delta=3$ for the RF class in $d=0$, but the magnitude seems underestimated by the one-loop formula, which suggests large higherloop corrections for this exponent. Note the pronounced bump in the plot of $S^{\tau} P(S)$ on Fig. 2. This bump is indeed seen in numerical simulations [26].

## C. Normalization and scales in the distribution of sizes

One can check that final formula (169) is properly normalized, i.e., $\int_{0}^{\infty} d S S P(S)=\langle S\rangle$, to leading and first order in $\epsilon$. Note that since $\tau>1$, this formula does not give information about typical avalanches, which are of the order of the cutoff $S_{\min }$, but about larger ones which control the moments $\left\langle S^{p}\right\rangle$ with $p>p_{c}=\tau-1$. It is then universal, i.e., it does not depend on the details of the small-scale cutoff $S_{\min }$. However the average avalanche size $\langle S\rangle$, which appears as a factor in distribution (169), is nonuniversal and cannot be computed from this theory. In Sec. V F we showed that it behaves as $\langle S\rangle \sim S_{\min }^{\tau-1} S_{m}^{2-\tau}$ but this is all one knows about it. In comparing with numerics or experiments, one replaces $\langle S\rangle$ by its actual measured value. Then there is only one free parameter left, the global scale $S_{m}$. Note that this scale is given by $S_{m}$ $=\left\langle S^{2}\right\rangle / 2\langle S\rangle$; hence it can also easily be measured. We recall that its value is predicted by field theory via the exact relation


FIG. 2. (Color online) $P(s)$, both for mean field (black) and $\epsilon=1,2,3$ (blue, red, green), i.e. from bottom to top at $\log (s)=0$.

$$
\begin{equation*}
S_{m}=\frac{\left|\Delta^{\prime}\left(0^{+}\right)\right|}{m^{4}}=\frac{K^{d / 2}}{m^{d+\zeta}} \frac{\left|\widetilde{\Delta}^{\prime}\left(0^{+}\right)\right|}{\epsilon \widetilde{I}_{2}}, \tag{173}
\end{equation*}
$$

$\epsilon \widetilde{I}_{2}=2(4 \pi)^{-d / 2} \Gamma\left(3-\frac{d}{2}\right)$ is a $d$-dependent number. A question is how predictive this formula and how universal this scale is. The answer depends on the universality class, RP, RF, or RB.

For the RP class, $\zeta=0$ and the fixed point for $\left|\widetilde{\Delta}^{\prime}\left(0^{+}\right)\right| / a$, where $a$ is the period, is a universal number. To two loops, this number is equal in the statics and at depinning and reads $\epsilon / 6+\epsilon^{2} / 9+O\left(\epsilon^{3}\right)$. In general $K$ and $m$ will experience small corrections for any, e.g., lattice model which does not exactly satisfy the STS symmetry. They can however be measured from large scale measurements on the system. Once they are extracted, then $S_{m}$ can be predicted. Alternatively, one can construct ratios which, for the RP class, are universal, for instance,

$$
\begin{equation*}
r_{\mathrm{RP}}=\frac{a S_{m}}{\overline{u^{2}} L^{d}}=\frac{a\left|\Delta^{\prime}\left(0^{+}\right)\right|}{\Delta(0)}=\frac{a\left|\widetilde{\Delta}^{\prime}\left(0^{+}\right)\right|}{\widetilde{\Delta}(0)}=6, \tag{174}
\end{equation*}
$$

where $u$ is the center of mass and $\overline{u^{2}}$ is the variance of its fluctuations. Its definition is $\overline{u^{2}} L^{d}=\Sigma_{x} u(0) u(x)$. Note that in the statics there is no higher order correction since the fixed point for $R(u)=a+b[u(1-u)]^{2}$ with some constants $a$ and $b$ which drop out from Eq. (174). Replacing by the depinning values gives $r=6+2 \epsilon+O\left(\epsilon^{2}\right)$. This universal ratio can thus be used to distinguish statics from depinning.

For the RF class, $\zeta=\epsilon / 3$ in the statics, as recalled in Appendix B. The fixed point values contain a scale $\xi$ (in the direction of $u$ ), which can be fixed if one knows the amplitude of the random field $\sigma=\int_{0}^{\infty} d u \Delta(u) . \sigma$ can be retrieved from large distances in experiments and is a parameter of simulations. There are various amplitude combinations which can be studied, depending on whether one is willing to measure $m$ and $K$ (see Appendix B and Ref. [92]). The nicest universal ratio is

$$
\begin{equation*}
r_{\mathrm{RF}}=\frac{\sigma S_{m}}{\left(m^{2} \overline{u^{2}} L^{d}\right)^{2}}=\frac{\sigma\left|\Delta^{\prime}\left(0^{+}\right)\right|}{\Delta(0)^{2}}=\gamma_{1}+\epsilon\left(\frac{2}{9} \gamma_{1}+\gamma_{2}\right)+O\left(\epsilon^{2}\right) \tag{175}
\end{equation*}
$$

with $\gamma_{1}=0.775304245188$ and $\gamma_{2}=-0.13945524$, where one can check from Appendix B that all dependence on $m, K$, and the scale $\xi$ cancel (thanks in part to the strict equality $\zeta=\epsilon / 3)$.

Finally, for the RB class (and for the RF class for depinning) there is the least universality: $\left|\widetilde{\Delta}^{\prime}\left(0^{+}\right)\right|$is nonuniversal, hence measuring $K$ and $m$ will not be enough. One has, for instance, also to measure $\overline{u^{2}} L^{d}$ to obtain the combination

$$
\begin{equation*}
\xlongequal[S_{m}^{2}]{u^{2} L^{d}}=\frac{\Delta^{\prime}\left(0^{+}\right)^{2}}{m^{4} \Delta(0)}=\frac{m^{-d} K^{d / 2}}{\epsilon \tilde{I}_{2}} \frac{\left[\widetilde{\Delta}^{\prime}\left(0^{+}\right)\right]^{2}}{\widetilde{\Delta}(0)} \tag{176}
\end{equation*}
$$

with the universal ratio

$$
\begin{equation*}
\frac{\widetilde{\Delta}^{\prime}\left(0^{+}\right)^{2}}{\widetilde{\Delta}(0)}=0.583405 \epsilon+0.294205 \epsilon^{2}+O\left(\epsilon^{3}\right) \tag{177}
\end{equation*}
$$

Of course, in all these cases there are other possibilities for interesting universal ratios. If one measures, for instance, the function $\Delta(u)$ as in Ref. [26], one can in some cases get rid of measuring $K$ and $m$. Eventually, a direct numerical test of the relation between $S_{m}$ and the cusp would also be welcome. Analytical solution of toy models in $d=0$ [90] has also successfully tested this relation.

## D. Moments and universal ratios

Having obtained $P(S)$ we can compute its moments. This is useful for comparison with numerics. Direct integration of Eq. (171) gives

$$
\begin{align*}
\frac{\left\langle S^{n}\right\rangle}{\langle S\rangle S_{m}^{n-1}}= & \frac{4^{n-1} \Gamma\left(n-\frac{1}{2}\right)}{\sqrt{\pi}}+\frac{\epsilon\left(1-\zeta_{1}\right)}{3} \frac{4^{n-2}}{\sqrt{\pi}} \\
& \times\left\{4 \sqrt{\pi} \Gamma(n)-\Gamma\left(n-\frac{1}{2}\right)\left[\left(\psi\left(n-\frac{1}{2}\right)\right.\right.\right. \\
& +\log (4)+\gamma)(n-2)+4 n]\}+O\left(\epsilon^{2}\right) \tag{178}
\end{align*}
$$

a formula valid for any fixed real $n>1 / 2$. From this we can extract the universal dimensionless ratios,

$$
\begin{align*}
r_{n} & :=\frac{\left\langle S^{n+1}\right\rangle\left\langle S^{n-1}\right\rangle}{\left\langle S^{n}\right\rangle^{2}} \\
& =\frac{2 n-1}{2 n-3}-\frac{\epsilon}{3}\left(1-\zeta_{1}\right) \frac{n \Gamma\left(n-\frac{3}{2}\right)+\sqrt{\pi} \Gamma(n-1)}{(2 n-3)^{2} \Gamma\left(n-\frac{3}{2}\right)}+O\left(\epsilon^{2}\right) \tag{179}
\end{align*}
$$

for any real $n>3 / 2$, with $\zeta_{1}=1 / 3$ for RF, $\zeta_{1}=0$ for RP, and $\zeta_{1}=0.208298042$ for RB. The lowest-order integer ones are

$$
\begin{gather*}
r_{2}=3-\epsilon\left(1-\zeta_{1}\right),  \tag{180}\\
r_{3}=\frac{5}{3}-\frac{5}{27} \epsilon\left(1-\zeta_{1}\right) . \tag{181}
\end{gather*}
$$

Another useful form when comparing to numerics for $n$ near $\tau$ is to isolate the simple pole divergence which occurs in any dimension,

$$
\begin{gather*}
r_{n}=\frac{A_{d}}{n-\tau}+B_{n, d},  \tag{182}\\
A_{d}=1-\frac{1+\pi}{12} \epsilon\left(1-\zeta_{1}\right)+O\left(\epsilon^{2}\right), \\
B_{n, d}=1+\epsilon\left(1-\zeta_{1}\right) \frac{\pi \Gamma[n-1 / 2]-\sqrt{\pi} \Gamma(n-1)}{6(2 n-3) \Gamma\left(n-\frac{1}{2}\right)}+O\left(\epsilon^{2}\right) . \tag{183}
\end{gather*}
$$

Note that for comparison with numerics it is useful to estimate the corrections due to the small scale cutoff, assuming it just cuts the previous result at $S_{\text {min }}$. The correction to the above result for the dimensionless quantity $\frac{\left\langle S^{n}\right\rangle}{\langle S\rangle\rangle_{m}^{n-1}}$ is thus of order $\left(S_{\min } / S_{m}\right)^{n-\tau+1} /(n-\tau+1)$ and an approximation to $r_{n}$ is thus

$$
\begin{equation*}
r_{n}=A_{d} \frac{1-\left(S_{\min } / S_{m}\right)^{n-\tau}}{n-\tau}+B_{n, d} \tag{184}
\end{equation*}
$$

For the RF class one finds either from a direct expansion or using the $(0,1)$ Padé: $r_{2}=\{2.333,2.45455\}$, $r_{3}=\{1.5432,1.55172\} \quad(d=3), \quad r_{2}=\{1.667,2.07692\}$, $r_{3}=\{1.41975,1.45161\} \quad(d=2), \quad r_{2}=\{1,1.8\}$, $r_{3}=\{1.2963,1.36364\} \quad(d=1)$, and $r_{2}=\{0.3333,1.58824\}$, $r_{3}=\{1.17284,1.28571\}(d=0)$.

This can be compared with the exact result [62] for RF disorder in $d=0$,

$$
\begin{gather*}
r_{2}=1.2978  \tag{185}\\
r_{3}=1.17776 \tag{186}
\end{gather*}
$$

Using the exact result as a constraint yields three Padé approximants and the results

$$
\begin{align*}
& r_{2}=2.404 \pm 0.009, \quad d=3,  \tag{187}\\
& r_{2}=1.935 \pm 0.021, \quad d=2,  \tag{188}\\
& r_{2}=1.571 \pm 0.022, \quad d=1, \tag{189}
\end{align*}
$$

and

$$
\begin{gather*}
r_{3}=1.5427 \pm 0.0012, \quad d=3  \tag{190}\\
r_{3}=1.419 \pm 0.003, \quad d=2  \tag{191}\\
r_{3}=1.297 \pm 0.003, \quad d=1 \tag{192}
\end{gather*}
$$

where averages are over the three Padé approximants and error bars are the corresponding one sigma deviations.

## IX. SPATIAL STRUCTURE OF AVALANCHE DISTRIBUTIONS AND SELF-CONSISTENT EQUATION AT NONZERO MOMENTUM

In this section we discuss generating functions which encode for the spatial correlations in the avalanches. An explicit calculation is performed at the level of the improved tree approximation (mean field). It exhibits an interesting connection to instanton calculations in a cubic field theory.

## A. Generating function

To obtain information about the structure of avalanches in internal space one may define for avalanche $i$,

$$
\begin{equation*}
S_{i}^{\phi}=\int_{x} \phi(x) S_{i}^{x} \tag{193}
\end{equation*}
$$

where $\phi(x)$ is a given function. One recovers the standard definition of size for $\phi(x)=1$, i.e., $S_{i}=S_{i}^{1}$. One would like to compute averages such as [93]

$$
\begin{align*}
Z^{\phi}(\lambda) & =\frac{1}{\int_{x} \phi(x)} \sum_{i}\left[\exp \left(\lambda \int_{x} \phi(x) S_{i}^{x}\right)-1\right] \delta\left(w-w_{i}\right) \\
& =\frac{1}{\int_{x} \phi(x)} \int_{0}^{\infty} d S^{\phi}\left(e^{\lambda S^{\phi}}-1\right) \rho_{\phi}\left(S^{\phi}\right) \tag{194}
\end{align*}
$$

where $\rho_{\phi}(s)=\rho_{0} P_{\phi}(s)$ is the density of avalanches with $S_{\phi}=s$ and $P_{\phi}(s)$ is the normalized distribution of $S_{\phi}$. Note that STS implies $\int_{x} \phi(x) \overline{\left[u_{x}(w)-w\right]}=0$ for any $\phi(x)$ and $w$. Hence, taking a derivative with respect to $w$ one obtains

$$
\begin{equation*}
\int_{0}^{\infty} d S^{\phi} S^{\phi} \rho_{\phi}\left(S^{\phi}\right)=\int_{x} \phi(x) \tag{195}
\end{equation*}
$$

Note that $\rho_{0}$, the total density of avalanches, is independent of $\phi$ : hence one also has the exact relation

$$
\begin{equation*}
\left\langle S^{\phi}\right\rangle=\langle S\rangle L^{-d} \int_{x} \phi(x) \tag{196}
\end{equation*}
$$

Extending the arguments leading to Eqs. (A4) and (A5) we write

$$
\begin{align*}
& \left.\partial_{w} \exp \left(\lambda\left\{\int_{x} \phi(x)\left[u_{x}(w)-w-u_{x}(0)\right]\right\}\right)\right|_{w=0^{+}} \\
& =\int d S^{\phi} \rho_{\phi}\left(S^{\phi}\right)\left(e^{\lambda S^{\phi}}-1-\lambda S^{\phi}\right) \tag{197}
\end{align*}
$$

Hence one needs to compute the generating function,

$$
\begin{gather*}
G^{\phi}(\lambda)=\frac{1}{\int_{x} \phi(x)} \exp \left(\lambda\left\{\int_{x} \phi(x)\left[u_{x}(w)-w-u_{x}(0)\right]\right\}\right)  \tag{198}\\
G^{\phi}(\lambda)=\hat{Z}^{\phi}(\lambda) w+O\left(w^{2}\right),  \tag{199}\\
\hat{Z}^{\phi}(\lambda)=\frac{1}{\left\langle S^{\phi}\right\rangle}\left(\left\langle e^{\lambda S^{\phi}}\right\rangle-1-\lambda\left\langle S^{\phi}\right\rangle\right) \tag{200}
\end{gather*}
$$

for $w>0$, with again $Z^{\phi}(\lambda)=\lambda+\hat{Z}^{\phi}(\lambda)$. Note that here we consider only a uniform $w_{x}=w$.

## B. Self-consistent equation

Let us now study this quantity in the improved tree approximation (also called mean field above). We can show that it is given by

$$
\begin{equation*}
Z^{\phi}(\lambda)=\frac{1}{\int_{x} \phi(x)} \int_{x} Z_{x}^{\phi}(\lambda) \tag{201}
\end{equation*}
$$

where $Z_{x}^{\phi}(\lambda)$ is solution to the self-consistent equation,

$$
\begin{equation*}
Z_{x}^{\phi}(\lambda)=\lambda \phi(x)+\left|\Delta^{\prime}\left(0^{+}\right)\right| \int_{z y} g_{x-y} g_{x-z} Z_{y}^{\phi}(\lambda) Z_{z}^{\phi}(\lambda) \tag{202}
\end{equation*}
$$

$g_{x}=\int_{k} g_{k} e^{i k x}$ is the free (elastic) propagator. One can check that for $\phi(x)=1, Z_{x}(\lambda)=Z(\lambda)$ and one recovers the tree-level recursion given in the text for $Z(\lambda)$, and, upon replacing $\Delta^{\prime}\left(0^{+}\right) \rightarrow 1$ and $g_{k} \rightarrow \widetilde{g}_{k}=\left.g_{k}\right|_{m=1}$, the one for the rescaled function $\tilde{Z}(\lambda)$. A derivation, resumming all diagrams, is given in Appendix F 1.

We now go to rescaled quantities, using Eq. (91), i.e., $Z(\lambda)=\widetilde{Z}\left(\lambda S_{m}\right) / S_{m}$. Then we can write for an arbitrary function $\phi(k)$ in Fourier space

$$
\begin{equation*}
\widetilde{Z}_{k}(\lambda)=\lambda \phi(k)+\int_{q} \frac{m^{2}}{q^{2}+m^{2}} \frac{m^{2}}{(k-q)^{2}+m^{2}} \widetilde{Z}_{q}(\lambda) \widetilde{Z}_{k-q}(\lambda) \tag{203}
\end{equation*}
$$

In real space, this is

$$
\begin{equation*}
\tilde{Z}_{z}(\lambda)=\lambda \phi(z)+m^{4} \int_{x y} g_{z-x} g_{z-y} \widetilde{Z}_{x}(\lambda) \widetilde{Z}_{y}(\lambda) \tag{204}
\end{equation*}
$$

We note that this self-consistent equation is simplified by defining $Y_{k}(\lambda):=m^{2} g_{k} \widetilde{Z}_{k}(\lambda)$ or in real space

$$
\begin{gather*}
Y(z m, \lambda):=m^{2} \int_{x} g_{z-x} \tilde{Z}_{x}(\lambda)  \tag{205}\\
\varphi(m z):=\phi(z) \tag{206}
\end{gather*}
$$

This results in (suppressing from now on the explicit dependence on $\lambda$ when convenient)

$$
\begin{equation*}
\left(-\nabla^{2}+1\right) Y(x)=\lambda \varphi(x)+Y(x)^{2} \tag{207}
\end{equation*}
$$

## C. Solution for $\varphi$ localized on a codimension one hyperplane

The function $Z(\lambda)$ will be qualitatively different, depending on whether $\varphi$ is extended on the scale of the inverse mass (absorbed in $x$ ) or is a $\delta$ distribution. We now study one special case, namely, $\varphi(x)=\delta\left(x_{1}\right)$. The function $Y(x)$ will then be constant along the directions $x_{2}, \ldots, x_{d}$, and for simplicity of notation we will denote $x_{1} \rightarrow x$ and suppress $x_{i}, i$ $>1$. Thus we effectively consider a one-dimensional problem.

Equation (207) can then be integrated analytically. Consider first the homogenous equation $(\lambda=0)$,

$$
\begin{equation*}
Y^{\prime \prime}(x)=Y(x)-Y(x)^{2} \tag{208}
\end{equation*}
$$

Multiplying with $Y^{\prime}$ and integrating once give

$$
\begin{equation*}
\left[Y^{\prime}(x)\right]^{2}=\text { const }-\frac{2}{3} Y(x)^{3}+Y(x)^{2} \tag{209}
\end{equation*}
$$

If Eq. (208) is viewed as the equation of motion of a particle, then setting const $\rightarrow 0$ will be the solution which has zero kinetic energy at the saddle point $Y=0$ of the potential $V(Y)=\frac{1}{3} Y^{3}-\frac{1}{2} Y^{2}$. It is the unique solution which decays (exponentially fast) to $Y=0$ for $x \rightarrow \pm \infty$. The other solutions are either oscillating or unbounded. Integrating once more for const $=0$ gives $-2 \operatorname{arctanh}\left(\frac{\sqrt{3-2 y}}{\sqrt{3}}\right)=x$ or equivalently

$$
\begin{equation*}
Y(x)=Y_{0}(x):=\frac{3}{\cosh (x)+1} \tag{210}
\end{equation*}
$$

where the center of the solution has been chosen to be at $x$ $=0$.

A symmetric solution $Y(x)=Y(-x)$ of Eq. (207) with $\varphi(x)=\delta(x)$ can now be constructed as follows:

$$
\begin{equation*}
Y(x, \lambda)=\frac{3}{\cosh \left(x+x_{0}(\lambda)\right)+1} \quad \text { for } x>0 \tag{211}
\end{equation*}
$$

Inserting into Eq. (207) and integrating from $-\delta$ to $\delta$ gives

$$
\begin{equation*}
-2 \partial_{\delta} Y(\delta, \lambda)=\lambda+O(\delta) \tag{212}
\end{equation*}
$$

Thus, in the limit of $\delta \rightarrow 0$,

$$
\begin{equation*}
\frac{\sinh \left(x_{0}(\lambda)\right)}{\left[1+\cosh \left(x_{0}(\lambda)\right)\right]^{2}}=\frac{\lambda}{6} \tag{213}
\end{equation*}
$$

Let us consider the unique real branch, such that $x_{0}(\lambda) \rightarrow \infty$ when $\lambda \rightarrow 0$, which gives $Z^{\phi}(\lambda) \rightarrow 0$ in the same limit. Increasing $\lambda$ from 0 , the solution breaks down when $\lambda$ reaches $\lambda_{c}$ with

$$
\begin{equation*}
\lambda_{c}=\frac{2}{\sqrt{3}} \tag{214}
\end{equation*}
$$

such that $\cosh \left[x_{0}\left(\lambda_{c}\right)\right]=2$. Now we need $Z^{\phi}(\lambda)$ defined in Eq. (201) or the rescaled version $\widetilde{Z}^{\phi}(\lambda)=\frac{1}{\int_{x} \phi(x)} \int_{x} \widetilde{Z}_{y}^{\phi}(\lambda)$,

$$
\begin{equation*}
\widetilde{Z}^{\phi}(\lambda)=\frac{1}{\int_{x} \phi(x)} \int_{x} \widetilde{Z}_{x}^{\phi}(\lambda)=\frac{1}{\int_{x} \varphi(x)} \int_{x}\left(-\nabla^{2}+1\right) Y(x, \lambda), \tag{215}
\end{equation*}
$$

where from the first to the second line we switched to dimensionless variables. Inserting Eq. (207) and using that $\varphi(x)$ $=\delta(x)$ yield

$$
\begin{align*}
\widetilde{Z}^{\phi}(\lambda) & =\int_{x} \lambda \varphi(x)+Y^{2}(x, \lambda) \\
& =\lambda+\int_{0}^{\infty} d x \frac{18}{\left[1+\cosh \left(x+x_{0}(\lambda)\right)\right]^{2}} \\
& =\lambda+12 \frac{1+3 e^{x_{0}(\lambda)}}{\left[1+e^{x_{0}(\lambda)}\right]^{3}}=\frac{12}{1+e^{x_{0}(\lambda)}}, \tag{216}
\end{align*}
$$

where in the last line Eq. (213) was used. Therefore $x_{0}(\lambda)$ can be expressed in terms of $\widetilde{Z}^{\phi}(\lambda)$, and inserted into Eq. (213), with the result

$$
\begin{equation*}
\lambda=\frac{\widetilde{Z}^{\phi}\left(\widetilde{Z}^{\phi}-6\right)\left(\widetilde{Z}^{\phi}-12\right)}{72} \tag{217}
\end{equation*}
$$

Note that this result was derived for $\lambda \in\left[0, \lambda_{c}\right], \widetilde{Z}^{\phi} \in[0,6$ $-2 \sqrt{3}]$, which corresponds to the red (solid) branch in Fig. 3 and then (implicitly) continued analytically to all $\lambda$ and all $\widetilde{Z}^{\phi}$. We note that the green part of the curve (long dashes) can also be obtained analytically using the second point for which Eq. (213) is satisfied and the two solutions for negative $\lambda$. One easily checks that the symmetry $Y(x)=Y(-x)$ implies point-reflection symmetry around the point $\left(\lambda, \widetilde{Z}^{\phi}\right)$ $=(0,6)$. This already strongly suggests that $\lambda$ is a third-order polynomial in $\widetilde{Z}^{\phi}$, symmetric around the point $\widetilde{Z}^{\phi}=6$ (in our chosen units).

For large negative $\lambda,\left|\widetilde{Z}^{\phi}(\lambda)\right| \sim \sqrt[3]{|\lambda|}$, thus from our discussion following Eq. (153), $p_{c}=\frac{1}{3}$, and the avalanche exponent $\tau$ is


FIG. 3. (Color online) $Z^{\phi}(\lambda)$ for $\phi(x)=\delta(x)$, as explained in the main text.

$$
\begin{equation*}
\tau^{\phi}=\frac{4}{3} \tag{218}
\end{equation*}
$$

Note that this value is consistent with a generalized conjecture that we can put forward, $\tau^{\phi}=2-\frac{2}{d_{\phi}+\zeta}$, where $d_{\phi}=d-1$ for a codimension one subspace, inserting $d=4$ and $\zeta=0$ since the above result was derived at mean-field level. Again, much work remains to validate or invalidate this generalized conjecture (e.g., a two loop calculation).

Interestingly, the probability distribution is nontrivial and different from the standard mean-field one (i.e., for $\phi=1$ ). It takes the form, for $S^{\phi} \gg S_{\min }^{\phi}$,

$$
\begin{equation*}
P^{\phi}\left(S^{\phi}\right)=\frac{1}{S_{m}} p^{\phi}\left(\frac{S^{\phi}}{S_{m}}\right) \tag{219}
\end{equation*}
$$

Here the avalanche sizes are defined as $S_{i}^{\phi}$ $=\frac{1}{m} \int_{x_{2}, \ldots, x_{d}} S_{x_{1}=0, x_{2}, \ldots, x_{d}}^{i} \sim m^{-d-\zeta}$ and $S_{m}$ is the same quantity as defined in Eq. (93). The distribution $p^{\phi}(s)$ can be calculated as follows:

$$
\begin{align*}
\langle s\rangle^{-1} p^{\phi}(s)= & \int_{-i \infty}^{i \infty} \frac{d \lambda}{2 \pi i} e^{-\lambda s} \widetilde{Z}^{\phi}(\lambda) \\
= & \int_{-i \infty}^{i \infty} \frac{d Z}{2 \pi i} \frac{d \lambda(Z)}{d Z} e^{-\lambda(Z) s} Z \\
= & \frac{18}{\pi} \int_{0}^{\infty} d x\left(3 x^{2}+1\right)\left\{x \sin \left[3 s x\left(x^{2}+1\right)\right]\right. \\
& \left.-\cos \left[3 s x\left(x^{2}+1\right)\right]\right\} \\
= & \frac{2 K_{1 / 3}\left(\frac{2 s}{\sqrt{3}}\right)}{\pi s} \tag{220}
\end{align*}
$$

Here $s=S^{\phi} / S_{m}$ and $\langle s\rangle=\left\langle S^{\phi}\right\rangle / S_{m}$, and $K_{1 / 3}$ is a Bessel function. The steps of the derivation are a change in variables from $\lambda$ to $Z=\widetilde{Z}^{\phi}(\lambda)$ a change in variables $Z=6 i(x-i)$, and combining the integrand for $x$ and $-x$. For large $s$ the asymptotic behavior is similar to the standard mean-field result $P(s) \sim s^{-3 / 2} e^{-\lambda_{c} s}$,

$$
\begin{equation*}
p^{\phi}(s)=\langle s\rangle e^{-2 s / \sqrt{3}}\left(\frac{\sqrt[4]{3}\left(\frac{1}{s}\right)^{3 / 2}}{\sqrt{\pi}}-\frac{5\left(\frac{1}{s}\right)^{5 / 2}}{48(\sqrt[4]{3} \sqrt{\pi})}+\cdots\right) \tag{221}
\end{equation*}
$$

This was expected since in both cases the solution for $Z(\lambda)$ ends at a $\lambda_{c}$ with a square-root singularity. For small $s$ the asymptotic behavior is in accordance with Eq. (218),

$$
\begin{equation*}
\langle s\rangle^{-1} p^{\phi}(s)=\frac{\sqrt[6]{3} \Gamma\left(\frac{1}{3}\right)}{\pi s^{4 / 3}}+\frac{\Gamma\left(-\frac{1}{3}\right)}{\sqrt[6]{3} \pi s^{2 / 3}}+\frac{\sqrt[6]{3} \Gamma\left(\frac{1}{3}\right) s^{2 / 3}}{2 \pi}+\cdots . \tag{222}
\end{equation*}
$$

Finally, from Eq. (220) one obtains the moments:

$$
\begin{equation*}
\frac{\left\langle\left(S^{\phi}\right)^{n}\right\rangle}{\left\langle S^{\phi}\right\rangle}=\frac{3^{n / 2} \Gamma\left(\frac{n}{2}-\frac{1}{6}\right) \Gamma\left(\frac{n}{2}+\frac{1}{6}\right)}{2 \pi} S_{m}^{n-1} \tag{223}
\end{equation*}
$$

for $n>1 / 3$, from which the universal ratios $r_{n}$ can be computed. In particular that $\left\langle\left(S^{\phi}\right)^{2}\right\rangle /\left\langle S^{\phi}\right\rangle=\frac{1}{2} S_{m}$ instead of $2 S_{m}$ for $\phi=1$.

Note that for $\phi$ localized on a hyperplane, Eq. (196) implies

$$
\begin{equation*}
\left\langle S^{\phi}\right\rangle=\frac{1}{L m}\langle S\rangle . \tag{224}
\end{equation*}
$$

This is understood from the observation, which only a fraction $1 /(L m)$ of all avalanches leads to an advance of the interface (avalanche) in the hyperplane. Thus, if one was to define an avalanche distribution by considering only the hyperplane, one would naturally choose a different normalization.

## D. Diagrammatic expansion

Here we have solved the codimension one case, i.e., $d_{\phi}$ $=d-1$. More generally one may consider the case $d_{\phi}=d$ $-d^{\prime}$, i.e., a function $\varphi(x)$ localized on a $d^{\prime} \leqslant d$ dimensional hyperplane. This amounts to study Eq. (207) in dimension $d^{\prime}$. This can be done graphically, as Eq. (207) can be used to define, in any dimension $d^{\prime}$, a diagrammatic expansion,


$$
\begin{equation*}
+O\left(\lambda^{5}\right) \tag{225}
\end{equation*}
$$

where independent $d^{\prime}$-dimensional momenta flow in each loop, the solid lines are propagators $\widetilde{g}_{k}=1 /\left(k^{2}+1\right)$, and each loop contains a momentum integration. Note that this expansion is valid, i.e., IR and UV convergent in all dimensions $d^{\prime}<4$ (hence any $d>0$ for the original problem). In $d^{\prime}=0$ it
reproduces Eq. (89). In $d^{\prime}=1$ the integrals are easily computed by switching from Fourier to real space where $\widetilde{g}_{x}$ $=\frac{1}{2} e^{-|x|}$, and one obtains

$$
\begin{equation*}
Z(\lambda)=\lambda+\frac{1}{4} \lambda^{2}+\frac{1}{9} \lambda^{3}+\left[\frac{11}{864}+4 \frac{83}{6912}\right] \lambda^{4}+\cdots \tag{226}
\end{equation*}
$$

This is in agreement with analytical result (217),

$$
\begin{align*}
Z(\lambda)= & \lambda+\frac{\lambda^{2}}{4}+\frac{\lambda^{3}}{9}+\frac{35 \lambda^{4}}{576}+\frac{\lambda^{5}}{27}+\frac{1001 \lambda^{6}}{41472}+\frac{4 \lambda^{7}}{243} \\
& +\frac{46189 \lambda^{8}}{3981312}+\frac{55 \lambda^{9}}{6561}+\frac{5311735 \lambda^{10}}{859963392}+O\left(\lambda^{11}\right) \tag{227}
\end{align*}
$$

It is amazing that the quite nontrivial series (225) is indeed resummed by the instanton solution, Eq. (211) in the case $d^{\prime}=1$. Study of higher $d^{\prime}$ and extension to the loop expansion is left for the future.

## X. NONLOCAL ELASTICITY AND AVALANCHE DISTRIBUTIONS FOR A CONTACT LINE

The calculations performed in this paper can be extended to a broader class of elasticity functions $g_{q}^{-1}$ in Eq. (2). One possible generalization is

$$
\begin{equation*}
g_{q}^{-1}=|q|^{\gamma}+\mu^{\gamma} \tag{228}
\end{equation*}
$$

where $\gamma=2$ corresponds to the choice studied above of local elasticity, while the case $\gamma=1$ is relevant for long-range elasticity as in contact lines of fluids. In fact, a more realistic form in that case is

$$
\begin{equation*}
g_{q}^{-1}=\sqrt{q^{2}+\mu^{2}} \tag{229}
\end{equation*}
$$

The mass is provided by capillarity, i.e., the interplay of surface tension and gravity. We keep the notation $m^{2}=g_{q=0}^{-1}$ $=\mu^{\gamma}$ for the curvature of the quadratic well, while distinguishing it from $\mu$ which is an inverse characteristic internal length in the interface (the inverse capillary length). One parametrization which contains these three cases is

$$
\begin{equation*}
g_{q}^{-1}=\mu^{\gamma} g(q / \mu) \tag{230}
\end{equation*}
$$

for some function $g(y) \sim y^{\gamma}$ at large $y$ and $g(0)=1$. Of course other scales may be present in a realistic problem, and form (230) assumes that $\mu^{-1}$ is the single largest internal length scale which cuts off the avalanches.

The upper critical dimension $d_{c}$ in all cases is given by the UV divergence of

$$
\begin{equation*}
I_{2}=\int_{q} g_{q}^{2}=C_{d, \gamma} \mu^{-\epsilon} \frac{1}{\epsilon} \tag{231}
\end{equation*}
$$

where $\epsilon=d_{c}-d, d_{c}=2 \gamma$. The constant $C_{d, \gamma}=\epsilon \tilde{I}_{2}$ with $\tilde{I}_{2}$ $=\left.\int_{q} g_{q}^{2}\right|_{\mu=1}$ remains finite at $d_{c}$.

With these choices all calculations of this paper are easily extended. Appendix B can still be used, up to trivial changes, e.g., replacing $m$ by $\mu$ in all final formula there, including in definition (B3) of the rescaled disorder (we can set $K=1$ ). One has again

$$
\begin{equation*}
\alpha=-\widetilde{\Delta}^{\prime \prime}\left(0^{+}\right)=-\frac{\epsilon}{3}\left(1-\zeta_{1}\right)+O\left(\epsilon^{2}\right) \tag{232}
\end{equation*}
$$

to one loop, with the same values for $\zeta_{1}$, independent of $\gamma$ [and $g(y)$ ] and of course now $\epsilon=d_{c}-d$ everywhere. For more details, including the calculation of the exponent $\zeta$ (and $z$ for the depinning) to two loops for the nonlocal elasticity see Refs. [36,49]. Here we compute the avalanche distribution to one loop.

To do that one starts again from Eq. (151). It generalizes as follows:

$$
\begin{equation*}
\widetilde{Z}(\lambda)=: \widetilde{Z}=\lambda+\widetilde{Z}^{2}+\alpha \mathcal{J}_{\gamma}, \tag{233}
\end{equation*}
$$

$\mathcal{J}_{\gamma}=\frac{1}{\epsilon \widetilde{I}_{2}} \int_{k} \frac{\widetilde{Z}^{2}}{\left(\widetilde{g}_{k}^{-1}-2 \widetilde{Z}\right)^{2}}+\frac{\tilde{Z}}{\widetilde{g}_{k}^{1}-2 \tilde{Z}}-\widetilde{Z} \widetilde{g}_{k}-3 \widetilde{Z}^{2} \widetilde{g}_{k}^{2}$,
where $\widetilde{g}_{k}=\left.g_{k}\right|_{\mu=1}$. Note that the definition of the rescaled $\tilde{Z}$ is still given by Eq. (152) with now

$$
\begin{equation*}
S_{m}=m^{-4} \Delta^{\prime}\left(0^{+}\right)=\mu^{-2 \gamma} \Delta^{\prime}\left(0^{+}\right)=\left(\epsilon \tilde{I}_{2}\right)^{-1} \widetilde{\Delta}^{\prime}\left(0^{+}\right) \mu^{-d+\zeta} \tag{235}
\end{equation*}
$$

The calculations are performed in Appendix E. For the choice of elasticity (228) we find again, for any $\gamma$, the distribution $P(S)$ given by Eq. (169), with amplitudes

$$
\begin{gather*}
A=1+\frac{1}{4 \gamma}\left(2-3 \gamma_{E}\right) \alpha, \quad B=1-\frac{2 \alpha}{\gamma}\left(1+\frac{\gamma_{E}}{4}\right), \\
C=-\frac{1}{\gamma} \sqrt{\pi} \alpha \tag{236}
\end{gather*}
$$

and exponents

$$
\begin{gather*}
\tau=\frac{3}{2}+\frac{3}{4 \gamma} \alpha=\frac{3}{2}-\frac{1}{4 \gamma}\left(1-\zeta_{1}\right) \epsilon+O\left(\epsilon^{2}\right)  \tag{237}\\
\delta=1-\frac{\alpha}{2 \gamma}=1+\frac{1}{6 \gamma}\left(1-\zeta_{1}\right) \epsilon+O\left(\epsilon^{2}\right) \tag{238}
\end{gather*}
$$

Of course everywhere $\epsilon=2 \gamma-d$. Note that the one-loop result for the avalanche exponent $\tau$ is now compatible to $O(\epsilon)$ with the generalized conjecture

$$
\begin{equation*}
\tau=2-\frac{\gamma}{d+\zeta} \tag{239}
\end{equation*}
$$

This conjecture can again be re-expressed as

$$
\begin{equation*}
\rho=2 \tag{240}
\end{equation*}
$$

given that scaling form (22) generalizes into

$$
\begin{equation*}
\rho(S)=L^{d} m^{\rho} S^{-\tau} \widetilde{\rho}\left(S \mu^{d+\zeta}\right) \tag{241}
\end{equation*}
$$

and that STS relation (21) still holds: hence now the exponent relation (always true) reads $\rho=2(2-\tau)(d+\zeta) / \gamma$.

In the case $\gamma=1$ and for form (229) more suitable to describe the contact line, one finds again the same exponent $\tau$ [Eq. (237)]: however the shape of the distribution is different. We find (see Appendix E)

$$
\begin{align*}
P(S)= & A^{\prime} \frac{\langle S\rangle}{2 \sqrt{\pi}} S_{m}^{-2}\left[\left(\frac{S}{S_{m}}\right)^{-\tau}+D^{\prime}\right] \\
& \times \exp \left(C^{\prime} \sqrt{\frac{S}{S_{m}}}-\frac{B^{\prime}}{4}\left[\frac{S}{S_{m}}\right]^{\delta^{\prime}}\right) \tag{242}
\end{align*}
$$

for $S \gtrdot S_{\min }$, with amplitudes

$$
\begin{gather*}
A^{\prime}=1+\frac{1}{4}\left(8-3 \gamma_{E}\right) \alpha, \quad B^{\prime}=1-\frac{3}{2} \alpha\left(\gamma_{E}-2\right), \\
C^{\prime}=-\frac{3}{2} \sqrt{\pi} \alpha, \quad D^{\prime}=\frac{1}{8} \sqrt{\pi} \alpha \tag{243}
\end{gather*}
$$

The exponent $\delta$ is now given by

$$
\begin{equation*}
\delta^{\prime}=1-\frac{3 \alpha}{2}=1+\frac{1}{2}\left(1-\zeta_{1}\right) \epsilon+O\left(\epsilon^{2}\right) \tag{244}
\end{equation*}
$$

Note that the presence of the constant $D^{\prime}$ suggests that the pre-exponential power law at large $S$ is not $S^{-\tau}$ anymore.

## XI. CONCLUSION

In conclusion we have succeeded in computing from the FRG the distribution of shock or jump sizes which occur in the minimal energy configuration of an interface pinned in a random potential and tied to a spring of varying position. These are the static analog of the avalanches which occur in the dynamics if the interface is instead pulled from a metastable configuration to the next one. Hence it opens the way to the same calculation in the dynamics, performed in [67], which yields very similar results. Shocks in the statics are often called static avalanches as there are many analogies, as well as some differences.

We believe that this is an important achievement. First because the FRG has been around for a while but it was not understood previously how to extract the avalanche statistics in a controlled way. In the sandpile literature this is still an open question despite many exact results for other quantities. It turns out to be conceptually simple (a posteriori) to extract these distributions from the FRG. In fact, the beautifully simple relation (36) between the cusp of the FRG function $\Delta(u)$ and the avalanche-size's second moment, unveiled in this work, gives a very transparent physical picture of the cusp. Similar relations hold for all moments, and the challenge is to sum them. This part is a priori technically difficult but some surprising simplifications occur in the calculation, which in the end lead to a simple self-consistent equation, with an interpretation in graph theory. This suggests a, yet to be discovered, simpler and presumably more powerful structure behind the present state of the art of the theory.

The one-loop result for the probability distribution is equally striking. First the avalanche exponent $\tau$ is found equal to order $O(\epsilon)$ to a conjectured form, put forward for the depinning transition. Note that our result is a firstprinciples derivation of this exponent. Hence the conjecture is confirmed to one loop. It is then of high interest to look for possible deviations to two loop, both in the statics and dynamics. Alternatively, if the conjecture is true in general, it
would be interesting to derive this from first principles in the field theory. Besides the exponent, we obtained the general formula for the distribution $P(S)$. This includes the precise way in which it is cut off for large avalanches at a cutoff scale $S_{m}$. These large and rare events are the important ones in terms of moments of distributions, and, in real life, if one is interested in, e.g., earthquakes. We predict the scale $S_{m}$ and the distribution, which is insensitive to details of the model at short scales (up to a single and measurable global factor). Of course the distribution depends on the details of the model at large scales: here we use mainly a quadratic well to cut the size off, but this dependence can explicitly be computed within the theory. We illustrate this point by computing it for contact-line depinning which has a more complicated elastic energy.

Some progress was made to study the spatial structure of avalanches. Additional definitions of local avalanche sizes were given which integrate information about jumps within a subspace of the interface. It was shown that to obtain their probability distribution at mean-field level one must solve an instanton problem in a cubic field theory. An explicit solution was found for a subspace of codimension one, leading to an exponent $\tau=4 / 3$ at mean-field level, i.e., in $d=4$, and a size distribution involving a Bessel function. It is different from the usual mean-field size distributions, recovered here, with $\tau=3 / 2$, and an exponential function.

There are many open interesting questions which can now be addressed, both within the statics and the dynamics. First, one would like to know more about the spatial structure of the avalanches and their correlations. The first steps are to extend the instanton calculation to regions of arbitrary shapes and codimension and to study it within the loop expansion. The problem can also be extended to a manifold with $N$ components. Consequences for hysteresis loops of magnets deserve a new study. Applications to earthquakes are also of interest, especially if avalanche correlations can be handled (preshock and after shock). Plastic avalanches can now be studied, e.g., in the framework of Ref. [94].

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## APPENDIX A: SHOCK-SIZE DISTRIBUTIONS

## 1. Generating function

We start again from the shock decomposition [Eq. (16)]. The field $e^{\lambda L^{L}[u(w)-w]}$ has jump discontinuities at each shock
position, where its value is multiplied by a factor $e^{\lambda S_{i}}$. Hence one can write

$$
\begin{align*}
\left(\partial_{w}+\lambda L^{d}\right) e^{\lambda L^{d}[u(w)-w]} & =\sum_{i}\left(e^{\lambda S_{i}}-1\right) e^{\lambda L^{d}\left[u\left(w_{i}^{-}\right)-w\right]} \delta\left(w-w_{i}\right) \\
& =\sum_{i}\left(1-e^{-\lambda S_{i}}\right) e^{\lambda L^{d}\left[u\left(w_{i}^{+}\right)-w\right]} \delta\left(w-w_{i}\right), \tag{A1}
\end{align*}
$$

where $L^{d} u\left(w_{i}^{-}\right)=\sum_{j<i} S_{j}$ and $L^{d} u\left(w_{i}^{+}\right)=\sum_{j \leqslant i} S_{j}$. The labels of the shocks are ordered according to their spatial position $w_{i}$. Let us multiply these equations by the field $e^{\lambda L^{d}[u(w)-w]}$ taken at a different point, i.e., consider

$$
\begin{align*}
\left(\partial_{w_{1}}\right. & \left.+\lambda L^{d}\right) e^{\lambda L^{d}\left\{u\left(w_{1}\right)-w_{1}-\left[u\left(w_{2}\right)-w_{2}\right]\right\}} \\
& =\sum_{i}\left(e^{\lambda S_{i}}-1\right) e^{\lambda L^{d}\left[u\left(w_{i}^{-}\right)-w_{1}\right]} \delta\left(w_{1}-w_{i}\right) e^{-\lambda L^{d}\left[u\left(w_{2}\right)-w_{2}\right]} \tag{A2}
\end{align*}
$$

where $w_{1}$ and $w_{2}$ are generic points and different from the $w_{i}$ (even though the notation might suggest otherwise). If we consider the limit $w_{2}=w_{1}^{-}$in Eq. (A2) and average over disorder the term containing $u\left(w_{i}^{-}\right)$cancels and one obtains

$$
\begin{align*}
& \lim _{w_{2} \rightarrow w_{1}^{-}}\left(\partial_{w_{1}}+\lambda L^{d}\right) \overline{e^{\lambda L^{d}\left\{u\left(w_{1}\right)-w_{1}-\left[u\left(w_{2}\right)-w_{2}\right]\right\}}} \\
& \quad=\overline{\sum_{i}\left(e^{\lambda S_{i}}-1\right) \delta\left(w_{1}-w_{i}\right)} \\
& \quad=\int d S \rho(S)\left(e^{\lambda S}-1\right) \tag{A3}
\end{align*}
$$

A similar equation can be derived starting from Eq. (A1) and considering the limit $w_{2}=w_{1}^{+}$. Using translational invariance one finally obtains

$$
\begin{equation*}
\partial_{w}{\overline{\left.e^{\lambda L^{d}[u(w)-w-u(0)}\right]}}_{\mid w=0^{+}}=\int d S \rho(S)\left(e^{\lambda S}-1-\lambda S\right), \tag{A4}
\end{equation*}
$$

$$
\begin{equation*}
\left.\partial_{w} \overline{e^{\lambda L^{d}[u(w)-w-u(0)]}}\right|_{w=0^{-}}=\int d S \rho(S)\left(1-e^{-\lambda S}-\lambda S\right), \tag{A5}
\end{equation*}
$$

where we have used Eq. (20). One thus recovers Eq. (44) in the text. Another method to derive the part proportional to $|w|$ is to multiply Eq. (A1) at two different points (and setting $\lambda \rightarrow-\lambda$ for the second) and extract the single shock contribution,

$$
\begin{align*}
\left(\partial_{w_{2}}\right. & \left.+\lambda L^{d}\right)\left(\partial_{w_{1}}+\lambda L^{d}\right) \overline{e^{\lambda L^{d}\left\{u\left(w_{1}\right)-w_{1}-\left[u\left(w_{2}\right)-w_{2}\right]\right\}}} \\
& =\delta\left(w_{1}-w_{2}\right) \overline{\sum_{i}\left(e^{\lambda S_{i}}-1\right)\left(e^{-\lambda S_{i}}-1\right) \delta\left(w_{1}-w_{i}\right)}+\text { smooth } \tag{A6}
\end{align*}
$$

This yields, using translational invariance,

$$
\begin{align*}
& \left(-\partial_{w}^{2}+\lambda^{2} L^{2 d}\right) \overline{e^{\lambda L^{d}[u(w)-w-u(0)]}} \\
& \quad=2 \delta(w) \int d S \rho(S)[1-\cosh (\lambda S)]+\text { smooth } \tag{A7}
\end{align*}
$$

Integrating twice one recovers Eq. (44), but only the coefficient of $|w|$ is determined. To determine the coefficient of $w$ one needs another equation, e.g., as above, or has to use the symmetry under $w \rightarrow-w$.

Note that this can be generalized as follows: consider a set of $\lambda_{k}$ with $\sum_{k=1}^{n} \lambda_{k}=0$, then one has

$$
\begin{align*}
& \prod_{k=1}^{n}\left(\partial_{w_{k}}+\lambda L^{d}\right) \exp \left(L^{d} \sum_{k=1}^{n} \lambda_{k}\left[u\left(w_{k}\right)-w_{k}\right]\right) \\
& = \\
& \quad \delta\left(w_{1}-w_{2}\right) \cdots \delta\left(w_{1}-w_{n}\right)  \tag{A8}\\
& \quad \times \int d S \rho(S) \prod_{k=1}^{n}\left(e^{\lambda_{k} S}-1\right)+\text { less singular. }
\end{align*}
$$

## 2. Multishock size distribution

The multishock size distributions are encoded in the higher pinning force cumulants [Eq. (13)]. For instance, the third cumulant can be written as

$$
\begin{align*}
-\partial_{w_{1}} \partial_{w_{2}} \partial_{w_{3}} \hat{C}\left(w_{1}, w_{2}, w_{3}\right)= & m^{6} L^{2 d}{\overline{\left[u^{\prime}\left(w_{1}\right)-1\right]\left[u^{\prime}\left(w_{2}\right)-1\right]\left[u^{\prime}\left(w_{3}\right)-1\right]}}^{c} \\
= & m^{6} L^{-d} \int d S S^{3} \rho(S) \delta\left(w_{1}-w_{2}\right) \delta\left(w_{1}-w_{3}\right) \\
& +m^{6} L^{-d}\left[\int d S_{1} d S_{3} S_{1}^{2} S_{3} \rho_{c}\left(S_{1}, S_{3}, w_{1}, w_{3}\right) \delta\left(w_{1}-w_{2}\right)+2 \text { perm }\right] \\
& +m^{6} L^{-d} \int d S_{1} d S_{2} d S_{3} S_{1} S_{2} S_{3} \rho_{c}\left(S_{1}, S_{2}, S_{3}, w_{1}, w_{2}, w_{3}\right)  \tag{A9}\\
\rho_{c}\left(S_{1}, \ldots, S_{n}, w_{1}, \ldots, w_{n}\right)= & \sum_{i_{a} \text { all distinct }} \frac{\delta\left(S_{1}-S_{i_{1}}\right) \cdots \delta\left(S_{n}-S_{i_{n}}\right) \delta\left(w_{1}-w_{i_{1}}\right) \cdots \delta\left(w_{n}-w_{i_{n}}\right)}{}{ }^{c} \tag{A10}
\end{align*}
$$

which is easily generalized to any cumulant. Note that the connected parts and the -1 substractions conspire, using Eq. (20), to give the correct connected shock distributions (such that scaling with volume is straighforward). Such formulas show explicitly the structure of the singularities expected in the force $n$ cumulants as $p$ points are brought together, as a consequence of the assumption of a finite density of shocks (dilute shocks). In principle, it allows us to check this assumption by computing all moments within the $\epsilon$ expansion.

The generating function generalizing $Z$, and allowing to extract $\rho_{c}\left(S_{1}, S_{2}, w_{1}, w_{2}\right)$, which is a function of $w_{1}-w_{2}$, is constructed in analogy with Eqs. (44) and (45),

$$
\begin{gather*}
G\left(\lambda_{1}, \lambda_{2}, w_{1}, w_{2}, \delta_{1}, \delta_{2}\right)=Z\left(\lambda_{1}, \lambda_{2}, w_{1}-w_{2}\right) \delta_{1} \delta_{2}+\cdots,  \tag{A11}\\
G\left(\lambda_{1}, \lambda_{2}, w_{1}, w_{2}, \delta_{1}, \delta_{2}\right)=L^{-d}{\overline{\left(e^{\lambda_{1} L^{d}\left[u\left(w_{1}+\delta_{1}\right)-u\left(w_{1}\right)-\delta_{1}\right]}-1\right)\left(e^{\lambda_{2} L^{d}\left[u\left(w_{2}+\delta_{2}\right)-u\left(w_{2}\right)-\delta_{2}\right]}-1\right)}}_{c}^{c} \tag{A12}
\end{gather*}
$$

for $\delta_{1}, \delta_{2}>0$.

## APPENDIX B: REVIEW OF BASIC FRG RESULTS

We review basic equations for functional RG, mostly at one-loop order, making explicit the dependence on the elastic coefficient $K$, which usually is set to 1 . Details (mostly at $K=1$ ) can be found in $[36,49]$. The FRG equation for the function $\Delta(u)$ to one loop is

$$
\begin{equation*}
-m \partial_{m} \Delta(u)=-\frac{1}{2}\left(-m \partial_{m} I_{2}\right)\left[[\Delta(u)-\Delta(0)]^{2}\right]^{\prime \prime} \tag{B1}
\end{equation*}
$$

where $I_{n}=\int_{k}\left(K k^{2}+m^{2}\right)^{-n}$, with an implicit UV cutoff $\Lambda$. One has $-m \partial_{m} I_{2}=4 m^{2} I_{3}$. One should distinguish $d<6$ and $d>6$. The latter is dominated by the UV cutoff and is briefly discussed at the end. For $d<6$, in the limit $m \ll \Lambda$ one has $I_{3}=\widetilde{I}_{3} m^{d-6} K^{-d / 2}$ with $\tilde{I}_{n}=\int_{k}\left(k^{2}+1\right)^{-n}$ (defined with infinite

UV cutoff whenever convergent). Thus $\widetilde{I}_{3}=\left(\epsilon \widetilde{I}_{2}\right) / 4$ with $\widetilde{I}_{2}=(4 \pi)^{-d / 2} \Gamma\left(2-\frac{d}{2}\right)$. Hence for $d<6$,

$$
\begin{equation*}
-m \partial_{m} I_{2}=4 m^{2} I_{3}=m^{d-4}\left(\epsilon \tilde{I}_{2}\right) K^{-d / 2} \tag{B2}
\end{equation*}
$$

where the combination $\epsilon \widetilde{I}_{2}=2(4 \pi)^{-d / 2} \Gamma\left(3-\frac{d}{2}\right)$ is well defined for all $d<6$ (the pole at $d=4$ is suppressed). One defines the rescaled (dimensionless) function $\widetilde{\Delta}(u)$ through

$$
\begin{equation*}
\Delta(u)=\frac{K^{d / 2}}{\epsilon \widetilde{I}_{2}} m^{\epsilon-2 \zeta} \widetilde{\Delta}\left(u m^{\zeta}\right) \tag{B3}
\end{equation*}
$$

with $\epsilon \tilde{I}_{2}=1 /\left(8 \pi^{2}\right)$ in $d=4$. It satisfies the dimensionless FRG equation,

$$
\begin{equation*}
-m \partial_{m} \widetilde{\Delta}(u)=(\epsilon-2 \zeta) \widetilde{\Delta}(u)+\zeta u \widetilde{\Delta}^{\prime}(u)-\frac{1}{2}\left\{[\tilde{\Delta}(u)-\widetilde{\Delta}(0)]^{2}\right\}^{\prime \prime} \tag{B4}
\end{equation*}
$$

This equation is valid for all $d<6$ and yields a fixed point $\widetilde{\Delta}^{*}(u)=O(\epsilon)$ for $d<4$ with $\zeta=\zeta_{1} \epsilon$ and a few universality classes. The two loop equation, not reproduced here, was also analyzed. We recall the main results:
(a) Random-bond class,

$$
\begin{gather*}
\zeta=0.208298042 \epsilon+0.006858 \epsilon^{2}+O\left(\epsilon^{3}\right),  \tag{B5}\\
\frac{\left[-\widetilde{\Delta}^{* \prime}\left(0^{+}\right)\right]^{2}}{\widetilde{\Delta}^{*}(0)}=0.583405 \epsilon+0.294205 \epsilon^{2}+O\left(\epsilon^{3}\right) . \tag{B6}
\end{gather*}
$$

Since Eq. (B4) is invariant under $\widetilde{\Delta}(u) \rightarrow \xi^{-2} \widetilde{\Delta}(\xi u), \widetilde{\Delta}(u)$ contains one nonuniversal scale, and only the above ratio is universal.
(b) Random-field class, with $R(u)=-\sigma|u|$ at large $u$,

$$
\begin{gather*}
\zeta=\frac{\epsilon}{3} \\
\widetilde{\Delta}^{*}(0)=\frac{\epsilon}{3} \dot{\xi}^{2}, \quad-\widetilde{\Delta}^{* \prime}\left(0^{+}\right)=\frac{\epsilon}{3} \xi\left[1+\frac{2 \epsilon}{9}+O\left(\epsilon^{2}\right)\right] \tag{B7}
\end{gather*}
$$

where $\xi$ is the nonuniversal scale discussed above. Integrating $\Delta(u)$ from 0 to $\infty$ yields [49] [Eq. (4.39)]

$$
\begin{equation*}
K^{-d / 2}\left(\epsilon \tilde{I}_{2}\right) \sigma=\frac{\epsilon}{3} \xi^{3}\left[\gamma_{1}+\epsilon \gamma_{2}+O\left(\epsilon^{2}\right)\right] \tag{B8}
\end{equation*}
$$

with $\quad \gamma_{1}=\int_{0}^{1} d y \sqrt{2(y-1-\ln y)}=0.775304245188$, $\gamma_{2}=-0.13945524$ (see [49] [Eq. (4.41)])
(c) Periodic class with period $u=a$,

$$
\begin{gather*}
\zeta=0  \tag{B9}\\
\widetilde{\Delta}^{*}(0)=a^{2}\left[\frac{\epsilon}{36}+\frac{\epsilon^{2}}{54}+O\left(\epsilon^{3}\right)\right]  \tag{B10}\\
-\widetilde{\Delta}^{* \prime}\left(0^{+}\right)=a\left[\frac{\epsilon}{6}+\frac{\epsilon^{2}}{9}+O\left(\epsilon^{3}\right)\right] \tag{B11}
\end{gather*}
$$

For all classes,

$$
\begin{equation*}
\widetilde{\Delta}^{* \prime \prime}\left(0^{+}\right)=\epsilon\left(1-\zeta_{1}\right) / 3+O\left(\epsilon^{2}\right) \tag{B12}
\end{equation*}
$$

In dimension $d=4$ one defines

$$
\begin{equation*}
\widetilde{\Delta}(u)=\hat{\Delta}\left(u \ln \left(m_{0} / m\right)^{-\zeta_{1}}\right)\left[\ln \left(m_{0} / m\right)\right]^{-1+2 \zeta_{1}}, \tag{B13}
\end{equation*}
$$

and setting $\zeta=0$ and $\epsilon=0$ in Eq. (B4) one finds

$$
\begin{equation*}
\ell \partial_{\ell} \hat{\Delta}(u)=\left(1-2 \zeta_{1}\right) \hat{\Delta}(u)+\zeta_{1} u \hat{\Delta}^{\prime}(u)-\frac{1}{2}\left\{[\hat{\Delta}(u)-\hat{\Delta}(0)]^{2}\right\} \tag{B14}
\end{equation*}
$$

where $\ell=\ln \left(m_{0} / m\right)$. Hence $\hat{\Delta}(u)$ converges to the same fixed point $\hat{\Delta}^{*}(u)=\widetilde{\Delta}^{*}(u) / \epsilon$ as for $\epsilon>0$. This yields for the original function $\Delta(u)$ in $d=4$,

$$
\begin{equation*}
\Delta(u)=8 \pi^{2} K^{2} \hat{\Delta}^{*}\left(u \ln \left(m_{0} / m\right)^{-\zeta_{1}}\right)\left[\ln \left(m_{0} / m\right)\right]^{-1+2 \zeta_{1}} \tag{B15}
\end{equation*}
$$

up to subdominant terms of the order of $1 / \ln \left(m_{0} / m\right)$. Here $1 / m_{0}$ is a nonuniversal scale, presumably (at least) of the order of the Larkin scale. In dimension $4<d<6$, we set again $\zeta=0$. The equation for $\widetilde{\Delta}^{\prime \prime}(0)$ reads

$$
\begin{equation*}
-m \partial_{m} \widetilde{\Delta}^{\prime \prime}(0)=-(d-4) \widetilde{\Delta}^{\prime \prime}(0)-3 \widetilde{\Delta}^{\prime \prime}(0)^{2} \tag{B16}
\end{equation*}
$$

for an analytic disorder $\left[\Delta^{\prime \prime}(0)<0\right]$. Hence if the bare disorder is smooth one needs $\left|\widetilde{\Delta}_{0}^{\prime \prime}(0)\right|>(d-4) / 3$ to generate a cusp and metastability. Hence for sufficiently strong bare smooth disorder, or with (even weak) rough bare disorder, a cusp is generated. Eventually as $m \rightarrow 0$ the flows converge back to the attractive fixed point $\widetilde{\Delta}(u)=0$ as

$$
\begin{gather*}
-m \partial_{m} \widetilde{\Delta}(u)=-(d-4) \widetilde{\Delta}(u),  \tag{B17}\\
\widetilde{\Delta}(u)=\left(\frac{m}{m_{0}}\right)^{d-4} \widetilde{\Delta}^{*}(u),  \tag{B18}\\
\Delta(u)=\frac{K^{d / 2}}{\epsilon \widetilde{I}_{2} m_{0}^{d-4}} \widetilde{\Delta}^{*}(u), \tag{B19}
\end{gather*}
$$

where now $\widetilde{\Delta}^{*}(u)$ is nonuniversal and depends on details of the FRG flow at intermediate stages. It has a cusp, unless one starts from smooth weak disorder. Note that for $d \geqslant 4$, if we make the natural assumption (as for $d<4$ ) that no other strong-disorder fixed point exists, the asymptotic behavior is exactly given by the one-loop FRG equation (in $d=4$ ) and by perturbation theory in the renormalized disorder for $d>4$, i.e.,

$$
\begin{equation*}
\Delta(u)=\Delta_{0}(u)-\frac{1}{2} I_{2}\left\{\left[\Delta_{0}(u)-\Delta_{0}(0)\right]^{2}\right\}^{\prime \prime}+O\left(\Delta_{0}^{3}\right) \tag{B20}
\end{equation*}
$$

with $I_{2} \approx\left(\epsilon \tilde{I}_{2}\right)\left(\Lambda^{d-4}-m^{d-4}\right) /(d-4)$ (setting from now on $K$ $=1$ for simplicity). This equation is valid if we choose a model with weak and rough bare disorders, i.e., whose correlator exhibits a cusp. Of course, if we choose smooth bare disorder this perturbation formula fails, and one must run the RG to determine if the system is in the weak-smooth disorder phase where dimensional reduction holds (possibly up to rare events) or in the nonanalytic phase. Finally, for $d>6$ the situation is qualitatively similar up to additional dependence in the UV cutoff.

## APPENDIX C: GRAPHICAL INTERPRETATION OF $\mathcal{K} \Gamma[W]$ AT ONE-LOOP ORDER

In this appendix, we give an intuitive derivation of the self-consistent equation (151) at one-loop order. Two classes
of diagrams contribute, $\mathcal{C}_{1}$ and $\mathcal{C}_{2}$. We use the diagrammatics explained in Sec. V C which can be used both for statics and dynamics. Here we chose to show explicitly the arrows of causality. At this order, no difference is expected between static and dynamic shocks. We do not distinguish here between $Z$ and its rescaled version $\tilde{Z}$ and use loose notations.

Class $\mathcal{C}_{1}$. These are the diagrams in the calculation of $Z$ which, before expansion in $w$, are proportional to $\Delta(w)$ $-\Delta(0)$. Therefore in order to get the term of order $w, \Delta(w)$ $-\Delta(0)$ has to be expanded, and all other disorder vertices will be taken in the limit of $w \rightarrow 0$. For the $n$th cumulant, we obtain [with $\Delta(w)-\Delta(0)$ sitting at the bottom and $\Delta^{\prime \prime}$ at the top]

$$
\begin{align*}
& =n!\frac{1}{2} \sum_{l_{1} \geqslant 0, l_{2} \geqslant 0} \delta_{n-2=l_{1}+l_{2}}\left[-2 \Delta^{\prime}(w)\right]^{l_{1}}\left[-2 \Delta^{\prime}(w)\right]^{l_{2}} \\
& \times\left[\Delta^{\prime \prime}(w)+\Delta^{\prime \prime}(0)\right][\Delta(w)-\Delta(0)] \\
& =n!(n-1) 2^{n-3}\left[-\Delta^{\prime}(w)\right]^{n-2}\left[\Delta^{\prime \prime}(w)+\Delta^{\prime \prime}(0)\right][\Delta(w)-\Delta(0)] \\
& =-n!(n-1) 2^{n-2}\left|\Delta^{\prime}\left(0^{+}\right)\right|^{n-1} \Delta^{\prime \prime}(0) w \text {. } \tag{C1}
\end{align*}
$$

$l_{1}$ and $l_{2}$ denote the number of $\Delta^{\prime}$ in the left and right legs, respectively. Rescaling by $S_{m}$ amounts to set $\left|\Delta^{\prime}\left(0^{+}\right)\right| \rightarrow 1$. Hence we find that the coefficient $A_{n}$ defined in the text reads, for this class,

$$
\begin{equation*}
A_{n}^{\mathcal{C}_{1}}=n!\times 2^{n-2}(n-1) \tag{C2}
\end{equation*}
$$

Since in the self-consistent equation each outgoing line in the above is branched to a $Z$, this implies that the contribution of this class of diagrams to the right-hand side (rhs) of the selfconsistent equation can be summed into

$$
\begin{equation*}
\sum_{n=2}^{\infty} \frac{A_{n}^{\mathcal{C}_{1}}}{n!} I_{n} Z^{n}=Z^{2} \int_{k} \frac{1}{\left(k^{2}+m^{2}-2 Z\right)^{2}} \tag{C3}
\end{equation*}
$$

In the end one sets $m^{2} \rightarrow 1$ for rescaled quantities. Starting the sum instead at $n=3$, as we did in the main text, yields the additional subtraction $-Z^{2} \int_{k} \frac{1}{\left(k^{2}+m^{2}\right)^{2}}$. This procedure identifies the latter term as counterterm of the disorder renormalization, which is the correct expansion in terms of the renormalized disorder, whereas Eq. (C3) is the correct result in terms of the bare disorder.

We can now give a graphical interpretation to the dressed propagator $1 /\left(k^{2}+m^{2}-2 Z\right)$ which appears here. In the above formulas, the left and right legs are effectively dressed propagators, which we represent as a double line,

(C4)

The notation on the rhs of the equation is as follows: the left vertex of each disorder is at 0 and the right one at $w$ (the choice being dictated so as to have a nonvanishing vertex under application of the $\mathcal{K}$ operator). Arrows pointing to the right or left are going out-they will be branched to an external point sitting at 0 (left) or $w$ (right). From each disorder vertex $\Delta^{\prime}(w)$, there are two outgoing lines, resulting in a combinatorial factor of 2 for continuing either with the left (0) or right ( $w$ ) vertex. Also note that the object exists, independently of whether the incoming and outgoing lines are at 0 or $w$. Now class $\mathcal{C}_{1}$ can be written as

$$
\begin{align*}
\mathcal{C}_{1} & =0 \bigoplus_{w}^{w}+\cdots \cdots \bigoplus_{0} \\
& =\Delta^{\prime \prime}(0)[\Delta(w)-\Delta(0)] \int_{k} \frac{Z^{2}}{\left(k^{2}+m^{2}-2 Z\right)^{2}}+\mathcal{O}\left(w^{2}\right) . \tag{C5}
\end{align*}
$$

[Actually, the graphical notation is a little sloppy since the upper vertex can either be $\Delta^{\prime \prime}(w)$ or $\Delta^{\prime \prime}(0)$, but there is an additional combinatorial factor of $\frac{1}{2}$ in Eq. (C1), first line.] The result is in agreement with Eq. (C3).

Class $\mathcal{C}_{2}$. This class of diagrams looks like a correction to the critical force,

$$
\begin{align*}
\mathcal{C}_{2} & ={ }^{0} \cdots{ }^{w}+\cdots \\
& =\left[\Delta^{\prime}(w)-\Delta^{\prime}\left(0^{+}\right)\right] \int_{k} \frac{Z}{\left(k^{2}+m^{2}-2 Z\right)}-\frac{Z}{k^{2}+m^{2}} . \tag{C6}
\end{align*}
$$

Indeed it should be viewed as a circle of $\Delta^{\prime}$, of which exactly one is expanded in $w$, leading to a circle with one marked vertex. This is the vertex drawn above. Note that the double line needs at least one $\Delta^{\prime}(w)$, otherwise it cannot start at 0 and go to $w$ as indicated. This leads to the last term in Eq. (C6) being subtracted. If one wants the expression in terms of the renormalized disorder, one again has to subtract the contribution proportional to $\int_{k\left(k^{2}+m^{2}\right)^{2}}$, giving an additional term $-\left[\Delta^{\prime}(w)-\Delta^{\prime}\left(0^{+}\right)\right] \int_{k} \frac{2 Z}{\left(k^{2}+m^{2}\right)^{2}}$.

Equations (C5) and (C6) (with the proper subtraction) together give Eq. (151).

## APPENDIX D: MOMENTS FROM $\widetilde{Z}(\boldsymbol{\lambda})$

A direct series expansion of formula (166) in powers of $\lambda$, e.g., using MATHEMATICA, yields the following formula for the moments ( $S$ expressed in units of $S_{m}$ ):

$$
\begin{align*}
\frac{\left\langle S^{n}\right\rangle}{\langle S\rangle}= & \frac{(-4)^{n-1} \sqrt{\pi}}{\Gamma\left(\frac{3}{2}-n\right)}+\alpha\left[-\frac{3(-1)^{n} 4^{n-2} \sqrt{\pi} \psi\left(\frac{3}{2}-n\right)(n-1)}{\Gamma\left(\frac{3}{2}-n\right)}+\frac{(-1)^{n} 2^{2 n-3} \sqrt{\pi}(n-1)}{\Gamma\left(\frac{3}{2}-n\right)}-\frac{3(-1)^{n} 4^{n-2} \sqrt{\pi}[-2+\gamma+\log (4)](n-1)}{\Gamma\left(\frac{3}{2}-n\right)}\right. \\
& \left.-4^{n-1} \Gamma(n)-\frac{(-1)^{n} 2^{2 n-3} \sqrt{\pi} \psi\left(\frac{1}{2}-n\right)}{\Gamma\left(\frac{1}{2}-n\right)}+\frac{(-1)^{n} 2^{2 n-1} \sqrt{\pi}}{\Gamma\left(\frac{1}{2}-n\right)}-\frac{(-1)^{n} 2^{2 n-3} \sqrt{\pi}[-2+\gamma+\log (4)]}{\Gamma\left(\frac{1}{2}-n\right)}\right], \tag{D1}
\end{align*}
$$

which yields the universal ratios,

$$
\begin{equation*}
r_{n}=1+\frac{2}{2 n-3}+\frac{\alpha\left(n+\frac{(-1)^{n} \Gamma\left(\frac{5}{2}-n\right) \Gamma(n-1)}{\sqrt{\pi}}\right)}{(2 n-3)^{2}} \tag{D2}
\end{equation*}
$$

We note that these formula make sense only for $n$ integer, and in fact one can check that for integer $n=1,2, \ldots$ they give the same result as the formula in the text obtained from $P(S)$. Because of the factor $(-1)^{n}$ they are not suited for analytical continuation to any real $n$. However using reflection identities of $\Gamma(x)$ and $\psi(x)=\Gamma^{\prime}(x) / \Gamma(x)$ functions one can eliminate the factors $(-1)^{n}$ and get back the formula of the text which are real for all real $n$.

## APPENDIX E: CALCULATIONS FOR NONLOCAL ELASTICITY

Let us first recover the result for local elasticity $\gamma=2, d_{c}$ $=4$ with $\widetilde{g}_{k}^{-1}=k^{2}+1$. One can rewrite the momentum integral defined in Eq. (234) as

$$
\begin{align*}
\mathcal{J}_{2}= & N_{2} \int_{0}^{\infty} d s s^{d / 2-1}\left[\frac{\tilde{Z}^{2}}{(s+1-2 \widetilde{Z})^{2}}+\frac{\tilde{Z}}{s+1-2 \tilde{Z}}\right. \\
& \left.-\frac{\tilde{Z}}{s+1}-3 \frac{\widetilde{Z}^{2}}{(s+1)^{2}}\right] \tag{E1}
\end{align*}
$$

The constant $N_{2}$ is defined as

$$
N_{2}=\frac{1}{\epsilon} \frac{1}{\int_{0}^{\infty} d s s^{d / 2-1} \frac{1}{(s+1)^{2}}}
$$

Explicit calculation yields back, Eq. (149) and for $d \rightarrow 4$,

$$
\begin{equation*}
\mathcal{J}_{2}=\frac{1}{2} \tilde{Z}[2 \tilde{Z}+(1-3 \tilde{Z}) \log (1-2 \tilde{Z})]+O(\epsilon) \tag{E2}
\end{equation*}
$$

i.e., one recovers Eq. (150) by a different method, which is also a check of Eq. (234).

Let us now study the case of general $\gamma$ using model (228),

$$
\begin{align*}
\mathcal{J}_{\gamma}^{a}= & \frac{1}{\epsilon \tilde{I}_{2}} \int_{k}\left[\frac{\tilde{Z}^{2}}{\left(|k|^{\gamma}+1-2 \tilde{Z}\right)^{2}}+\frac{\tilde{Z}}{|k|^{\gamma}+1-2 \tilde{Z}}\right. \\
& \left.-\frac{\tilde{Z}}{|k|^{\gamma}+1}-3 \frac{\widetilde{Z}^{2}}{\left(|k|^{\gamma}+1\right)^{2}}\right] . \tag{E3}
\end{align*}
$$

Introducing $N_{1}^{a}=\frac{1}{2 \gamma-d}\left(\int_{0}^{\infty} d s s^{d / \gamma-1} \frac{1}{(s+1)^{2}}\right)^{-1}$, this is computed as

$$
\begin{align*}
\mathcal{J}_{\gamma}^{a}= & N_{1}^{a} \int_{0}^{\infty} d s s^{d / \gamma-1}\left[\frac{\widetilde{Z}^{2}}{(s+1-2 \tilde{Z})^{2}}+\frac{\tilde{Z}}{s+1-2 \tilde{Z}}\right. \\
& \left.-\frac{\tilde{Z}}{s+1}-3 \frac{\widetilde{Z}^{2}}{(s+1)^{2}}\right] \\
= & \frac{1}{\gamma} \tilde{Z}(2 \tilde{Z}+(1-3 \tilde{Z}) \log (1-2 \tilde{Z}))+O(2 \gamma-d) \tag{E4}
\end{align*}
$$

i.e., a function identical to Eq. (E2) at the critical dimension up to the global multiplication by $2 / \gamma$. Hence the distribution of avalanche sizes to one loop with this choice of nonlocal elasticity will be exactly given by Eq. (68) with the replacement of $\alpha \rightarrow 2 \alpha / \gamma$ in Eqs. (170) and (171) as detailed in the text.

Let us now study the case $\gamma=1$ and form (229) suitable to describe the contact line. We need to compute

$$
\begin{align*}
\mathcal{J}_{1}^{b}= & N_{1}^{b}\left[\int_{0}^{\infty} k^{d-1} d k \frac{\tilde{Z}^{2}}{\left(\sqrt{k^{2}+1}-2 \tilde{Z}\right)^{2}}+\frac{\tilde{Z}}{\sqrt{k^{2}+1}-2 \tilde{Z}}\right. \\
& \left.-\frac{\tilde{Z}}{\sqrt{k^{2}+1}}-3 \frac{\tilde{Z}^{2}}{k^{2}+1}\right] \tag{E5}
\end{align*}
$$

where $N_{1}^{b}=\left(\epsilon \int_{0}^{\infty} \frac{k^{d-1} d k}{k^{2}+1}\right)^{-1}, \epsilon=2-d$. We will use the following integral representation:

$$
\begin{equation*}
\frac{1}{\sqrt{1+k^{2}}-2 \tilde{Z}}=\int_{\alpha=0}^{\infty} \int_{t=0}^{\infty} \frac{e^{-\left(1+k^{2}\right) \alpha t^{2}+2 \tilde{Z} t-1 / 4 \alpha}}{2 \sqrt{\pi} \alpha^{3 / 2}} \tag{E6}
\end{equation*}
$$

This yields

$$
\begin{equation*}
N_{1}^{b} \int_{k}\left[\frac{1}{\sqrt{1+k^{2}}-2 \tilde{Z}}-\frac{1}{\sqrt{1+k^{2}}}\right]=\frac{2 \widetilde{Z}\left(1-4 \widetilde{Z}^{2}\right)^{d / 2-1}}{2-d}-\frac{\Gamma\left(\frac{1}{2}-\frac{d}{2}\right) \Gamma\left(\frac{d}{2}-1\right)\left[{ }_{2} F_{1}\left(1, \frac{1-d}{2} ; \frac{1}{2} ; 4 \widetilde{Z}^{2}\right)-1\right] \sin \left(\frac{d \pi}{2}\right)}{2 \pi^{3 / 2}} \tag{E7}
\end{equation*}
$$

Taking a derivative with respect to $\tilde{Z}$ yields

$$
\begin{equation*}
\int_{k} \frac{1}{\left(\sqrt{1+k^{2}}-2 \widetilde{Z}\right)^{2}}=\frac{\left(1-4 \widetilde{Z}^{2}\right)^{(d-4) / 2}\left[4(d-1) \widetilde{Z}^{2}-1\right]}{d-2}-\frac{4 \widetilde{Z} \Gamma\left(\frac{3}{2}-\frac{d}{2}\right) \Gamma\left(\frac{d}{2}-1\right){ }_{2} F_{1}\left(2, \frac{3}{2}-\frac{d}{2} ; \frac{3}{2} ; 4 \widetilde{Z}^{2}\right) \sin \left(\frac{d \pi}{2}\right)}{\pi^{3 / 2}} \tag{E8}
\end{equation*}
$$

Together this gives

$$
\begin{align*}
\mathcal{J}_{1}^{b} & =\frac{\widetilde{Z}^{2}\left(4(d+1) \widetilde{Z}^{2}-3\right)\left(1-4 \widetilde{Z}^{2}\right)^{(d-4) / 2}}{d-2}-\frac{3 \widetilde{Z}^{2}}{d-2}+\frac{\tilde{Z} \Gamma\left(\frac{1}{2}-\frac{d}{2}\right)\left[{ }_{2} F_{1}\left(2, \frac{1}{2}-\frac{d}{2} ; \frac{1}{2} ; 4 \widetilde{Z}^{2}\right)-1\right]}{2 \sqrt{\pi} \Gamma\left(2-\frac{d}{2}\right)}  \tag{E9}\\
& =\frac{2 \widetilde{Z}^{3}}{1-2 \tilde{Z}}-3 \widetilde{Z}^{2} \log (1-2 \tilde{Z})+O(2-d) .
\end{align*}
$$

This result (E10) has to be compared to Eq. (E4) for $\gamma=1$. One sees that the leading behavior for $\widetilde{Z} \rightarrow-\infty$, from which is extracted the exponent $\tau$, is identical hence the avalanche size exponent $\tau$ is still given by Eq. (237). The form of the avalanche distribution however will be different. To see this, first invert

$$
\begin{equation*}
\tilde{Z}=\lambda+\widetilde{Z}^{2}+\alpha\left[\frac{2 \widetilde{Z}^{3}}{1-2 \tilde{Z}}-3 \widetilde{Z}^{2} \log (1-2 \tilde{Z})\right] \tag{E11}
\end{equation*}
$$

This gives

$$
\begin{equation*}
\widetilde{Z}_{\gamma=1}^{b}(\lambda)=\frac{1}{2}(1-\sqrt{1-4 \lambda})-\alpha \frac{(\sqrt{1-4 \lambda}-1)^{2}\left[\frac{3}{2} \sqrt{1-4 \lambda} \log (1-4 \lambda)+\sqrt{1-4 \lambda}-1\right]}{4(1-4 \lambda)} . \tag{E12}
\end{equation*}
$$

One can now verify that to first order in $\alpha$ one can write (setting $S_{m} \rightarrow 1$, restored in the text)

$$
\begin{equation*}
\widetilde{Z}_{\gamma=1}^{b}(\lambda)=\int_{0}^{\infty} d S\left[e^{\lambda S}-1\right] P_{\gamma=1}^{b}(S) \tag{E13}
\end{equation*}
$$

with

$$
\begin{align*}
P_{\gamma=1}^{b}(S)= & \frac{1}{2 \sqrt{\pi}}\left[1+\frac{\alpha}{4}\left(8-3 \gamma_{E}\right)\right]\left[S^{-3 / 2-3 \alpha / 4}+\frac{\sqrt{\pi} \alpha}{8}\right] \\
& \times \exp \left[-\frac{1}{4}\left(1-\frac{3}{2}\left(\gamma_{E}-2\right) \alpha\right) S^{1-3 \alpha / 2}-\frac{3}{2} \sqrt{\pi} \alpha \sqrt{S}\right], \tag{E14}
\end{align*}
$$

which is the form given in the text. Note that to first order in $\alpha$ it can equally well be written as

$$
\begin{align*}
P_{\gamma=1}^{b}(S)= & \frac{1}{2 \sqrt{\pi}}\left[1+\frac{\alpha}{4}\left(8-3 \gamma_{E}\right)\right] S^{-3 / 2-3 \alpha / 4} \\
& \times \exp \left[-\frac{1}{4}\left(1-\frac{3}{2}\left(\gamma_{E}-2\right) \alpha\right) S^{1-3 \alpha / 2}\right. \\
& \left.-3 / 2 \sqrt{\pi} \alpha \sqrt{S}+\frac{\sqrt{\pi} \alpha}{8} S^{3 / 2}\right] \tag{E15}
\end{align*}
$$

This would suggest an abrupt jump in the effective stretched exponential decay exponent $\delta$ to the value $\frac{3}{2}$. However since the above results are valid for fixed $S$ at first order in $\alpha$ one cannot conclude on this point at this stage.

## APPENDIX F: MORE ON SPATIAL STRUCTURE OF AVALANCHE DISTRIBUTIONS

## 1. Derivation of the recursion relation

We want to show in Fourier space, using $\phi(x)=\cos (q x)$,

$$
\begin{equation*}
Z_{k}(\lambda)=\lambda \frac{1}{2}\left[\delta_{k+q}+\delta_{k-q}\right]+\left|\Delta^{\prime}\left(0^{+}\right)\right| \int_{p} g_{p} g_{k-p} Z_{p}(\lambda) Z_{k-p}(\lambda), \tag{F1}
\end{equation*}
$$

where $\delta_{q}=(2 \pi)^{d} \delta^{d}(q)$.
The calculation of $G^{\phi}(\lambda)$ is a generalization of the one of Sec. V to nonzero momentum graphs. One still takes $w$ uniform and now computes the spatially dependent correlations (for $q_{i}$ all nonvanishing),

$$
\begin{align*}
& g_{q}^{-2}{\overline{u_{q}\left(w_{1}\right) u_{q^{\prime}}\left(w_{2}\right)}}^{c}=-R_{q}^{\prime \prime}\left[w_{1}-w_{2}\right] \delta_{q+q^{\prime}},  \tag{F2}\\
& {\overline{u_{q_{1}}}\left(w_{1}\right) \cdots u_{q_{n}}\left(w_{n}\right)}^{c} \\
& \quad=(-1)^{n}{\hat{\hat{V}_{q_{1}}\left[w_{1}\right] \cdots \hat{V}_{q_{n}}^{\prime}\left[w_{n}\right]}}^{c} \\
& \quad=(-1)^{n} \hat{C}_{q_{1}, \ldots, q_{n}}^{(n)}\left(w_{1}, \ldots, w_{n}\right) \delta_{q_{1}+\cdots q_{n}}, \tag{F3}
\end{align*}
$$

where $\delta_{q}=(2 \pi)^{d} \delta^{d}(q)$. We use Fourier space versions of the definitions in Eq. (4). In particular we have defined

$$
\begin{equation*}
R_{q}^{\prime \prime}[w]:=\int_{x} e^{i q x} R_{0 x}^{\prime \prime}[w], \quad R_{x y}^{\prime \prime}[w]:=\left.\frac{\delta^{2} R[w]}{\delta w_{x} \delta w_{y}}\right|_{w_{x}=w} \tag{F4}
\end{equation*}
$$

using translational invariance for a uniform $w$ and the definition of the $R[w]$ functional [62] ${\hat{\hat{V}}\left[w_{1}\right] \hat{V}\left[w_{2}\right]}^{c}=R\left[w_{1}-w_{2}\right]$. The $q=0$ limit is recovered using $u_{q=0}(w)=L^{d} u(w)$. From these $\hat{C}$ one computes the associated momentum-dependent Kolmogorov moments, applying the operator $\mathcal{K}$. The choice $\phi(x)=\cos (q x)$ amounts to choose $q_{i}= \pm q$ on the external legs of the tree diagrams of Sec. V, as for the calculation of $\left|S_{q}\right|^{2 n}$; hence we need $q_{i}= \pm q$.

We extend the previous tree-level analysis and associate $a_{m, m^{\prime}}$ to the box with $m$ times $+q$ and $m^{\prime}$ times $-q$ entering from the top and a moment $\left(m-m^{\prime}\right) q$ exiting from the bottom of the tree diagram. Pasting together the trees as explained in Sec. V we obtain the recursion relation,

$$
\begin{equation*}
a_{m, m^{\prime}}=\sum_{p+l=m ; p^{\prime}+l^{\prime}=m^{\prime}} a_{p, p^{\prime}} a_{l, l^{\prime}} \frac{m!}{p!l!} \frac{m^{\prime}!}{p^{\prime}!l^{\prime}!} g_{\left(p-p^{\prime}\right) q} g_{\left(l-l^{\prime}\right) q} . \tag{F5}
\end{equation*}
$$

Introducing $b_{m, m^{\prime}}:=a_{m, m^{\prime}} / m!m^{\prime}$ ! one obtains

$$
\begin{align*}
b_{m, m^{\prime}} e^{i z\left(m-m^{\prime}\right) q}= & \int_{x y} g_{x} g_{y} \sum_{p+l=m ; p^{\prime}+l^{\prime}=m^{\prime}} b_{p, p^{\prime}} \\
& \times b_{l, l^{\prime}} e^{i(x+z)\left(p-p^{\prime}\right) q} e^{i(y+z)\left(l-l^{\prime}\right) q} . \tag{F6}
\end{align*}
$$

By definition,

$$
\begin{equation*}
Z_{z}(\lambda):=\sum_{m, m^{\prime}} b_{m, m^{\prime}}\left|\Delta^{\prime}\left(0^{+}\right)\right|^{m+m^{\prime}+1} e^{i z\left(m-m^{\prime}\right) q} \lambda^{m+m^{\prime}} \tag{F7}
\end{equation*}
$$

In sum (F5) the terms 1.0 and 0,1 are left out, so one finds in real space

$$
\begin{equation*}
Z_{z}(\lambda)=\lambda \cos (q z)+\left|\Delta^{\prime}\left(0^{+}\right)\right| \int_{x y} g_{z-x} g_{z-y} Z_{x}(\lambda) Z_{y}(\lambda) \tag{F8}
\end{equation*}
$$

The Fourier transform of this relation gives Eq. (F1) and it is a particular case of Eq. (202). The general case is obtained by attaching $\phi(x)$ to each external leg. At the end of the calculation one wants $Z_{k=0}(\lambda)$ since no momentum flows from the lower vertex, hence formula (201). Note that we have neglected everywhere the nonlocal component (i.e., the momentum dependence) of the $\Delta^{\prime}(0)$ vertex which we have taken at zero momentum. Near $d=4$ this should be sufficient, as discussed again below.

## 2. Generalized cusp-moment relation and check on the second order

Let us start from the general relation

$$
\begin{equation*}
\overline{\left[u_{x}\left(w_{1}\right)-w_{1}\right]\left[u_{y}\left(w_{2}\right)-w_{2}\right]^{c}}=-\int_{z z^{\prime}} g_{x z} g_{y z} R_{x y}^{\prime \prime}\left[w_{1}-w_{2}\right] . \tag{F9}
\end{equation*}
$$

Consider now the expansion to second order $\lambda^{2}$ of Eq. (198). It yields

$$
\begin{equation*}
\left.G^{\phi}(\lambda)\right|_{\lambda^{2}}=\frac{1}{\int_{x} \phi(x)} \int_{x y z z^{\prime}} \phi(x) \phi(y) g_{x z} g_{y z^{\prime}}\left(R_{z z^{\prime}}^{\prime \prime}[w]-R_{z z^{\prime}}^{\prime \prime},[0]\right) . \tag{F10}
\end{equation*}
$$

Taking a derivative with respect to $w$ we obtain

$$
\begin{equation*}
\left.\hat{Z}^{\phi}(\lambda)\right|_{\lambda^{2}}=\frac{1}{\int_{x} \phi(x)} \int_{x y z z^{\prime} t} \phi(x) \phi(y) g_{x z} g_{y z^{\prime}} R_{z z^{\prime} t^{\prime}}^{\prime \prime \prime}\left[w=0^{+}\right] . \tag{F11}
\end{equation*}
$$

Comparing with its definition, and using the relation $\left\langle S^{\phi}\right\rangle$ $=\langle S\rangle L^{-d} \int d x \phi(x)$, yields the result for the second moment of the avalanche amplitude at different points in internal space,

$$
\begin{equation*}
\frac{\left\langle S^{x} S^{y}\right\rangle}{\langle S\rangle}=2 L^{-d} \int_{z z^{\prime} t} g_{x z} g_{y z^{\prime}} R_{z z^{\prime} t^{\prime}}^{\prime \prime \prime}\left[w=0^{+}\right], \tag{F12}
\end{equation*}
$$

an exact relation which is the nonlocal generalization of Eq. (36).

To leading order in $\epsilon$ the third derivative vertex is local, i.e., $R_{z z^{\prime \prime} t}^{\prime \prime \prime}[w]=\delta_{z z^{\prime} t} R^{\prime \prime \prime}(w)$ : hence we find simply

$$
\begin{equation*}
\frac{\left\langle S^{x} S^{y}\right\rangle}{\langle S\rangle}=2 L^{-d} \int_{z} g_{x z} g_{y z} R^{\prime \prime \prime}\left(0^{+}\right) \tag{F13}
\end{equation*}
$$

or in $q$ space

$$
\begin{equation*}
\frac{\left\langle S^{q} S^{q^{\prime}}\right\rangle}{\langle S\rangle}=2\left|\Delta^{\prime}\left(0^{+}\right)\right| L^{-d} g_{q}^{2} \delta_{q+q^{\prime}} \tag{F14}
\end{equation*}
$$

This is what is predicted by the self-consistent equation,

$$
\begin{equation*}
\left\langle\left(S^{\phi}\right)^{2}\right\rangle /\langle S\rangle=2\left|\Delta^{\prime}\left(0^{+}\right)\right| L^{-d} \int_{x}\left(\int_{y} g_{x y} \phi(y)\right)^{2}, \tag{F15}
\end{equation*}
$$

a simple generalization of Eq. (36), valid to lowest order in $\epsilon$ (or in the locality expansion).

## APPENDIX G: DISTRIBUTION OF THE CENTER-OF-MASS FLUCTUATIONS

One defines the generating function for the probability distribution of the center-of-mass fluctuations as

$$
\begin{equation*}
f(\lambda)=L^{-d} \ln \overline{\exp \left(\lambda L^{d} u(0)\right)}=L^{-d} \overline{\ln \exp \left(\lambda \int_{x} u_{x}(0)\right)} \tag{G1}
\end{equation*}
$$

Result (135) implies that, to one-loop accuracy,

$$
\begin{align*}
f(\lambda)= & \frac{\lambda^{2}}{2} \frac{\Delta(0)}{m^{4}}-\int_{k} \sum_{p=2}^{\infty} \frac{1}{2 p} \frac{\Delta^{\prime}\left(0^{+}\right)^{2 p} \lambda^{2 p}}{m^{4 p}\left(k^{2}+m^{2}\right)^{2 p}} \\
= & \frac{\lambda^{2}}{2} \frac{\Delta(0)}{m^{4}}+\frac{1}{2} \int_{k}\left[\ln \left(1-\frac{\Delta^{\prime}\left(0^{+}\right)^{2} \lambda^{2}}{m^{4}\left(k^{2}+m^{2}\right)^{2}}\right)\right. \\
& \left.+\frac{\Delta^{\prime}\left(0^{+}\right)^{2} \lambda^{2}}{m^{4}\left(k^{2}+m^{2}\right)^{2}}\right] \tag{G2}
\end{align*}
$$

Note that the term $p=1$ is contained in the renormalized $\Delta(0)$. Since we work to one-loop order, we can now compute the integral in the second term in $d=4$. The calculation yields

$$
\begin{gather*}
f(\lambda)=\frac{\lambda^{2}}{2} \frac{\Delta(0)}{m^{4}}+S_{4} m^{d} g\left(\lambda m^{-4}\left|\Delta^{\prime}\left(0^{+}\right)\right|\right) \\
g(b)=\frac{1}{8}\left[(1+b)^{2} \ln (1+b)+(1-b)^{2} \ln (1-b)-3 b^{2}\right] \\
=-\frac{b^{4}}{48}-\frac{b^{6}}{240}-\frac{b^{8}}{672}-\frac{b^{10}}{1440}+O\left(b^{12}\right) \tag{G3}
\end{gather*}
$$

with $S_{4}=\epsilon \tilde{I}_{2}=1 /\left(8 \pi^{2}\right)$ in $d=4$. We note that the scale
$S_{m}=m^{-4}\left|\Delta^{\prime}\left(0^{+}\right)\right|=O\left(\epsilon m^{-d-\zeta}\right)$ naturally appears. Indeed we can rewrite the result for the generating function near $d=4$ using the rescaled correlator from Appendix B as

$$
\begin{align*}
& \exp \left(\lambda \int_{x} u_{x}(0) / S_{m}\right) \\
& =\exp \left([m L]^{d} \in \widetilde{I}_{2}\left[\frac{\lambda^{2}}{2} \frac{\widetilde{\Delta}(0)}{\widetilde{\Delta}^{\prime}\left(0^{+}\right)^{2}}+g(\lambda)\right]\right) \\
& =\exp \left(\frac{(m L)^{d}}{8 \pi^{2}}\left[\frac{\lambda^{2}}{2}\left(\frac{1}{\epsilon\left(1-2 \zeta_{1}\right)}+g_{2}\right)+g(\lambda)\right]\right) \tag{G4}
\end{align*}
$$

$\lambda$ is now a dimensionless number. The total factor in the exponential is obtained correctly to both orders $O(1 / \epsilon)$ and $O(1)$. In the last equation we have used the FRG fixed-point equation for $\widetilde{\Delta}(0)$. The term $g_{2}$ is zero if one uses the FRG equation to one loop. However to get the $\lambda^{2}$ term to $O(1)$ requires injecting the ratio $\frac{\tilde{\Delta}(0)}{\tilde{\Delta}^{\prime}\left(0^{+}\right)^{2}}$ to two loops. These ratio are given in the appendix and it is easy from there to find the number $g_{2}$ for each universality class.

Result (G4) seems to indicate [95] that the tail of $P(u)$ decays as $\exp \left(-|u|^{\delta}\right)$, with $\delta=2-\epsilon\left(1-2 \zeta_{1}\right) / 2<2$, contrary to the $d=0$ result $\delta=3$. It remains to be understood whether $\delta$ first decreases and then increases again for $d \rightarrow 0$ or whether the functional form is different, and the assumption that the $\ln$ can be re-exponentiated is unfounded. A larger tail ( $\delta<2$ ) seems to be counterintuitive since all connected moments in Eq. (G3) above the second are negative.

The main result (G4) can be interpreted as a nice example of the central-limit theorem. There are effectively $m L^{d}$ independent regions [96]. In each region the generating function is given by the same formula (G4) setting the $m L^{d}$ factor to unity. The final distribution of $u(0)$ is obtained as a convolution of these $m L^{d}$ identical ones. In each region one can write $u(0)=\sqrt{\epsilon} u_{1}+\epsilon u_{2}=S_{m}\left(\frac{1}{\sqrt{\epsilon}} \widetilde{u}_{1}+\widetilde{u}_{2}\right)$, where $u_{1}, u_{2}\left(\tilde{u}_{1}, \widetilde{u}_{2}\right)$ are independent random variables of order one, $u_{1}\left(\tilde{u}_{1}\right)$ has a Gaussian distribution, and $u_{2}\left(\tilde{u}_{2}\right)$ has a nontrivial distribution encoded in the function $g(\lambda)$. Hence near $d=4$ the center-of-mass fluctuations are Gaussian with a small correction encoded in $g(\lambda)$.
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to -1 below the cut and from -1 to $-u / 2$ above the cut has to be added. Its nonvanishing part comes from the $2 \pi i$ discontinuity of the $\ln$ across the cut, thus does not give a $\ln$ contribution. The final result at leading order is then obtained by replacing $\lambda$ by $\lambda-u / 2$, doing the (almost) Gaussian integral in $\lambda$ in imaginary direction, with the result that $P(u)$ $\left.\approx(G 4)\right|_{\lambda \rightarrow-u / 2}$. Exponentiating the $\ln$ leads to the quoted result for $\delta$.
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